

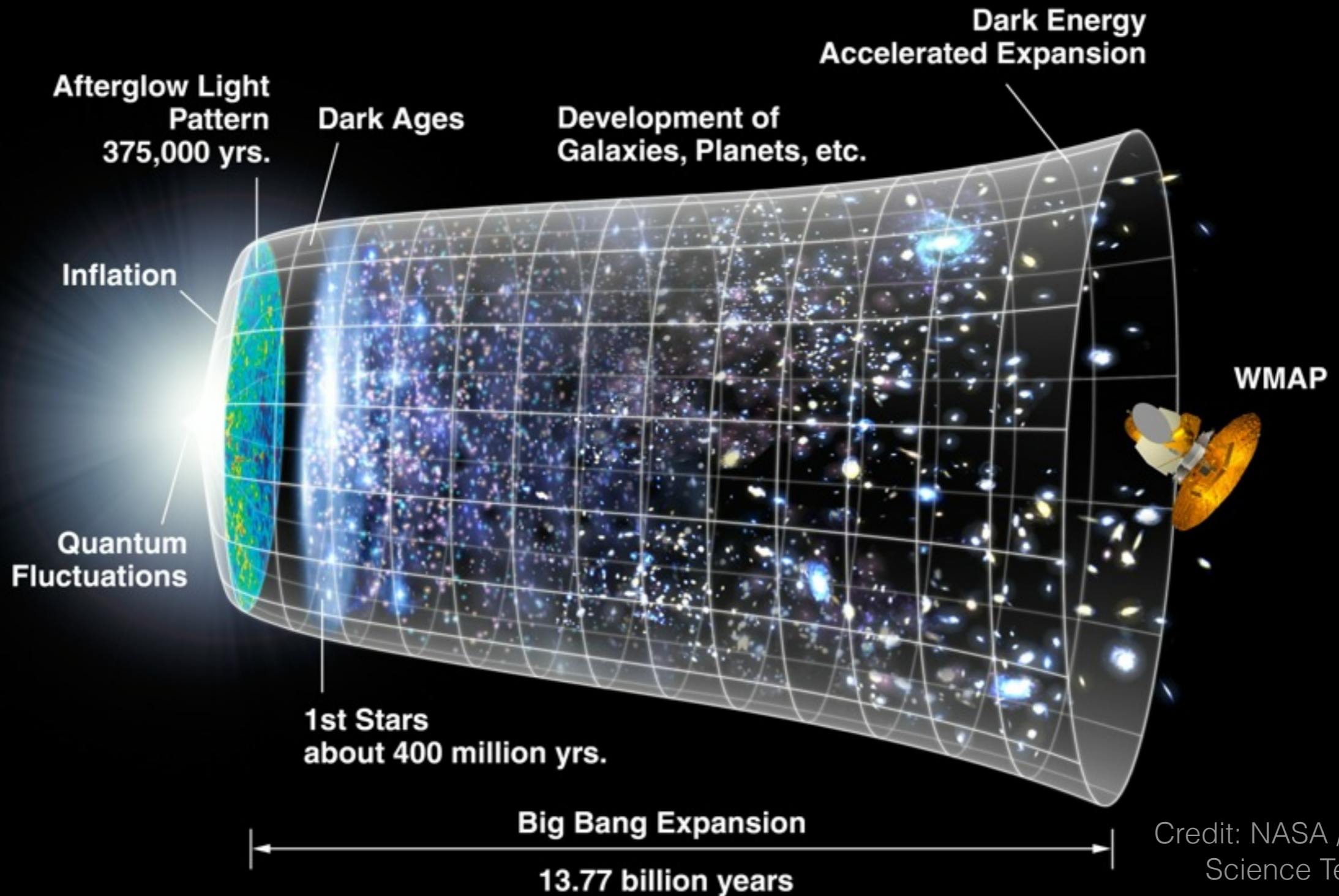
Computational Challenges in Cosmological Parameter Estimation

Sebastian Seehars and Joël Akeret

Joint work with Adam Amara, Alexandre Refregier, and André Csillaghy.
(In collaboration with University of Applied Sciences Northwestern Switzerland)

21.02.2014

Cosmology

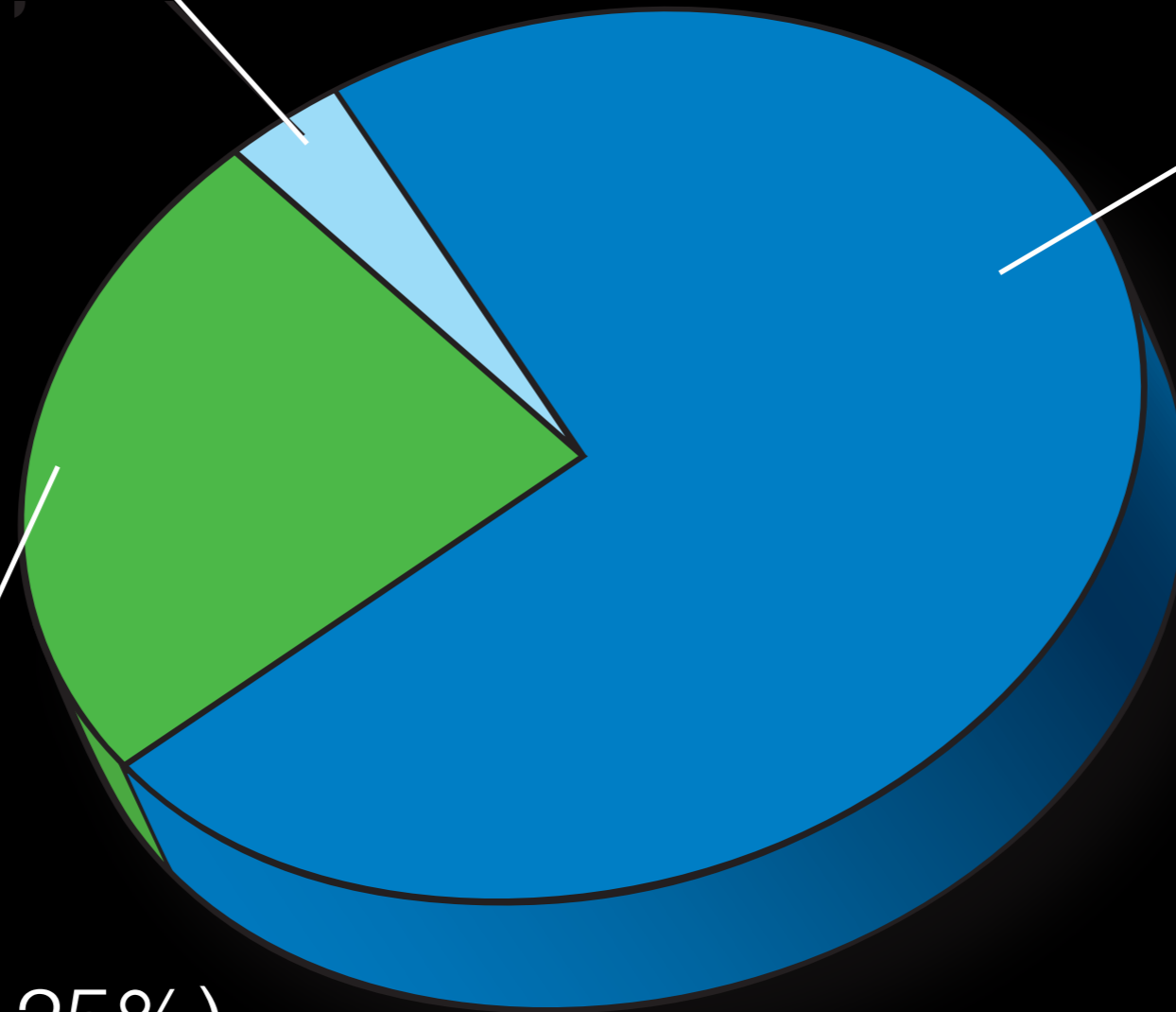


Credit: NASA / WMAP Science Team

Dark Components

Ordinary Matter (~5%)

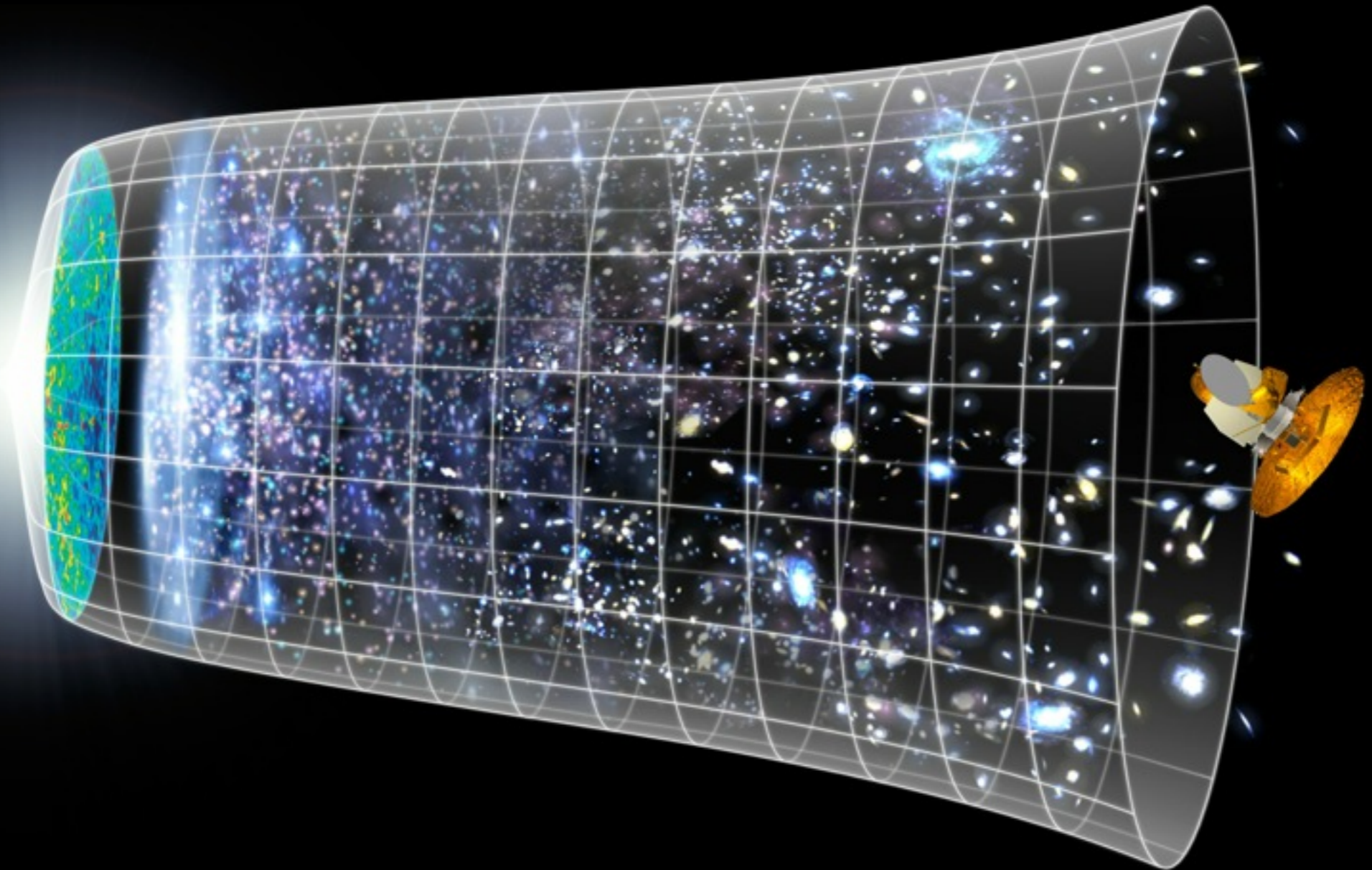
Dark Energy (~70%)



Dark Matter (~25%)

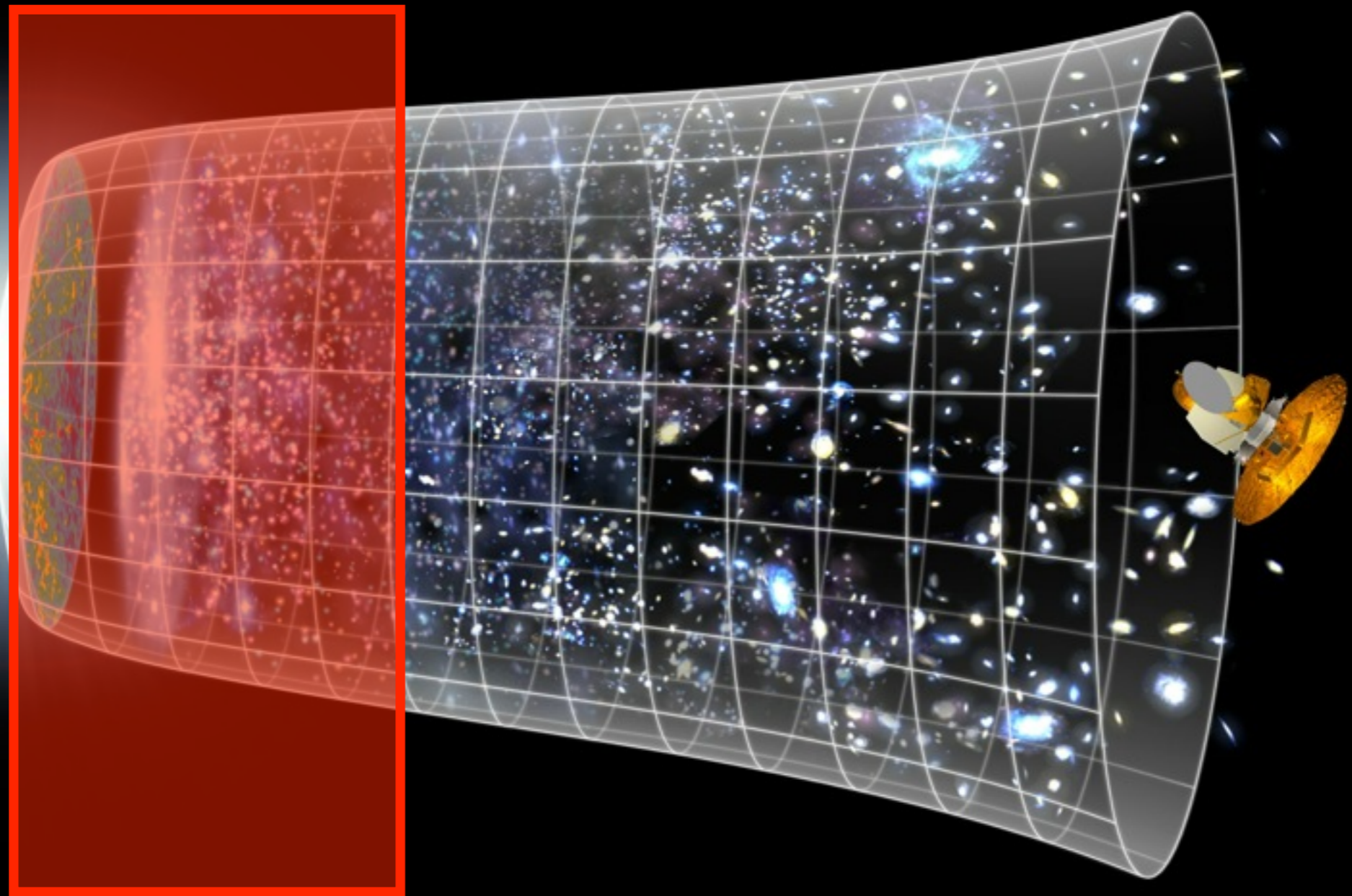
Credit: NASA / WMAP
Science Team

Computational Cosmology



Credit: NASA / WMAP
Science Team

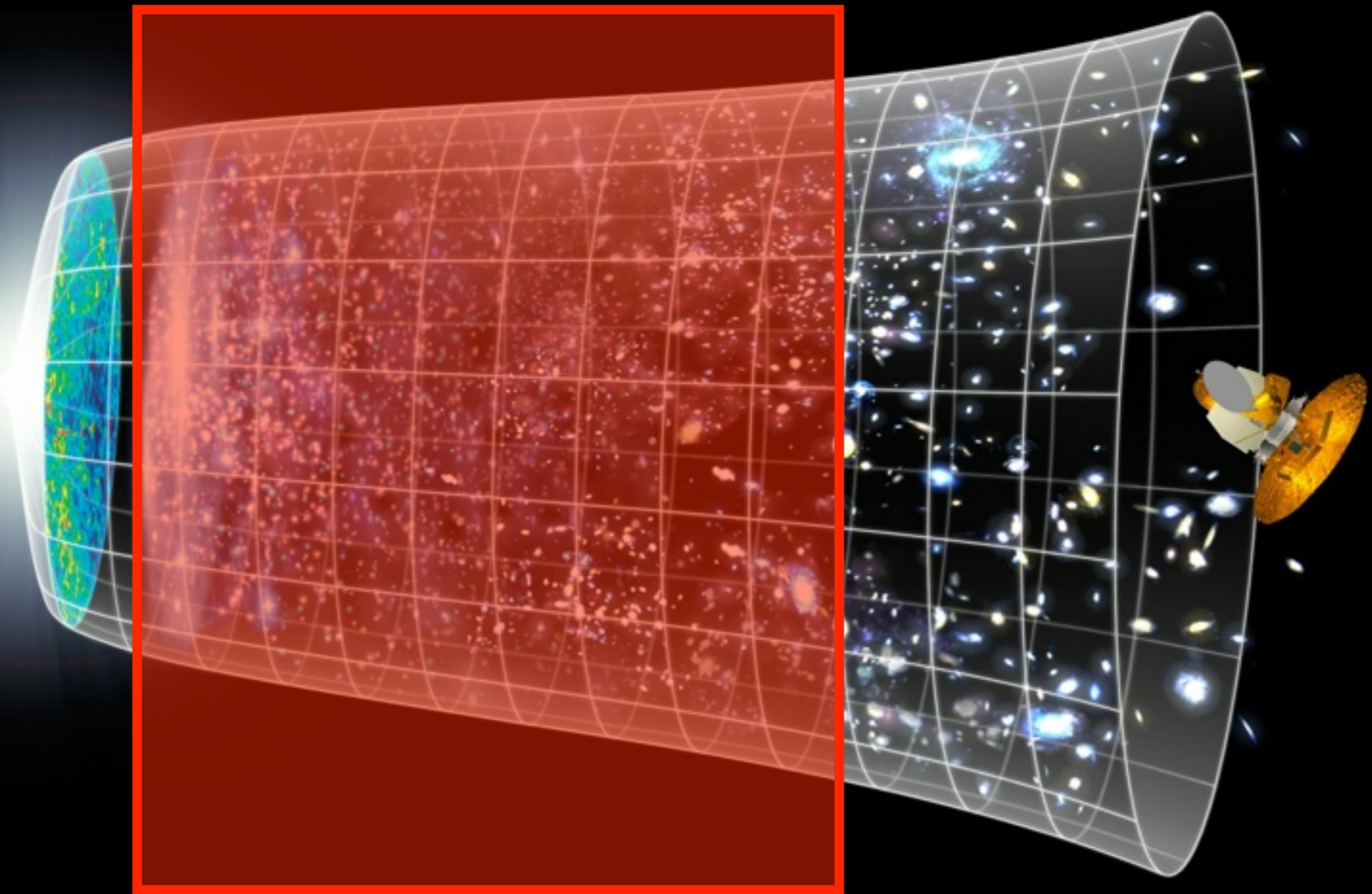
Computational Cosmology



Credit: NASA / WMAP
Science Team

Linear perturbation theory:
Solving Boltzmann equations numerically

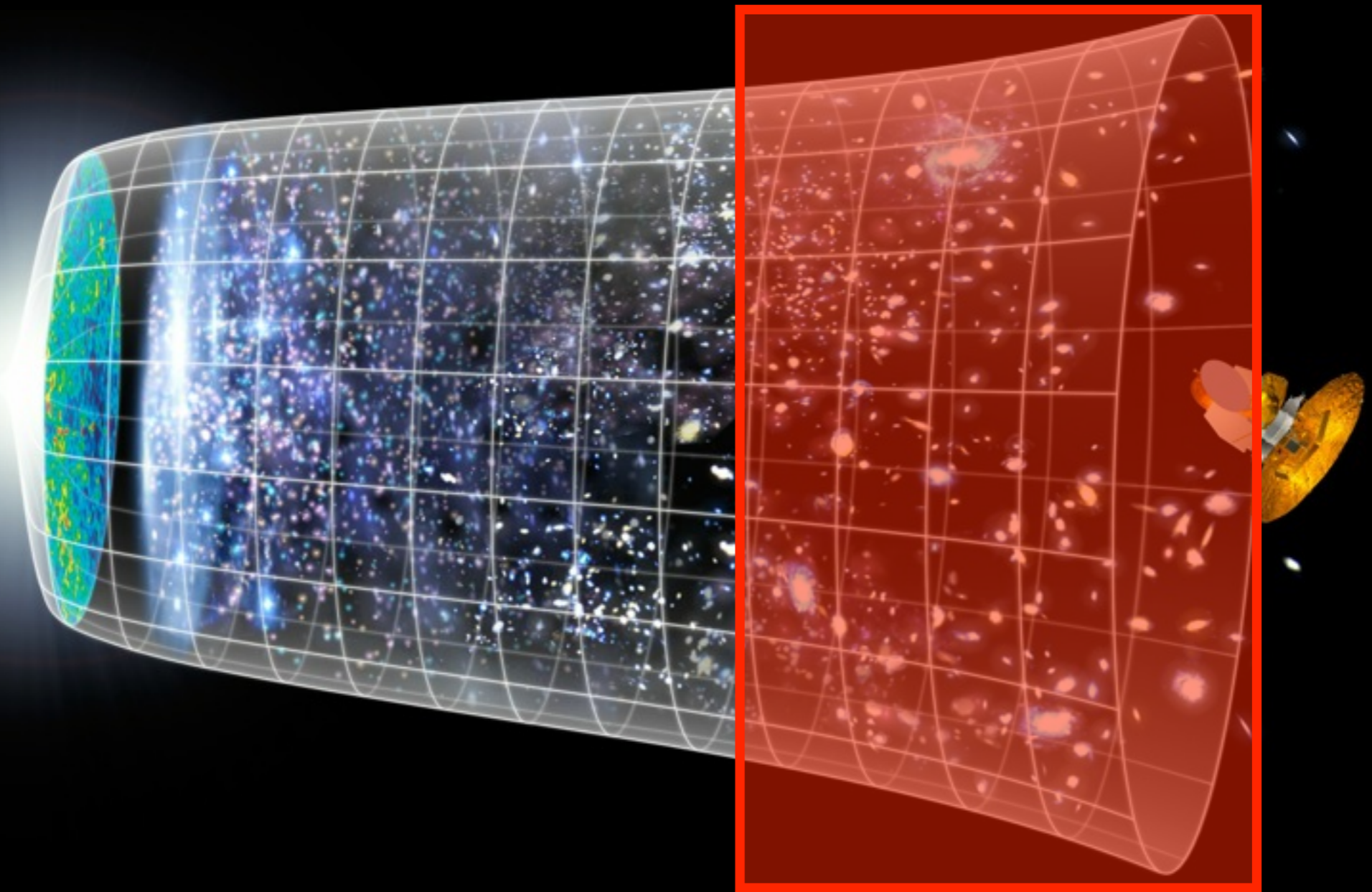
Computational Cosmology



Credit: NASA / WMAP
Science Team

Large scale structure formation:
N-body simulations of dark matter dynamics

Computational Cosmology



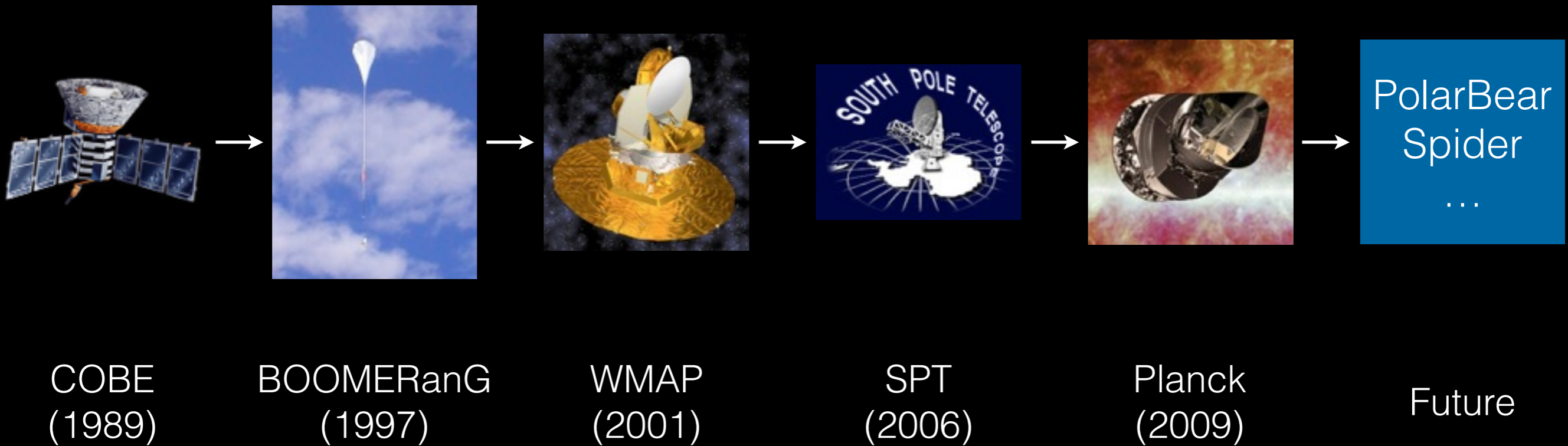
Credit: NASA / WMAP
Science Team

Galaxy formation:
Hydrodynamic simulations

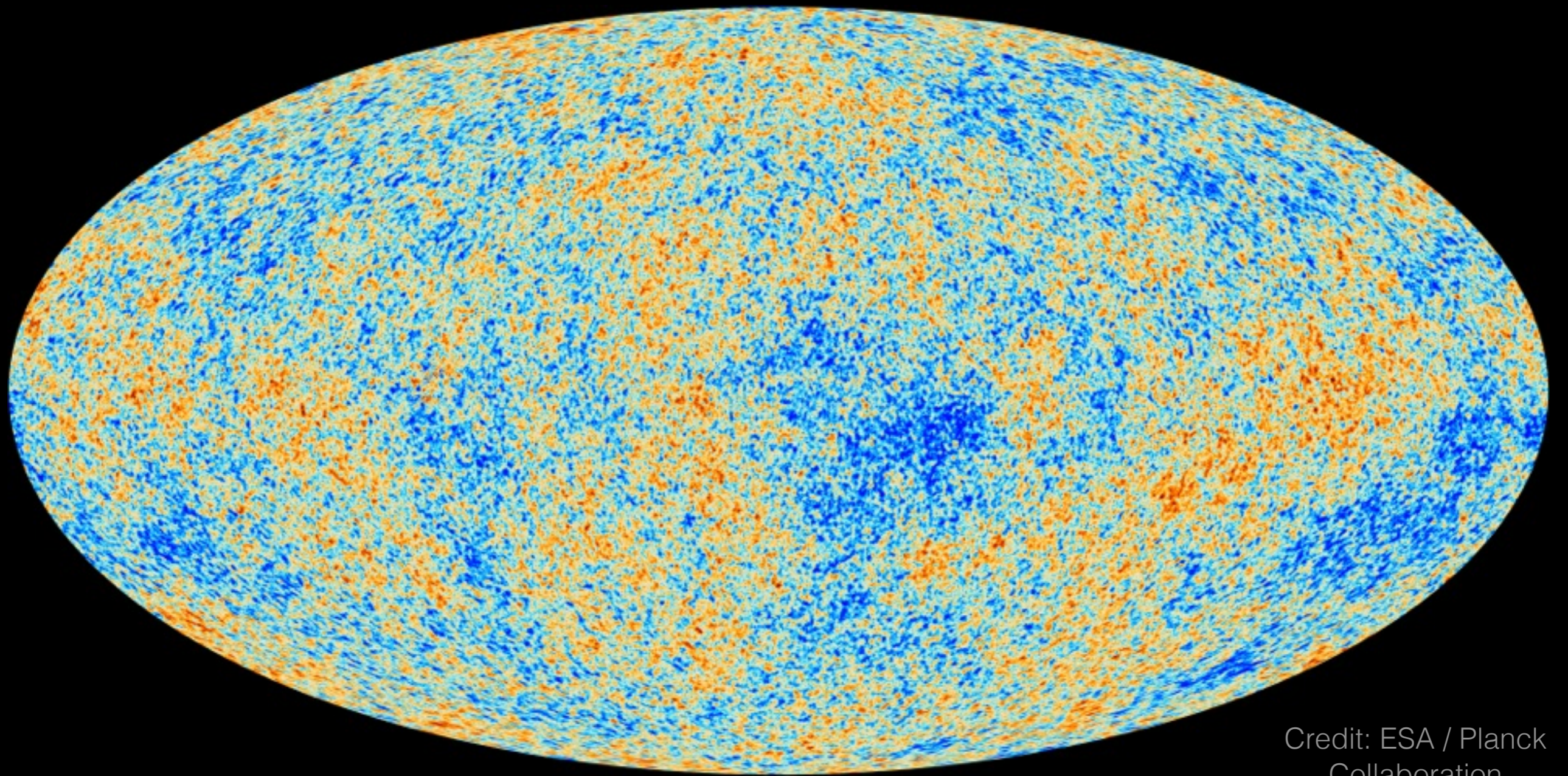
Data Analysis

- Processing and analysis of large datasets
- Forward modelling
 - Astronomical foregrounds
 - Instrument effects
- Statistical inference

Incomplete History of Cosmic Microwave Background (CMB) Surveys

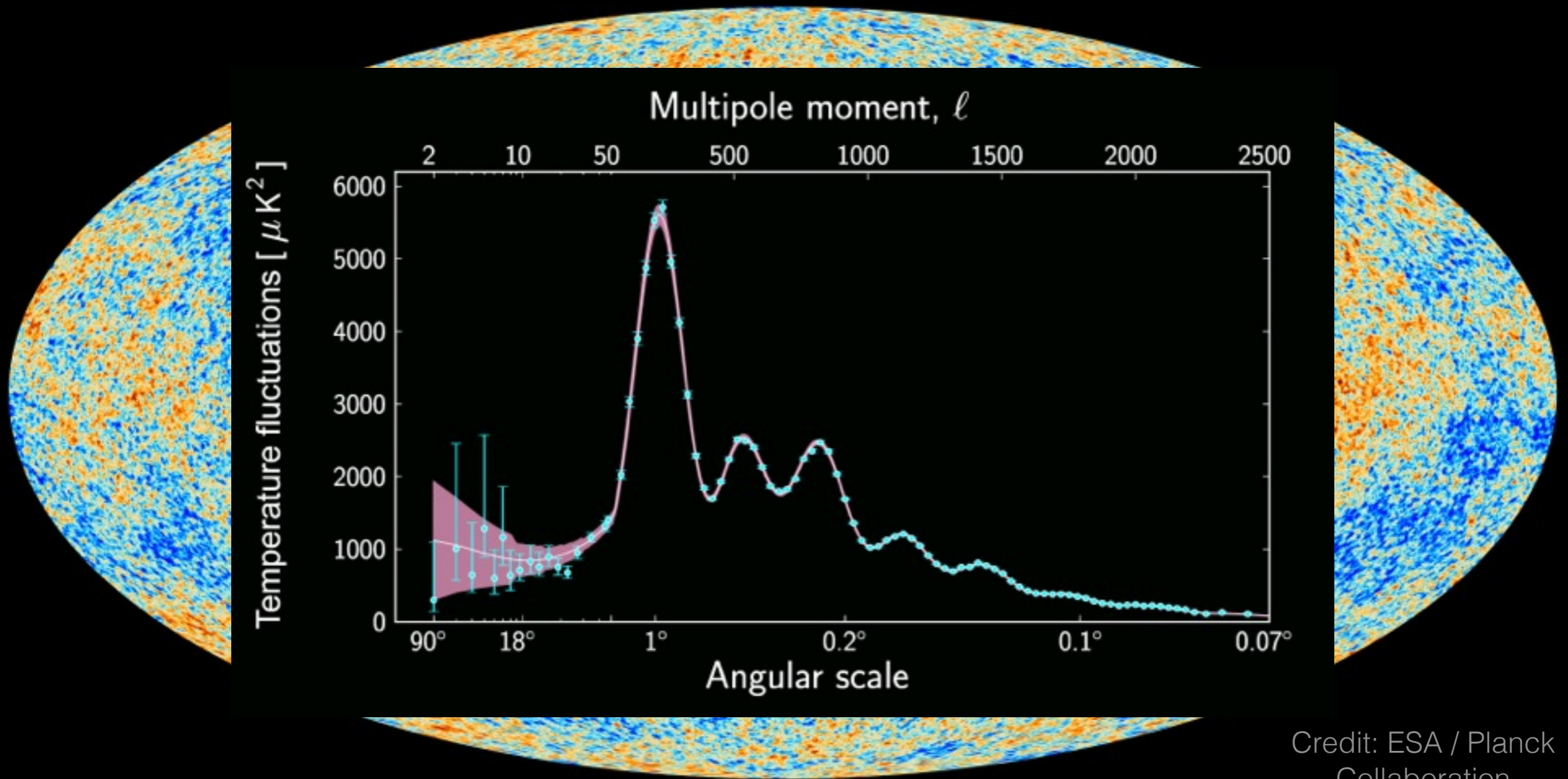


Map of the CMB

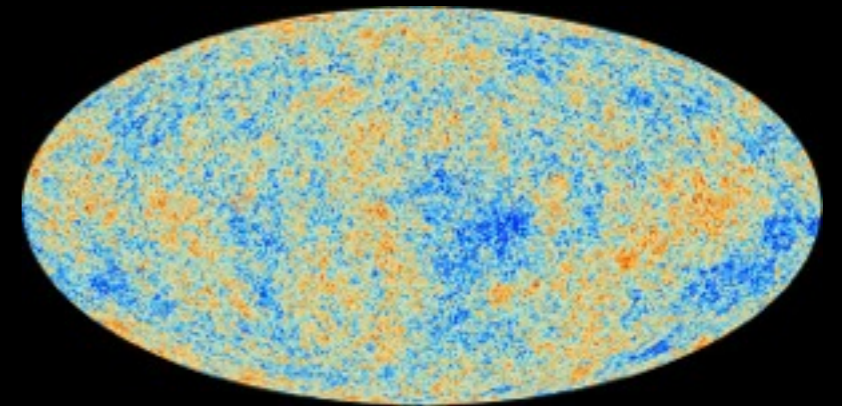
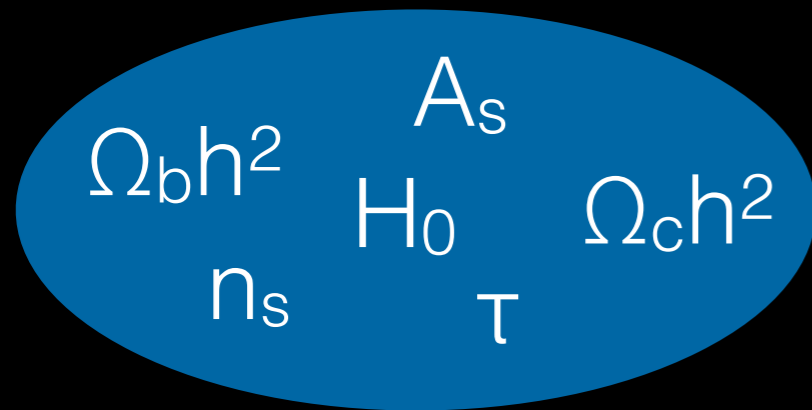


Credit: ESA / Planck
Collaboration

Map of the CMB

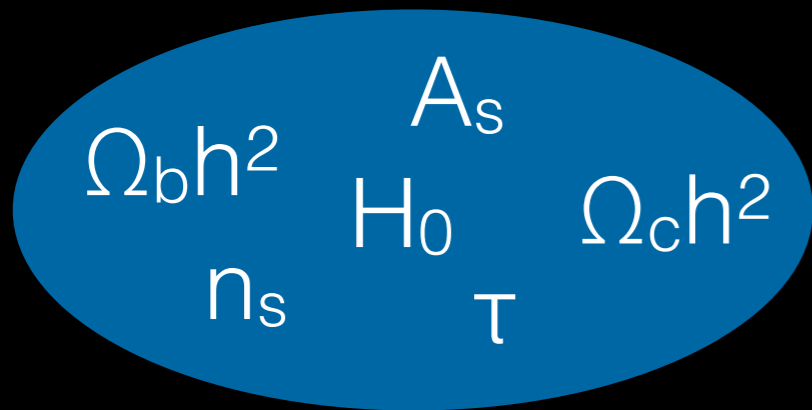


Parameter Estimation

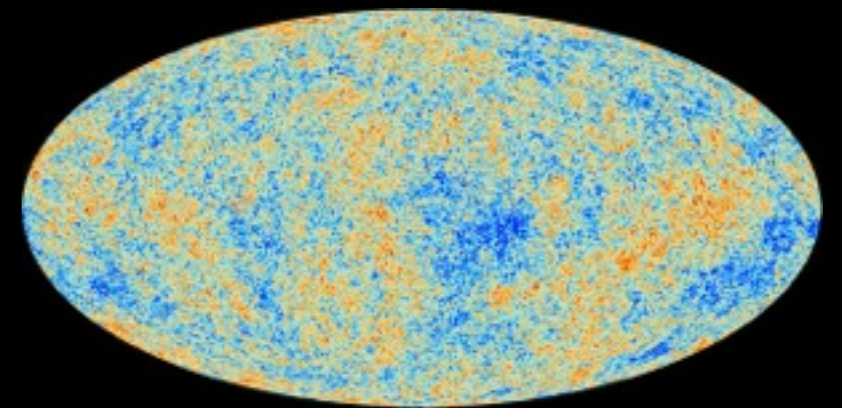


Credit: ESA / Planck
Collaboration

Parameter Estimation



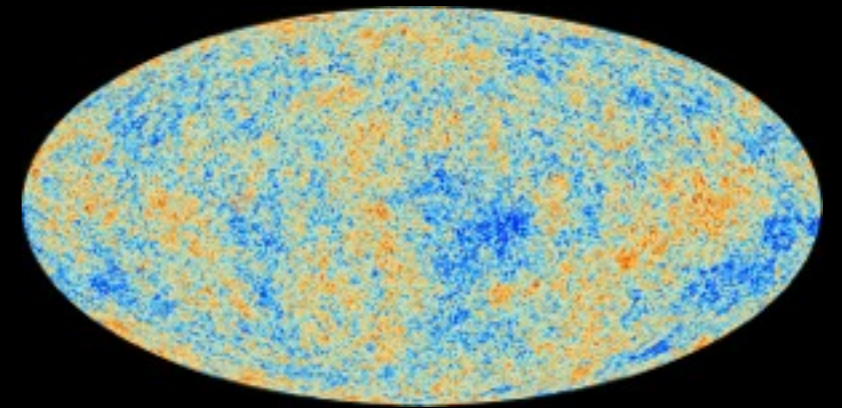
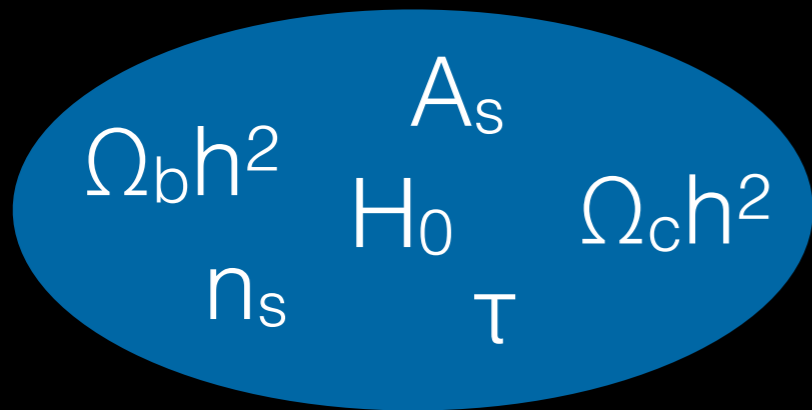
Predictions
from Theory



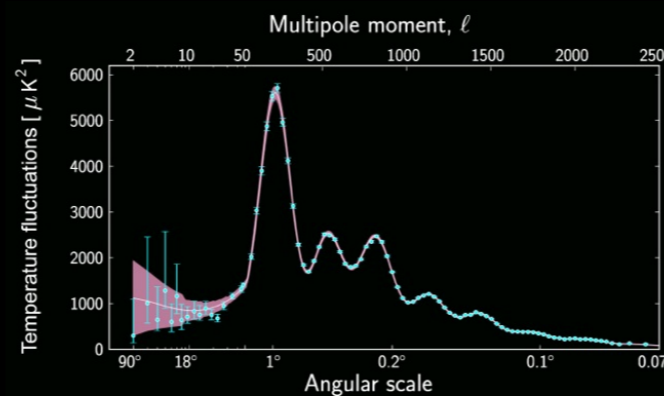
Data and
Errors

Credit: ESA / Planck
Collaboration

Parameter Estimation



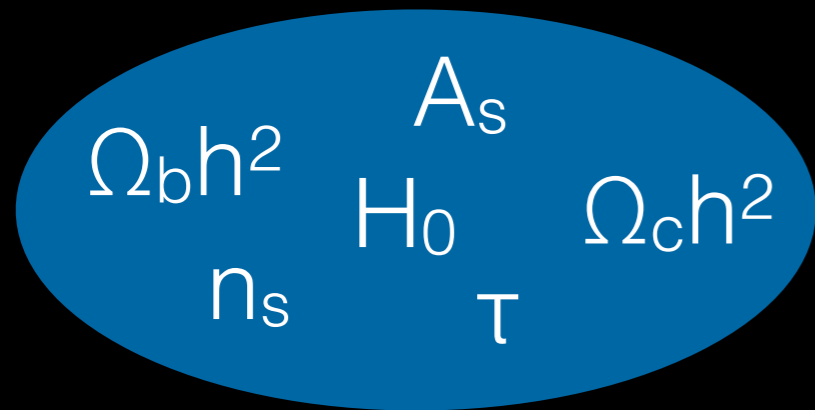
Predictions
from Theory



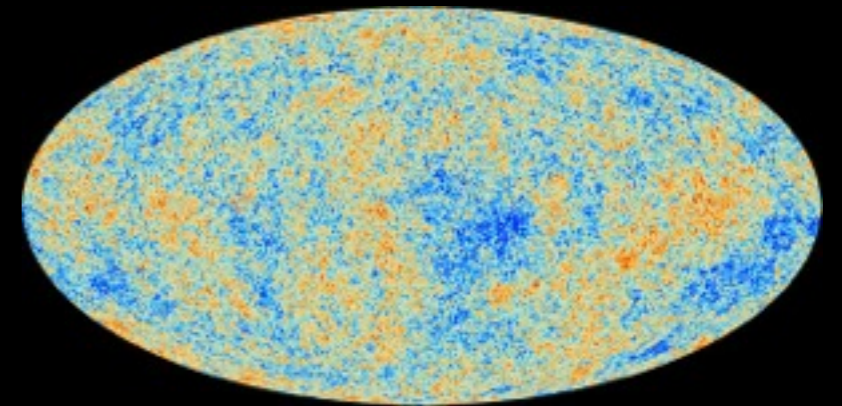
Data and
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Collaboration

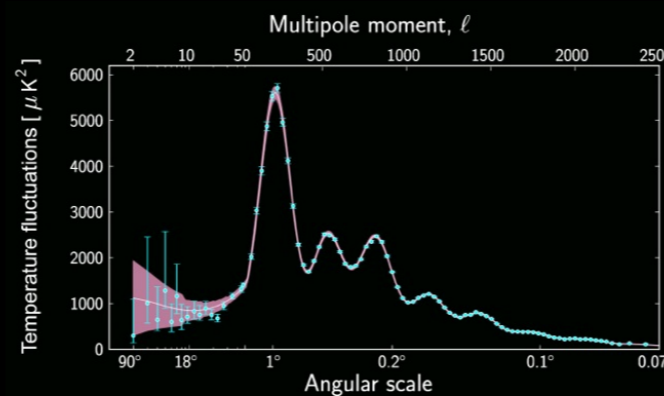
Parameter Estimation



Parameter Estimation



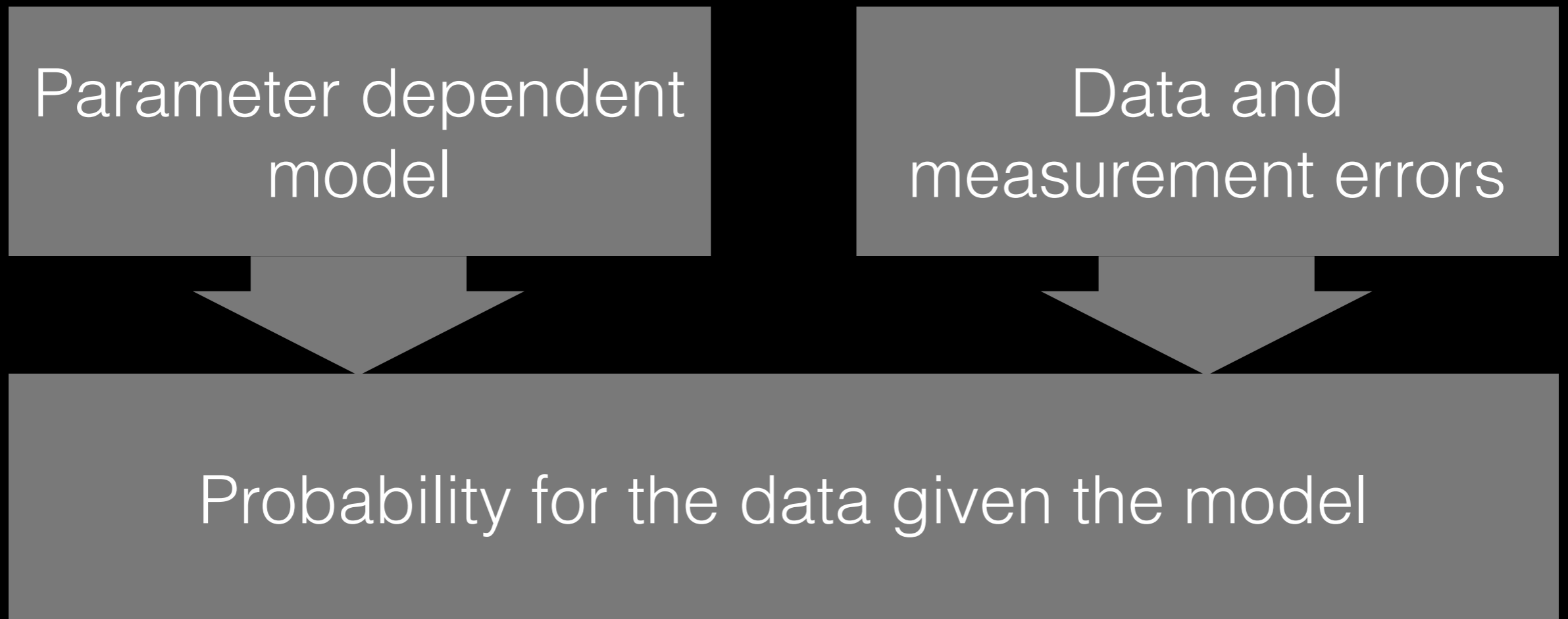
Predictions from Theory



Data and Errors

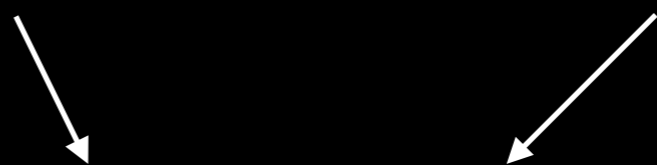
Credit: ESA / Planck Collaboration

Likelihood



Bayesian Inference

Likelihood Prior

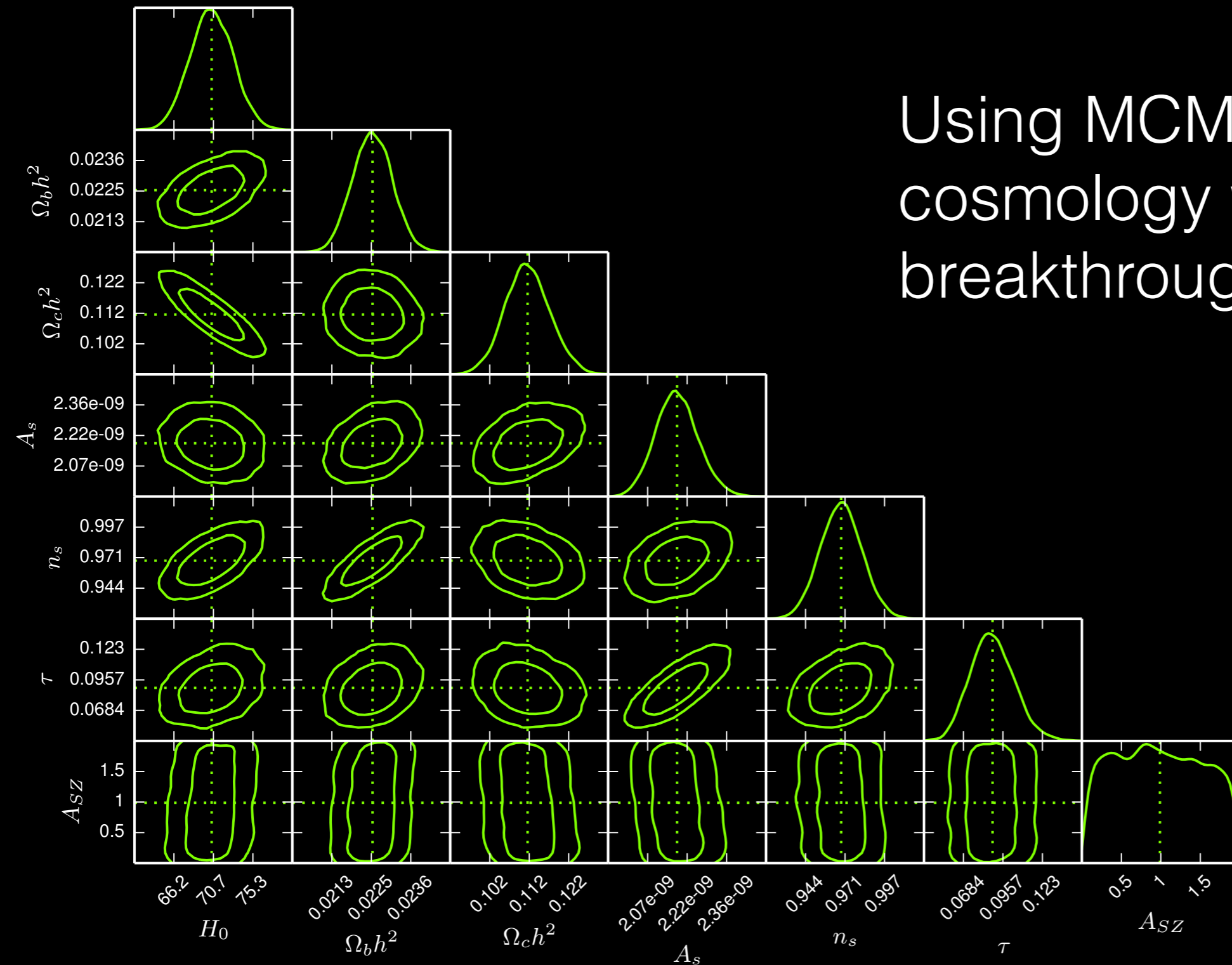

$$p_{new}(\Theta) = \frac{p(D_{new} | \Theta) p(\Theta)}{\int d\Theta p(D_{new} | \Theta) p(\Theta)}$$

Bayesian Inference in Cosmology

- Parameter space is greater or equal to six
- Model predictions have to be evaluated numerically
- The likelihood is computationally intensive

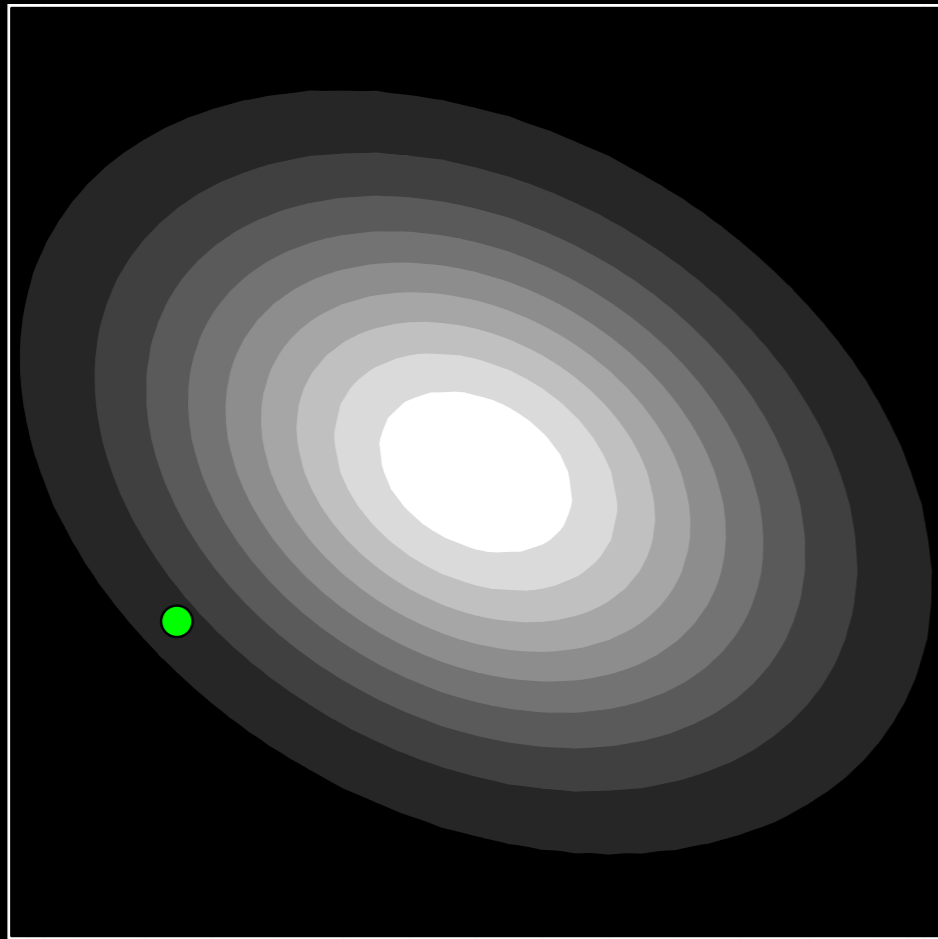
MCMC in Cosmology

Using MCMC in cosmology was a major breakthrough in 2001



Metropolis-Hasting

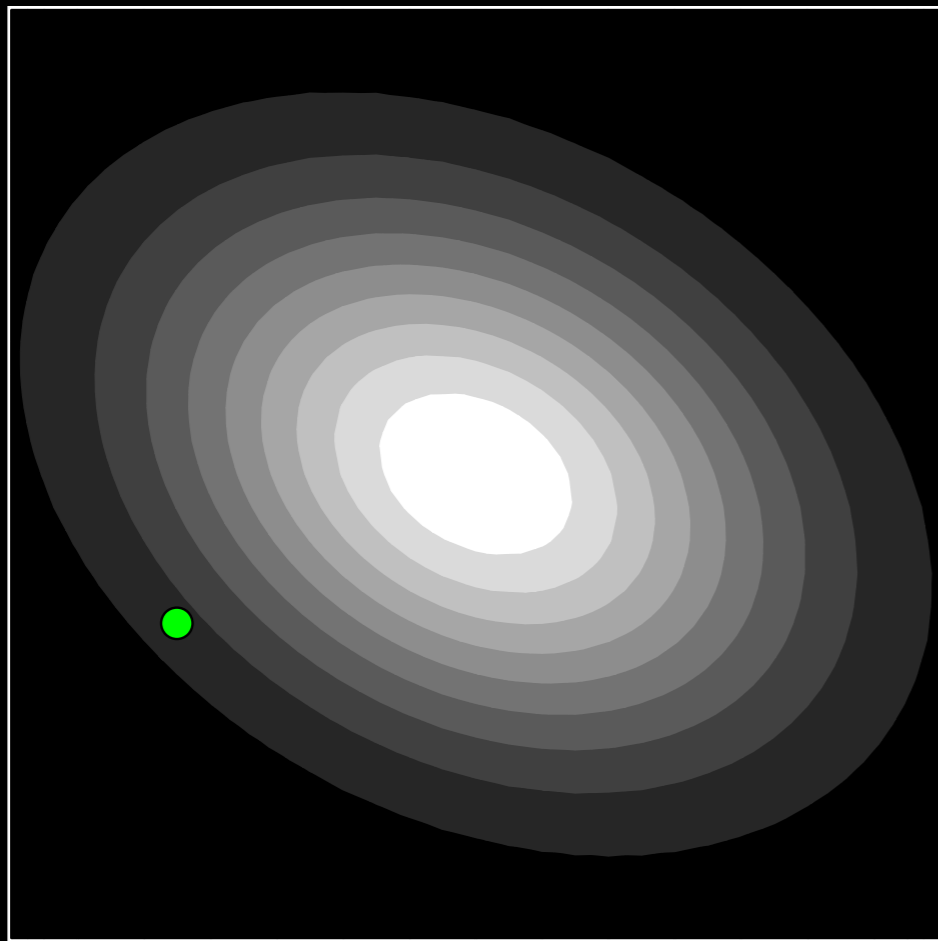
Metropolis+ 1953
Hastings 1970



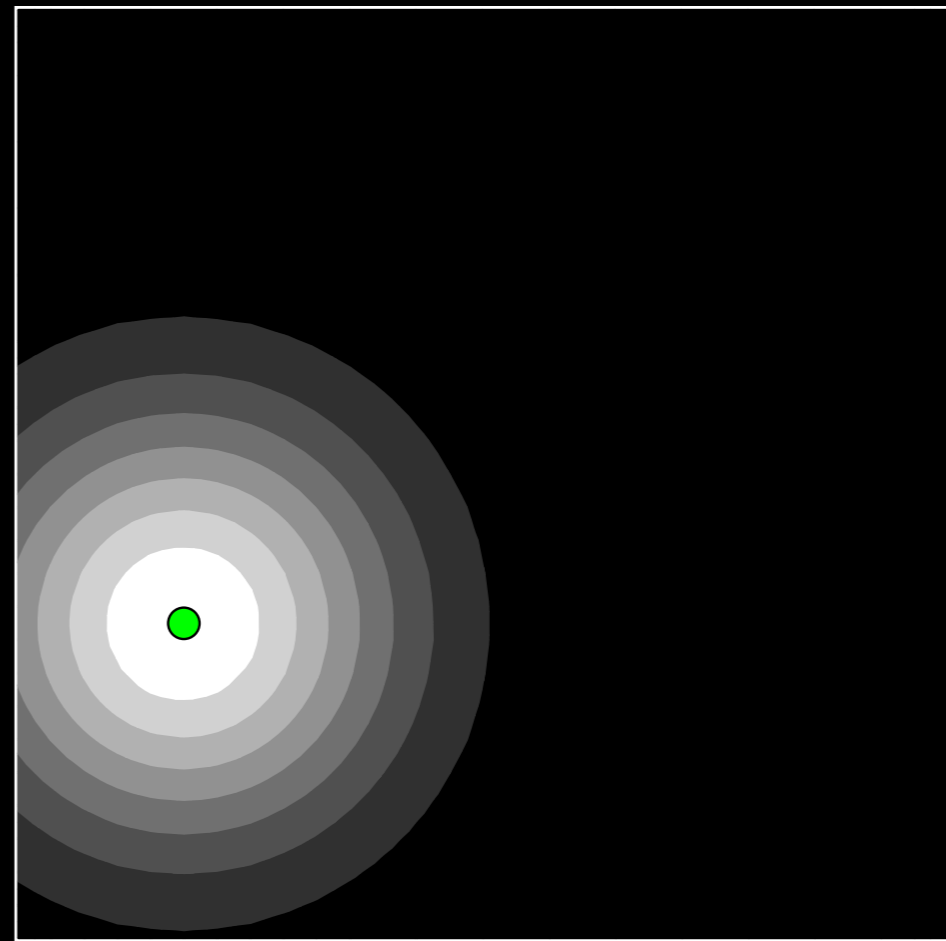
target distribution $P(\theta)$

Metropolis-Hasting

Metropolis+ 1953
Hastings 1970



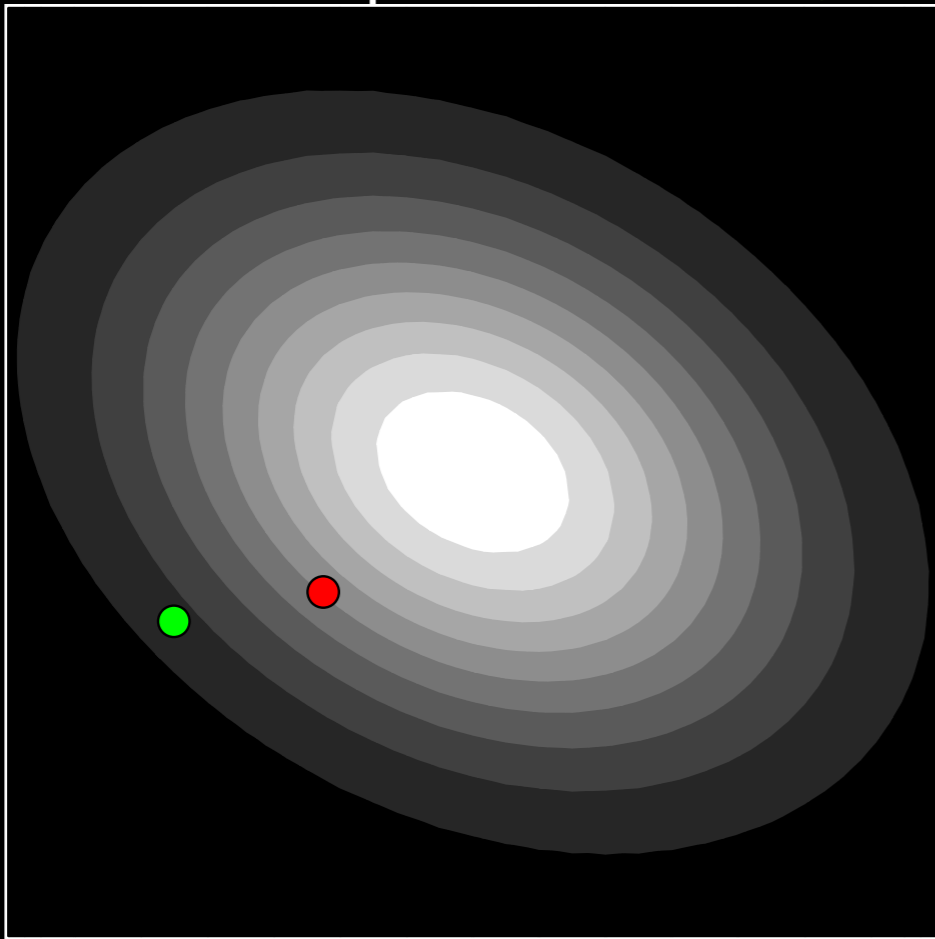
target distribution $P(\theta)$



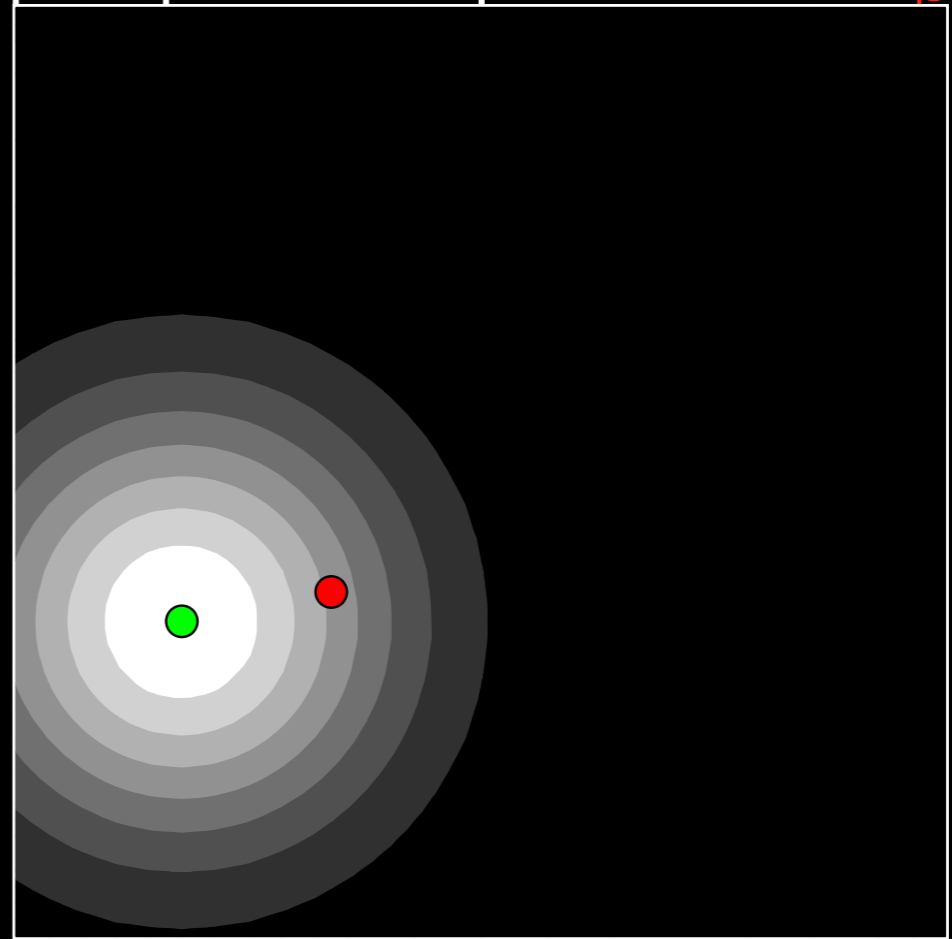
proposal distribution $Q(\theta)$

Metropolis-Hasting

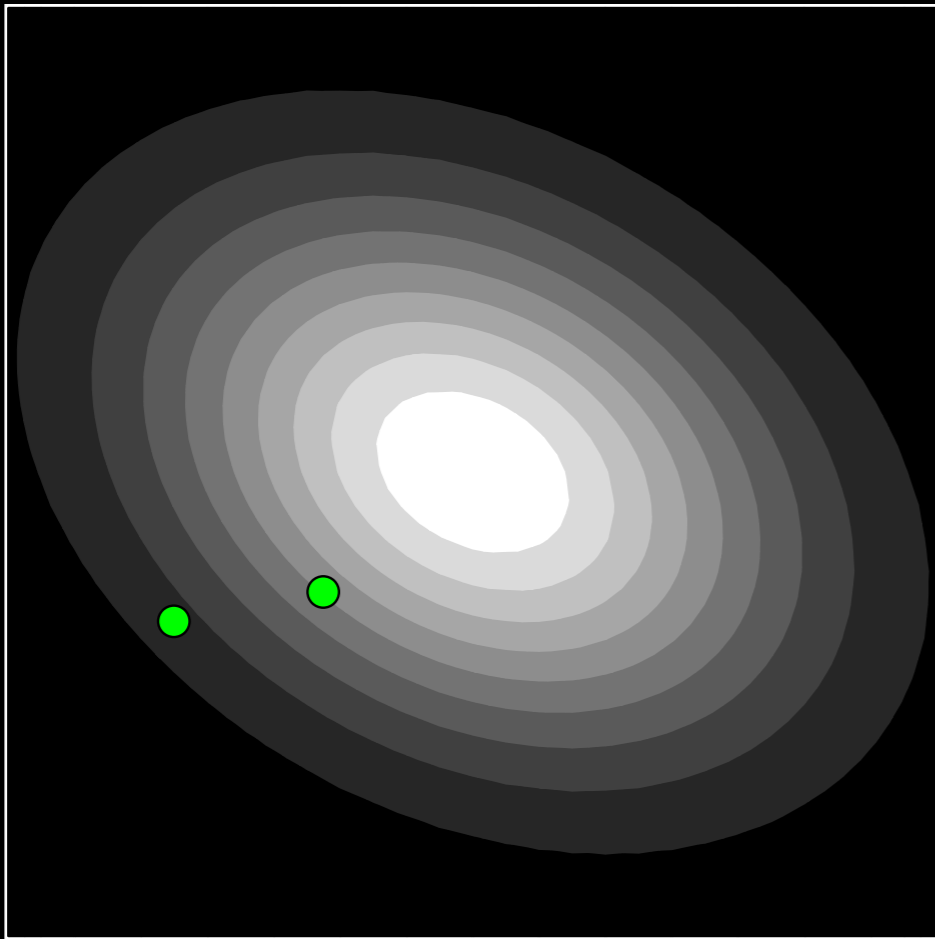
initial position θ_0



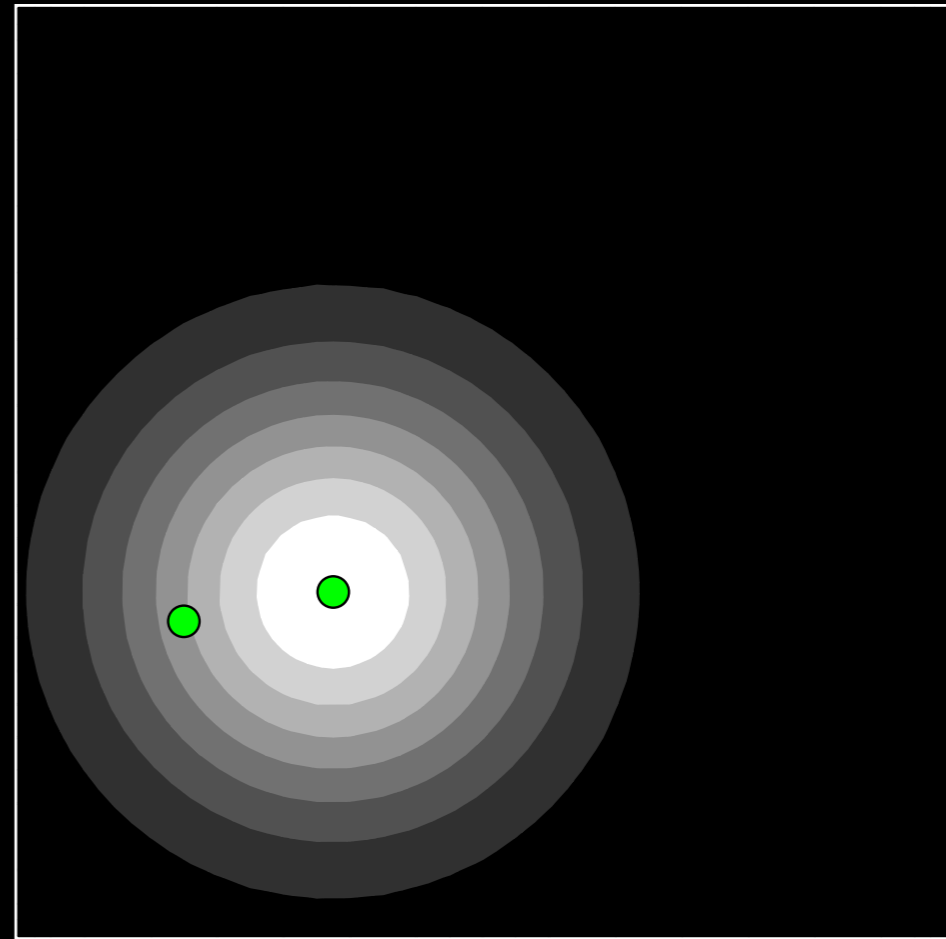
proposed position θ_p



Metropolis-Hasting

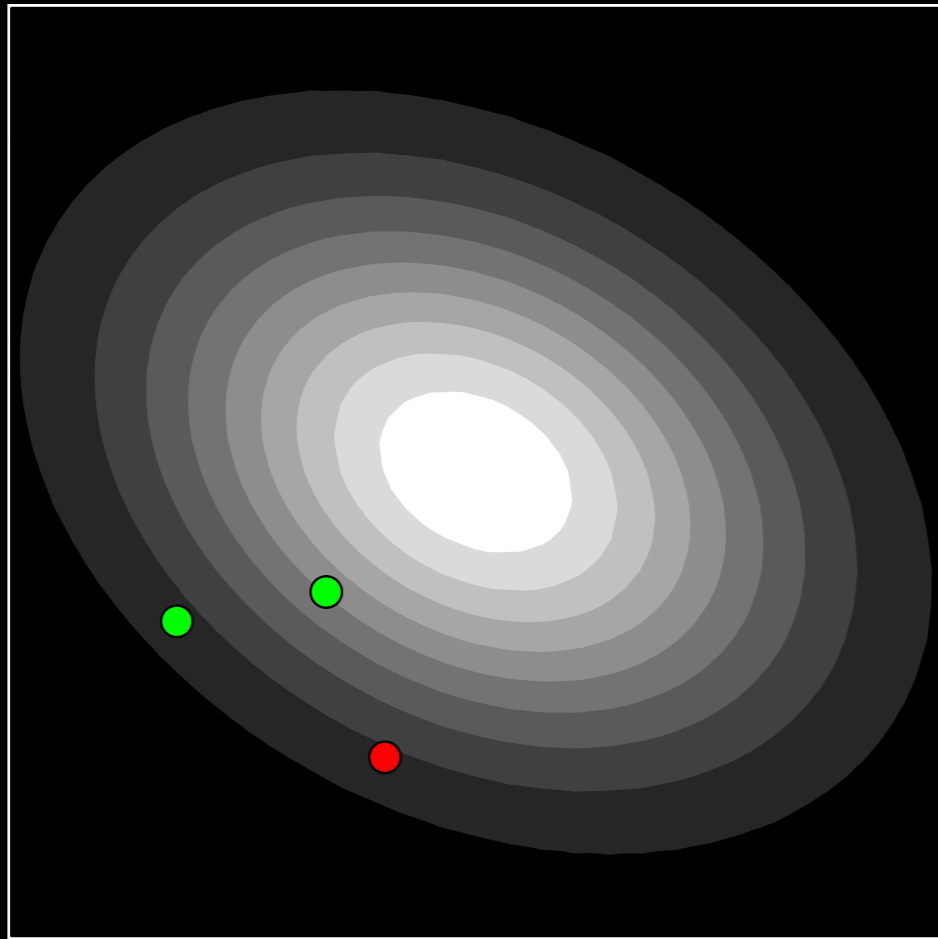


target distribution $P(\theta)$

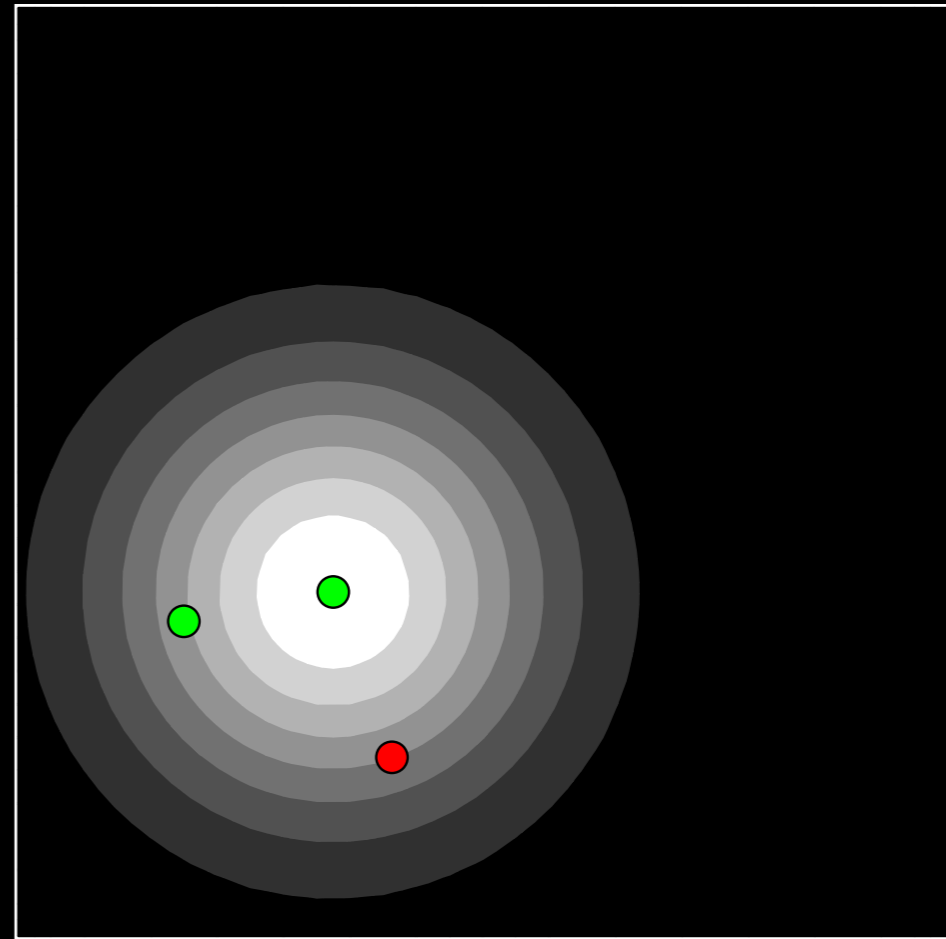


proposal distribution $Q(\theta)$

Metropolis-Hasting

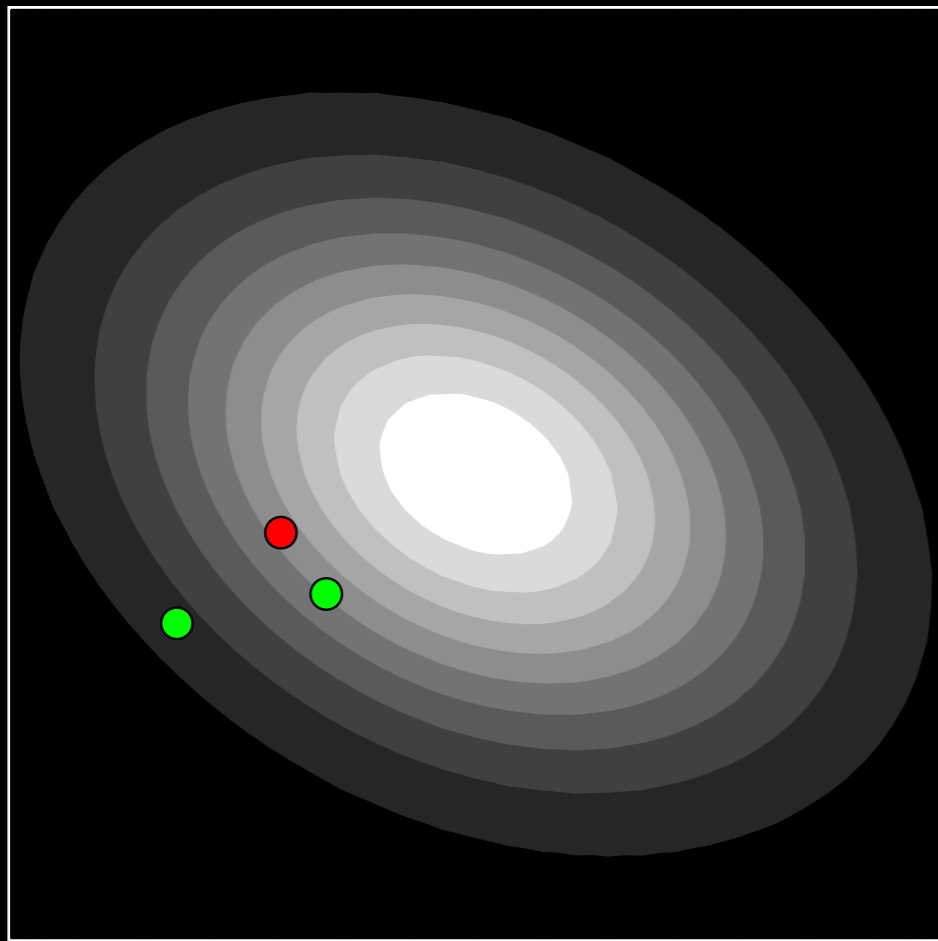


target distribution $P(\theta)$

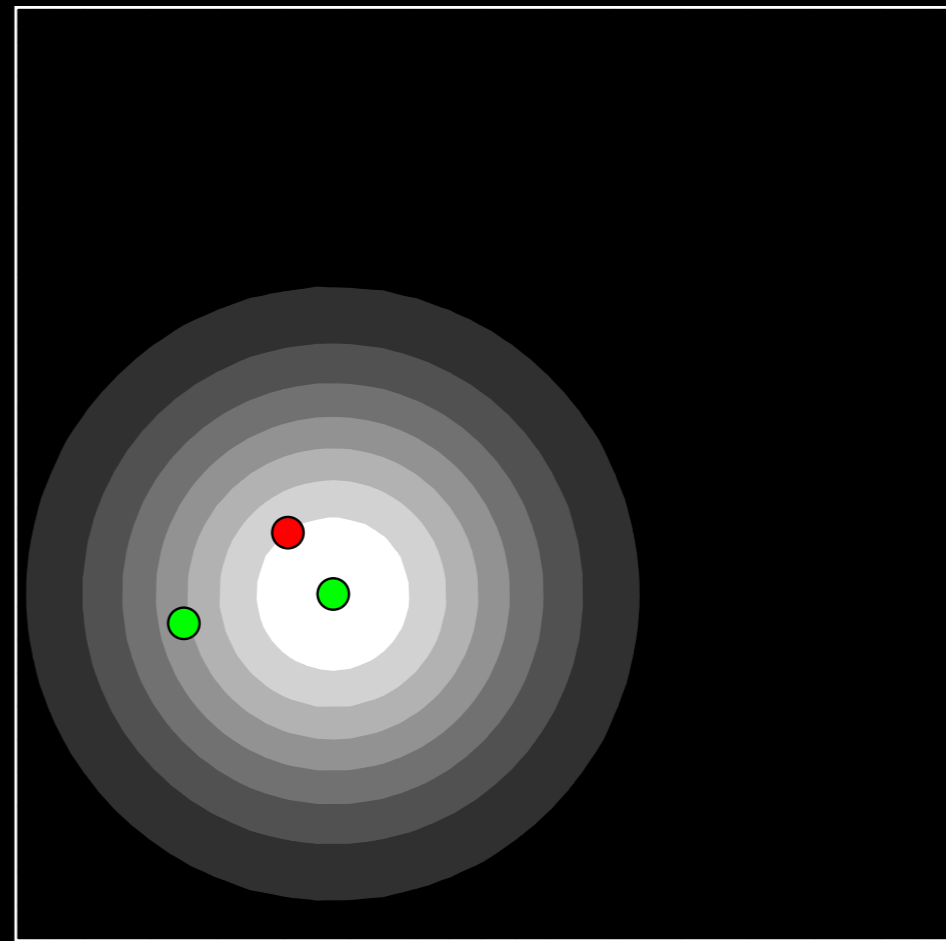


proposal distribution $Q(\theta)$

Metropolis-Hasting

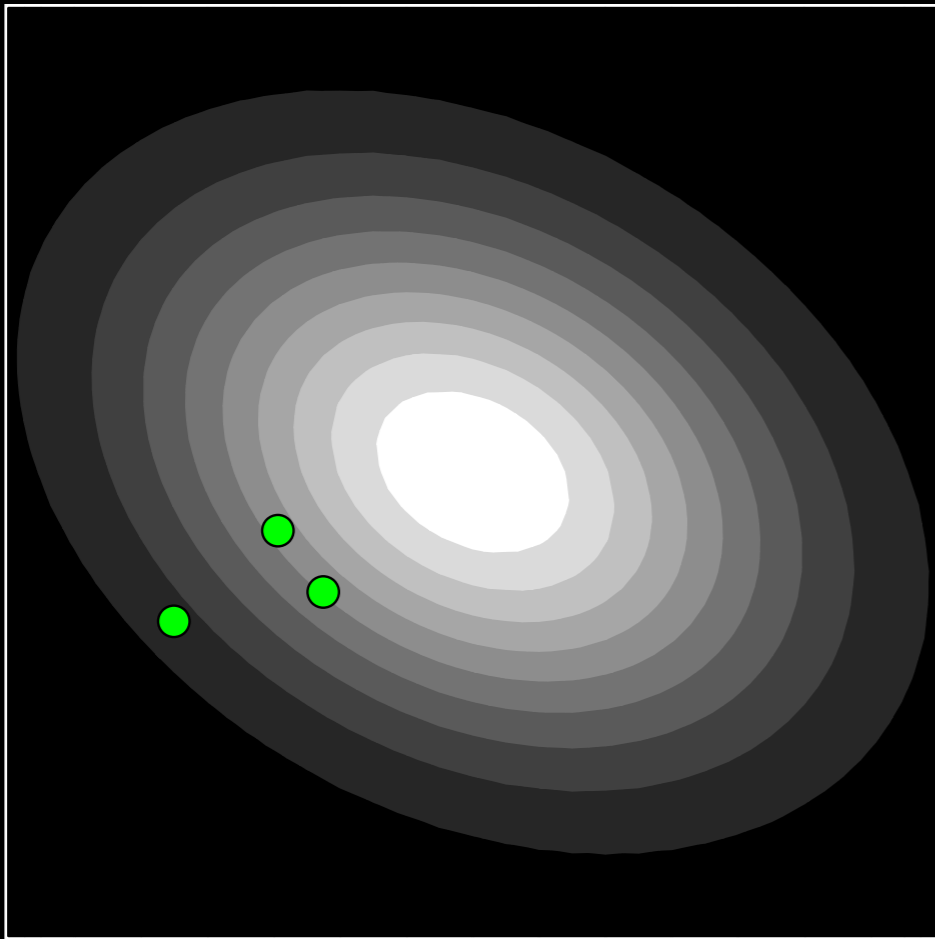


target distribution $P(\theta)$

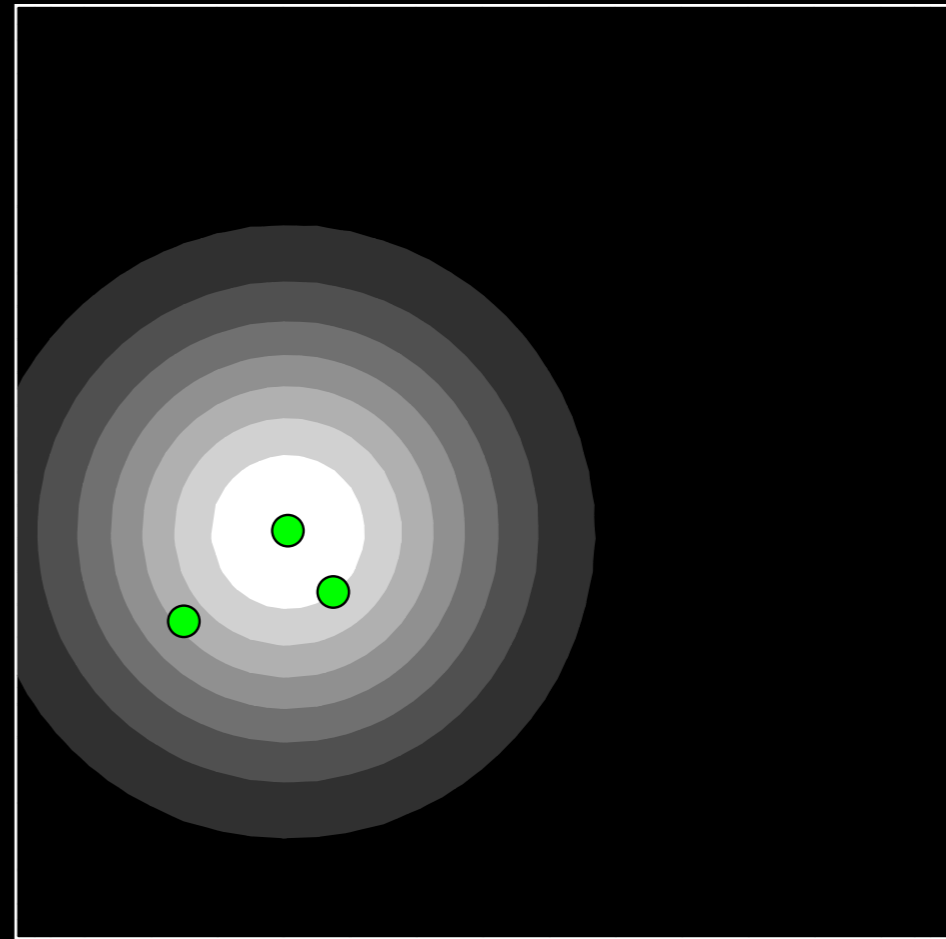


proposal distribution $Q(\theta)$

Metropolis-Hasting

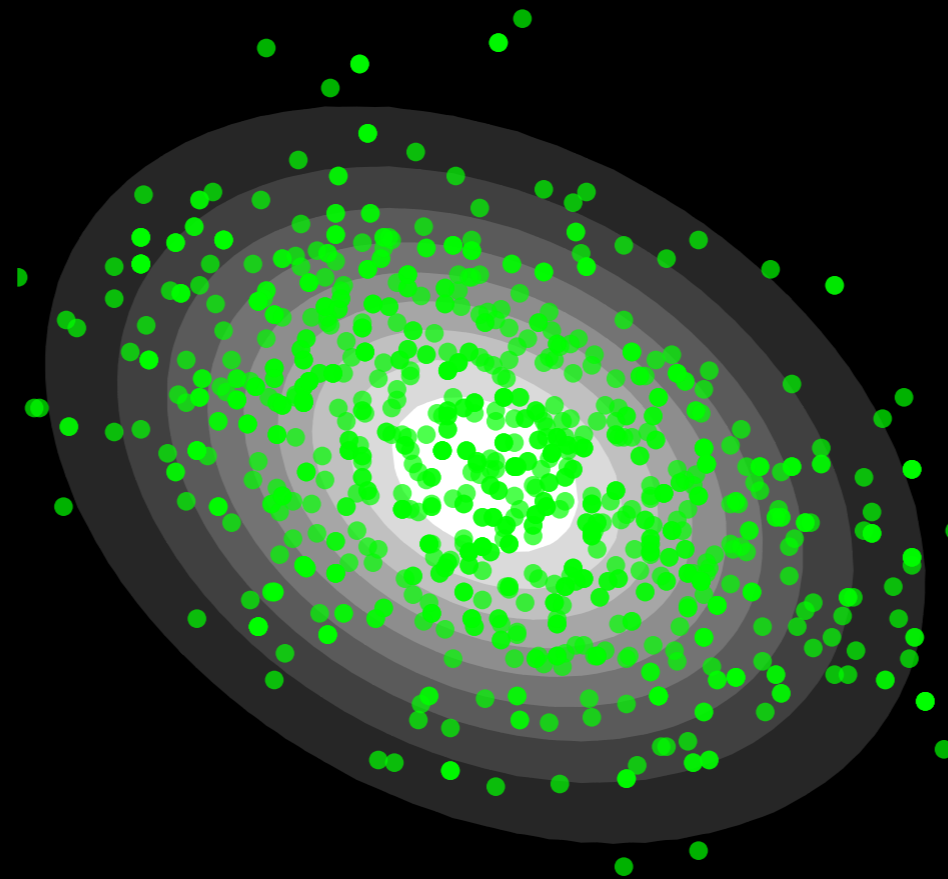


target distribution $P(\theta)$



proposal distribution $Q(\theta)$

Metropolis-Hasting



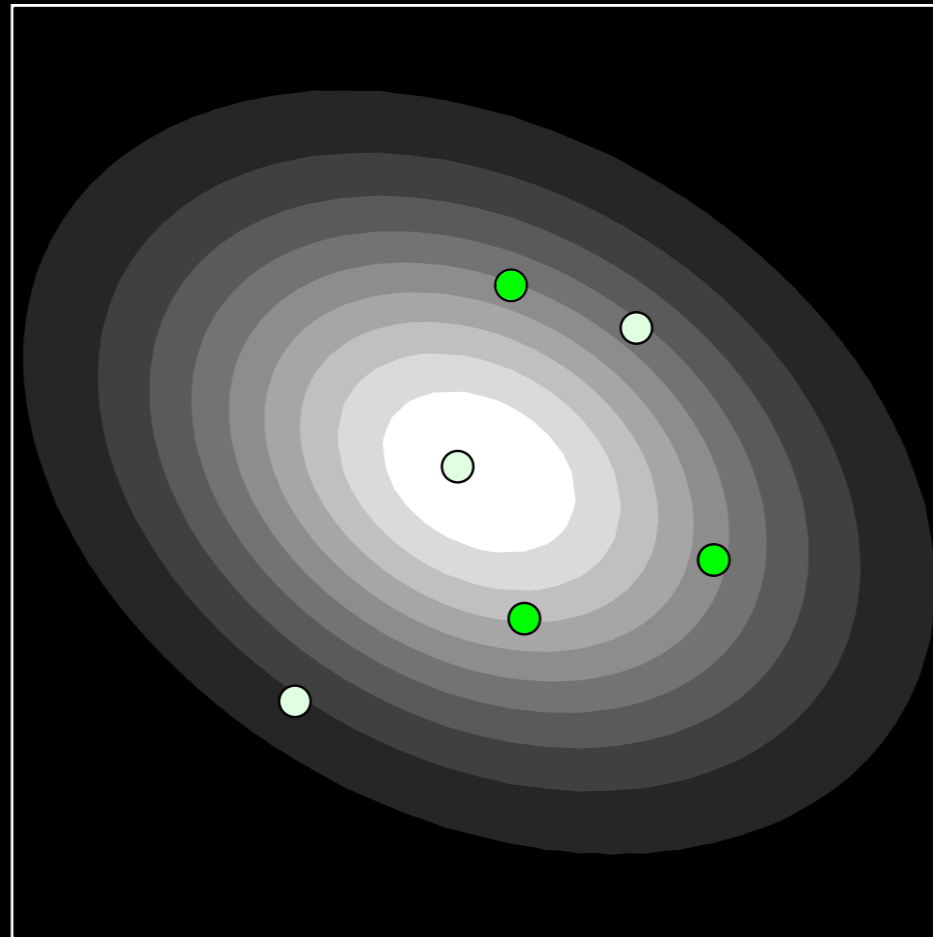
after 1,000 iterations

MCMC in Cosmology

- Estimation of cosmological parameters with MCMC is time consuming
- Multiple runs of the same likelihood
 - Learn about different models and configurations

emcee

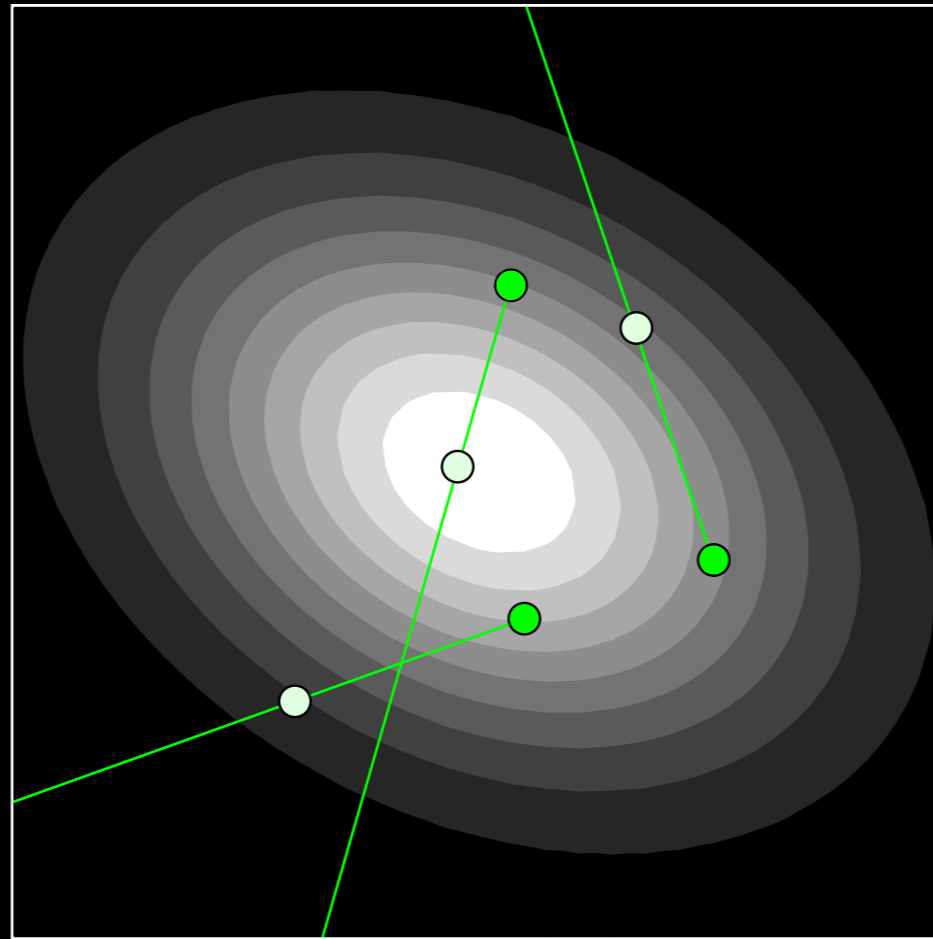
Foreman-Mackey+
2012
Goodman+ 2010



initial positions $\theta_0^1, \dots, \theta_0^m$

emcee

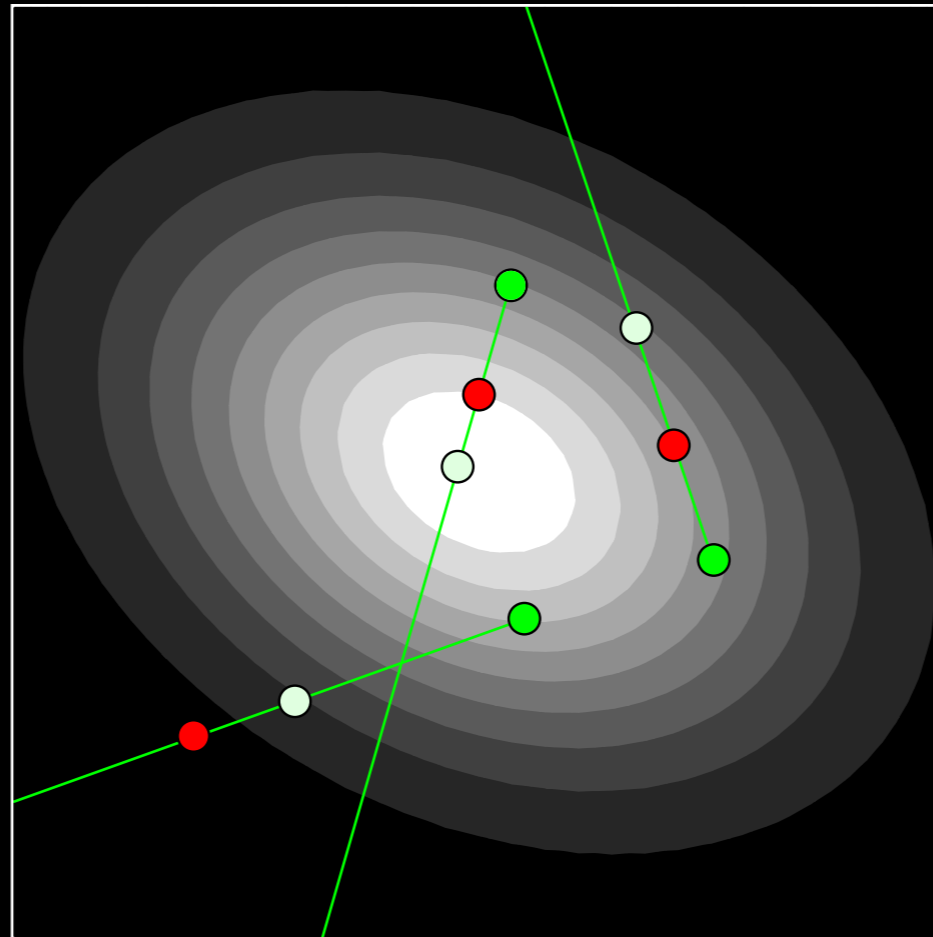
Foreman-Mackey+
2012
Goodman+ 2010



pick a partner at random

emcee

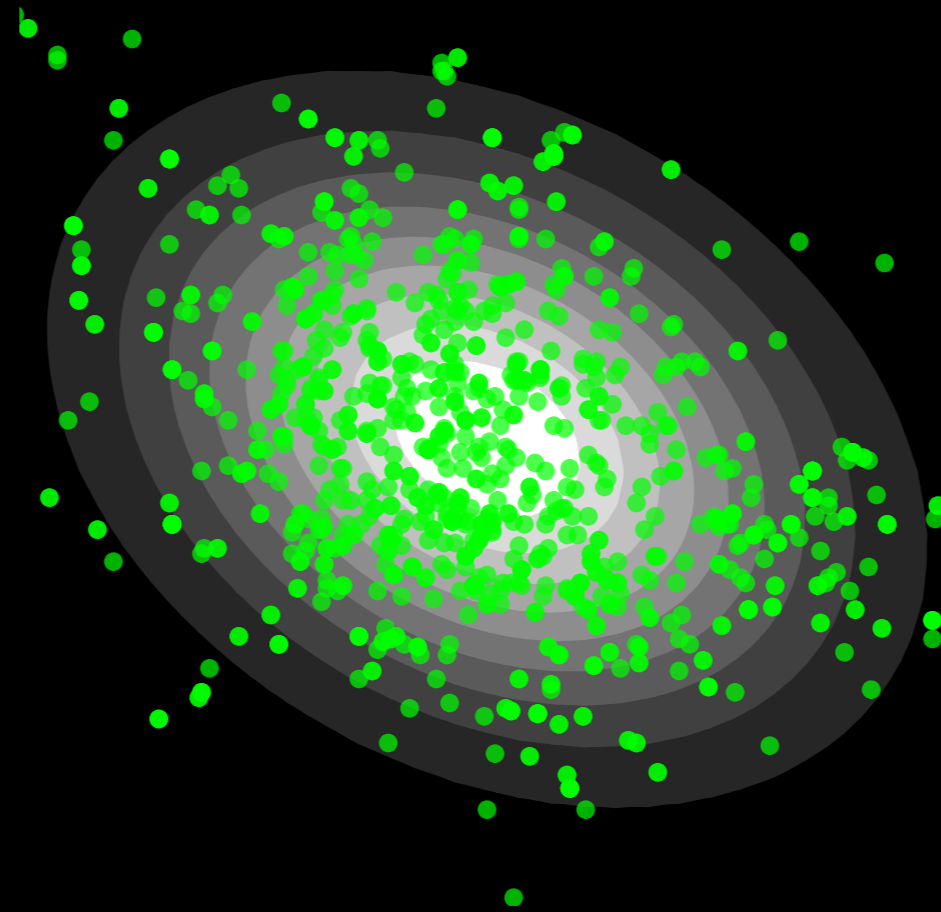
Foreman-Mackey+
2012
Goodman+ 2010



propose positions on connecting rays

emcee

Foreman-Mackey+
2012
Goodman+ 2010

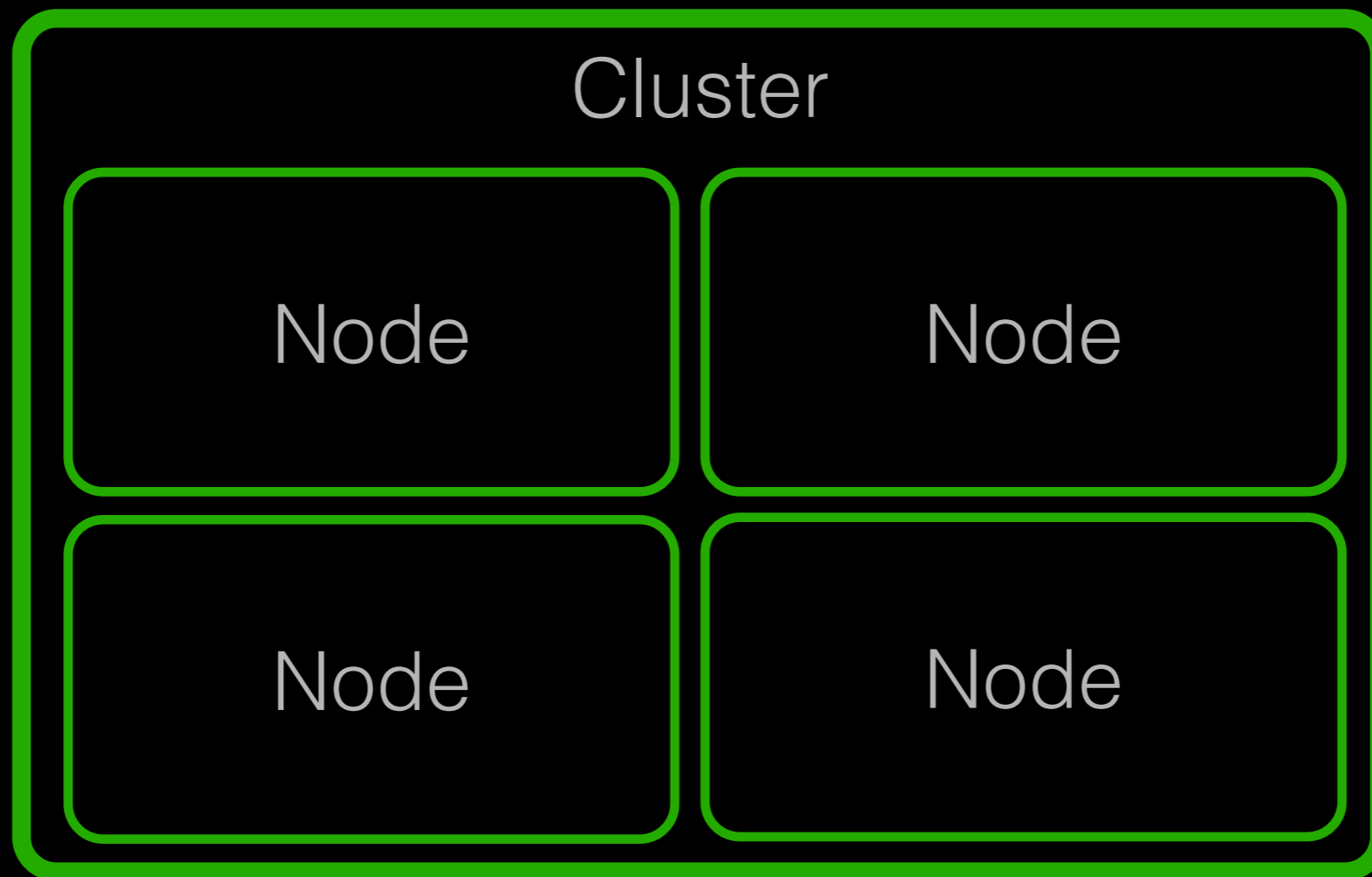


after 1,000 iterations

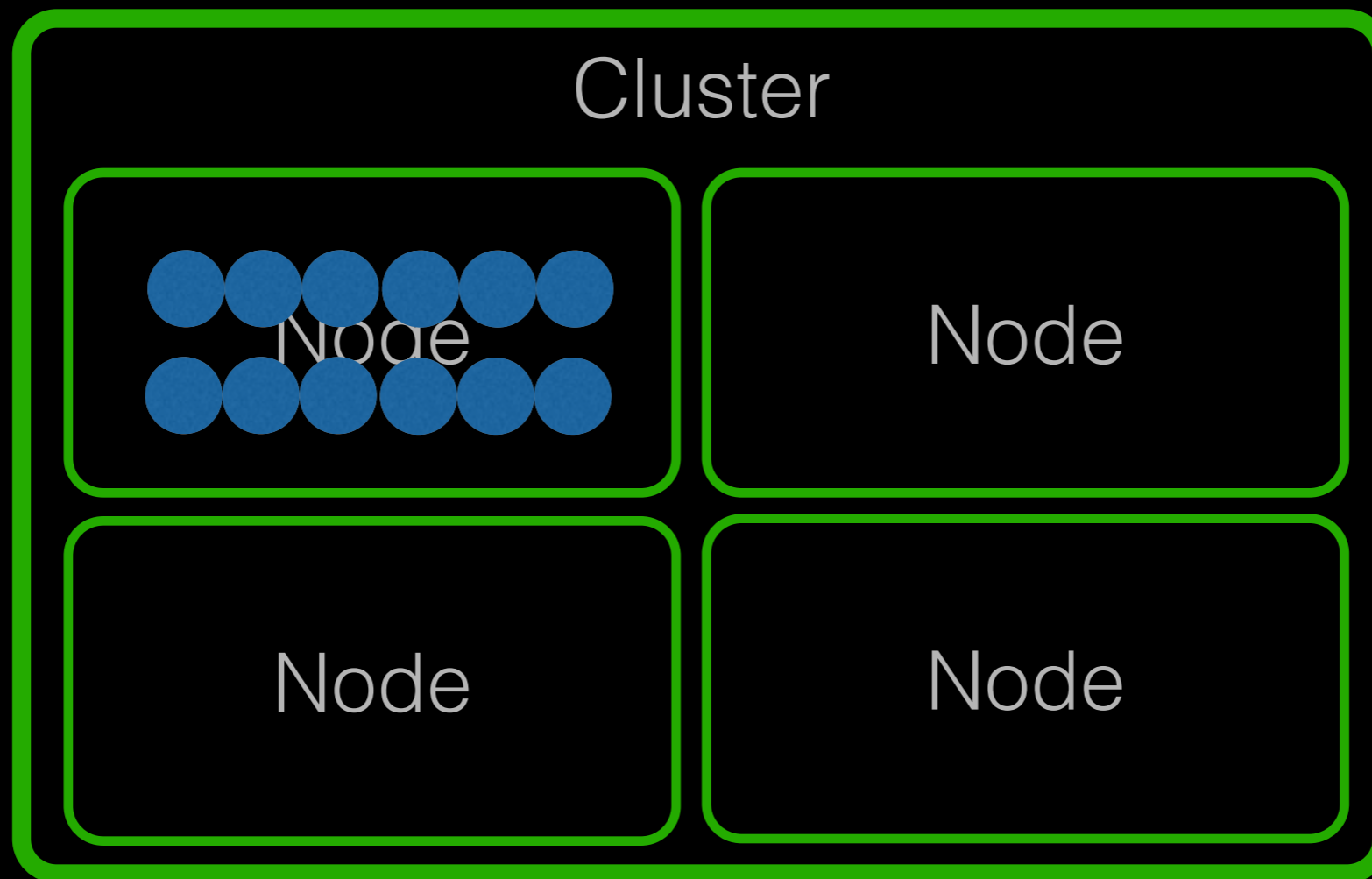
CosmoHammer

- Python framework for parallelised MCMC sampling with emcee
 - Flexible architecture to sample various models and data sets
- Distributes the workload over multiple nodes in a compute cluster
 - Built on MPI and Python multiprocessing

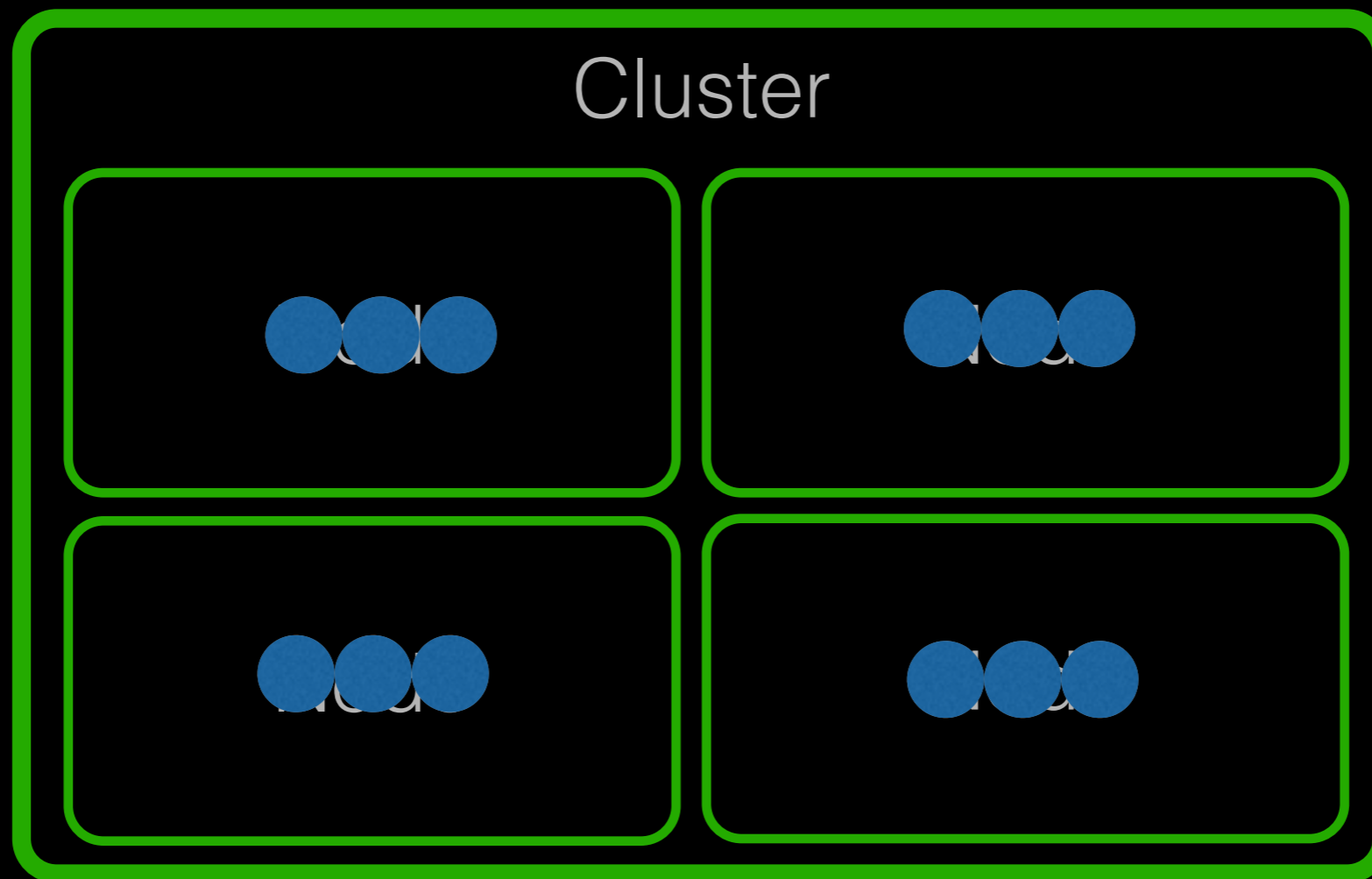
CosmoHammer



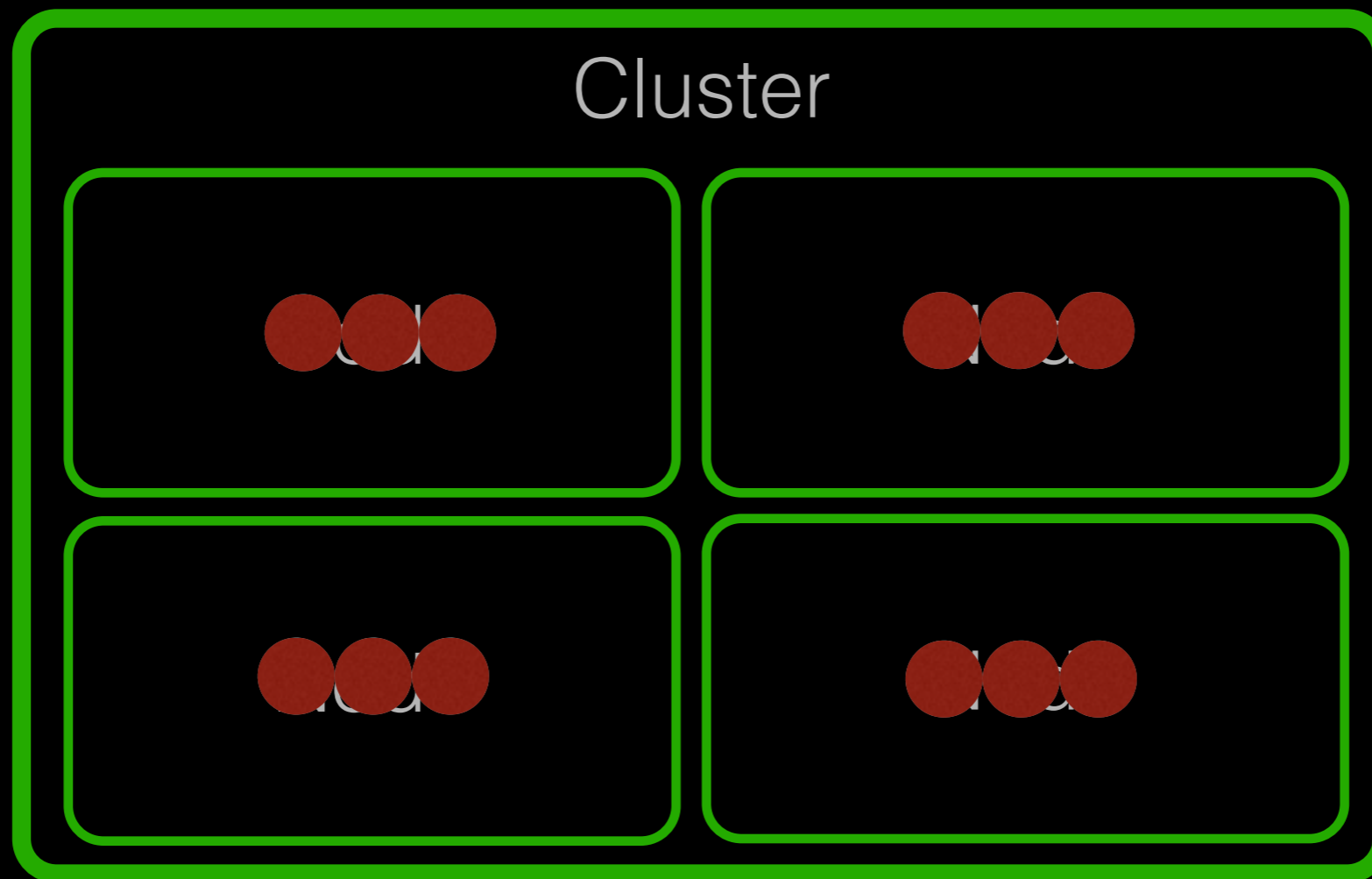
CosmoHammer



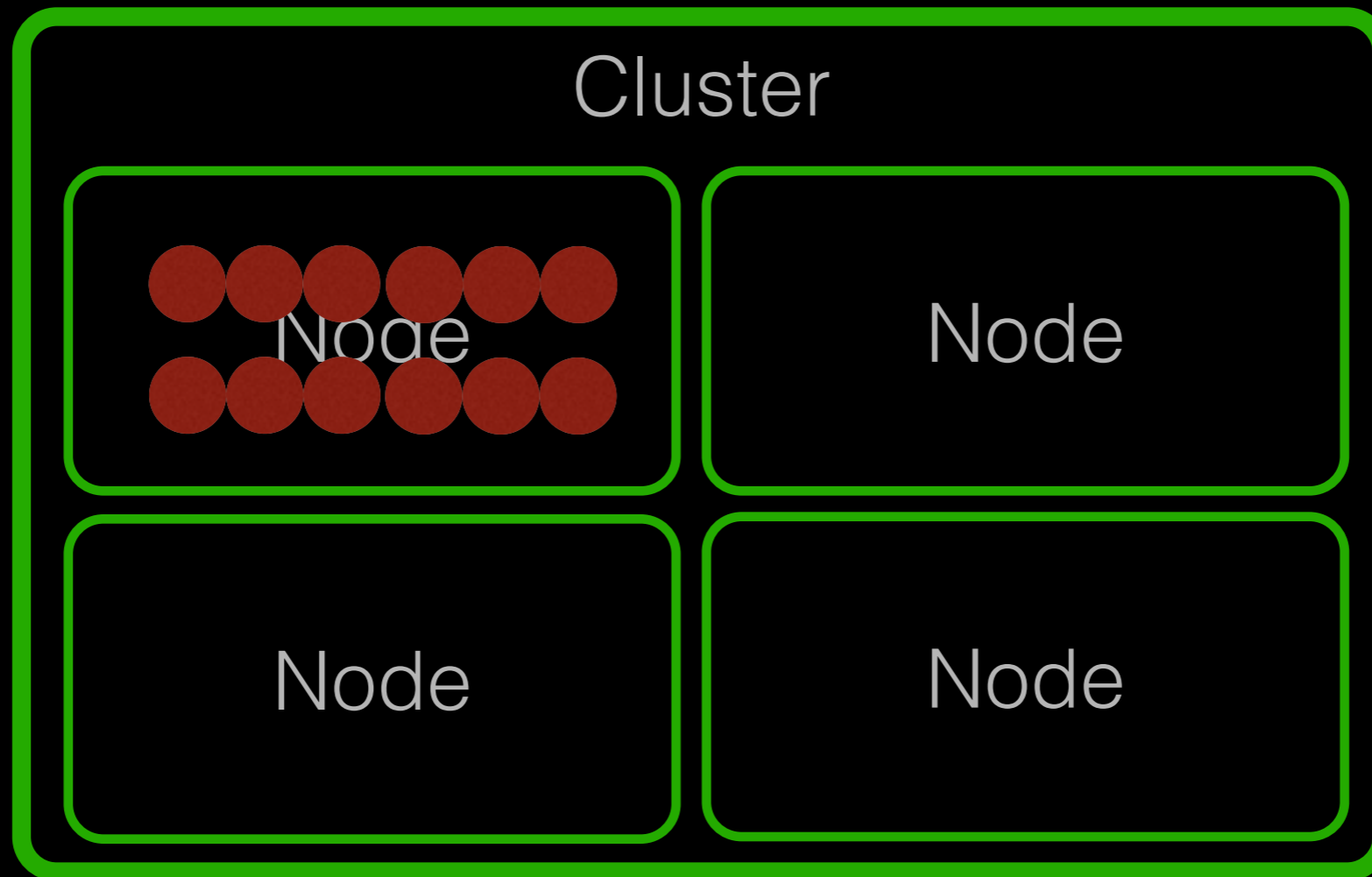
CosmoHammer



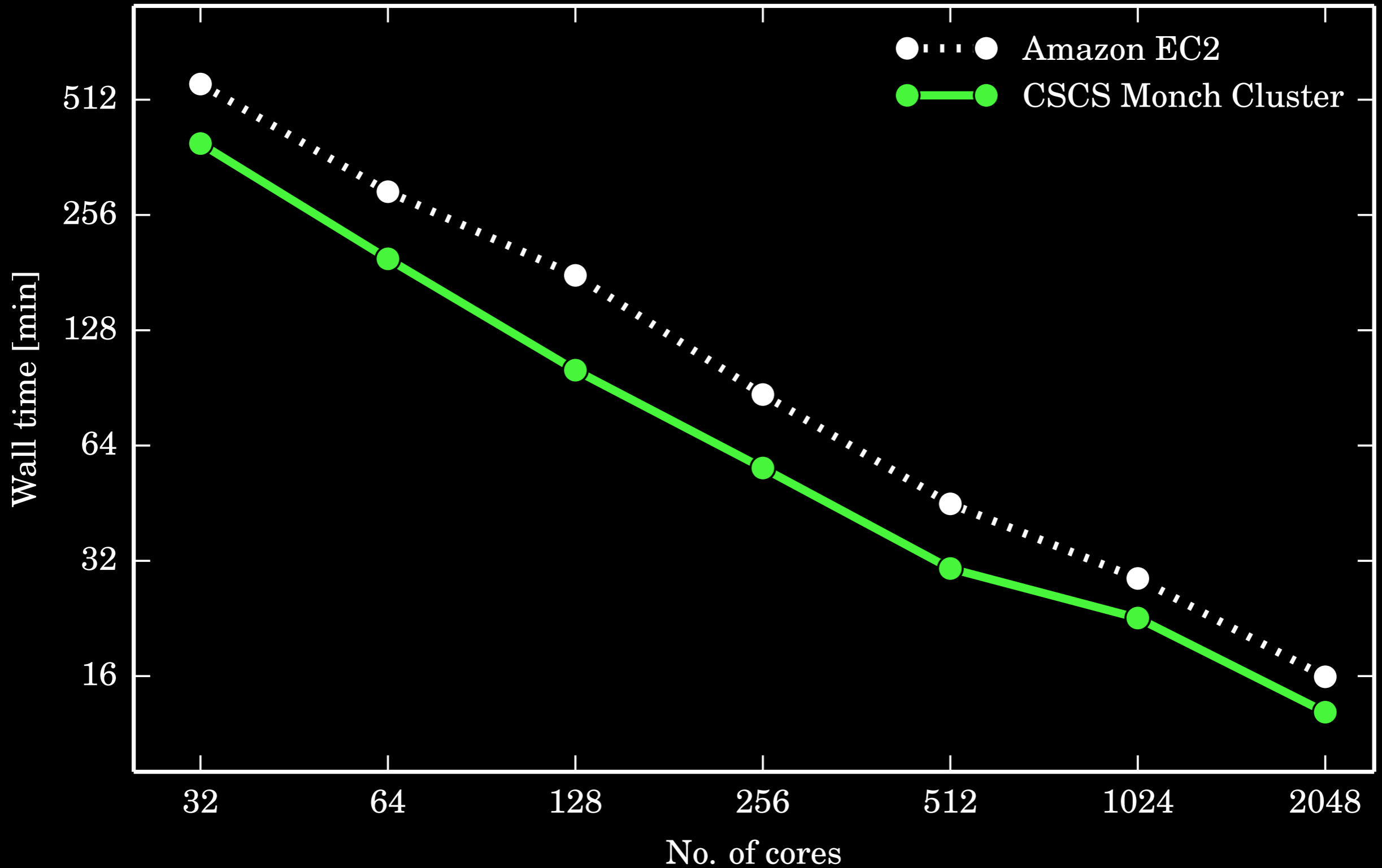
CosmoHammer



CosmoHammer

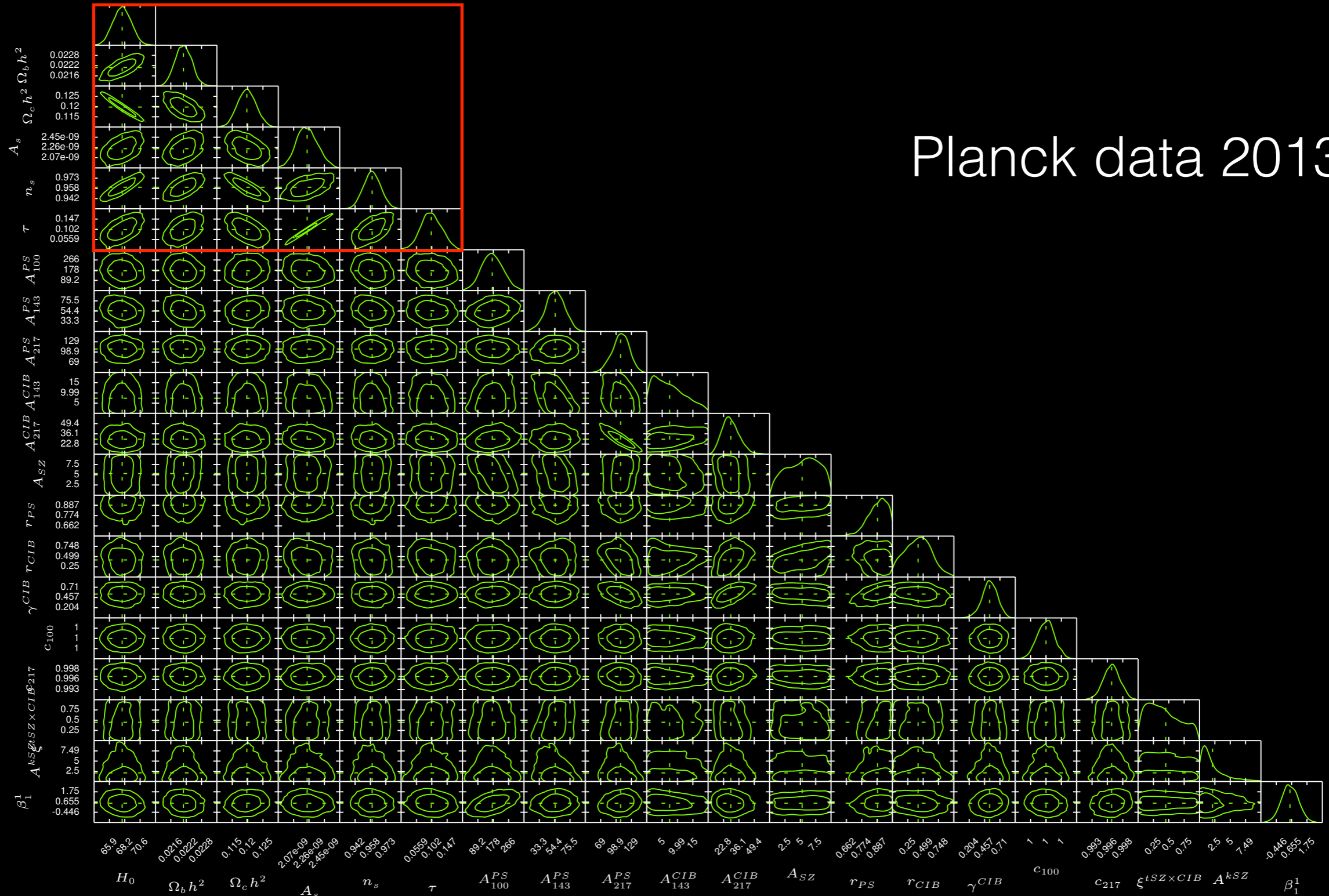


CosmoHammer



Current challenges

Planck data 2013

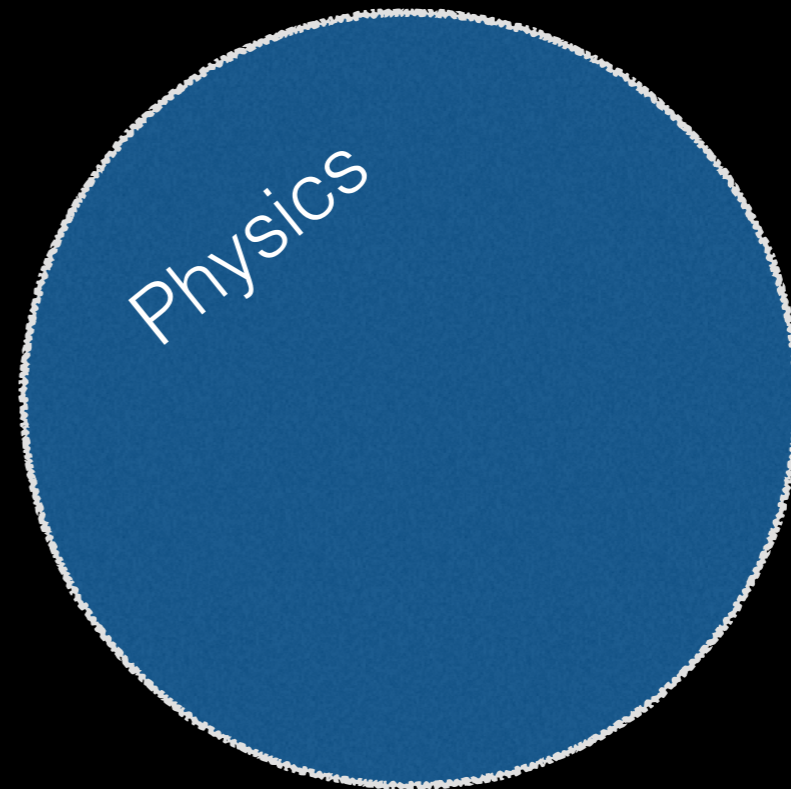


Current challenges

- Fast sampling in high dimensions
- Diagnostics & Convergence criteria
- Handling of systematic effects

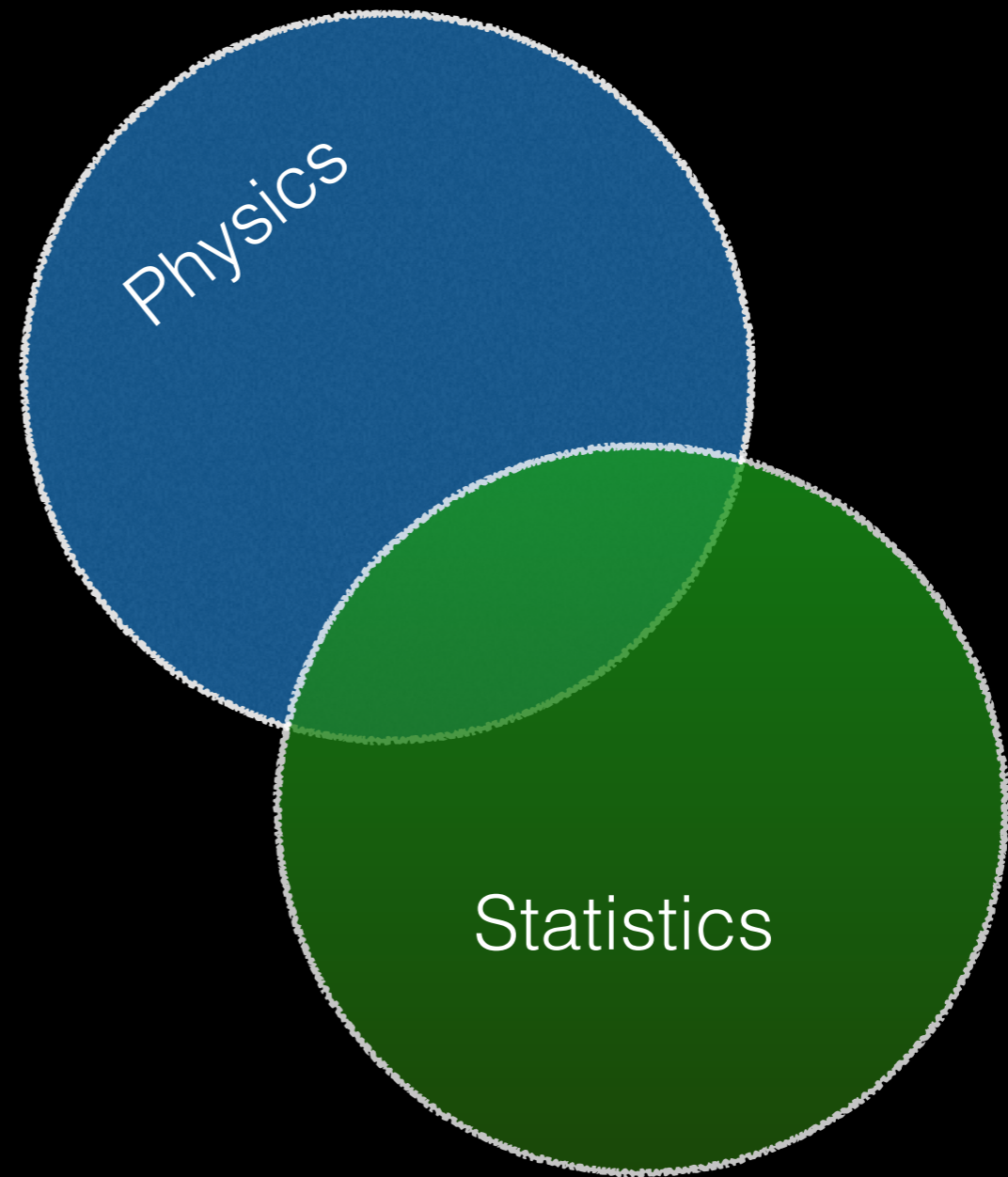
Current challenges

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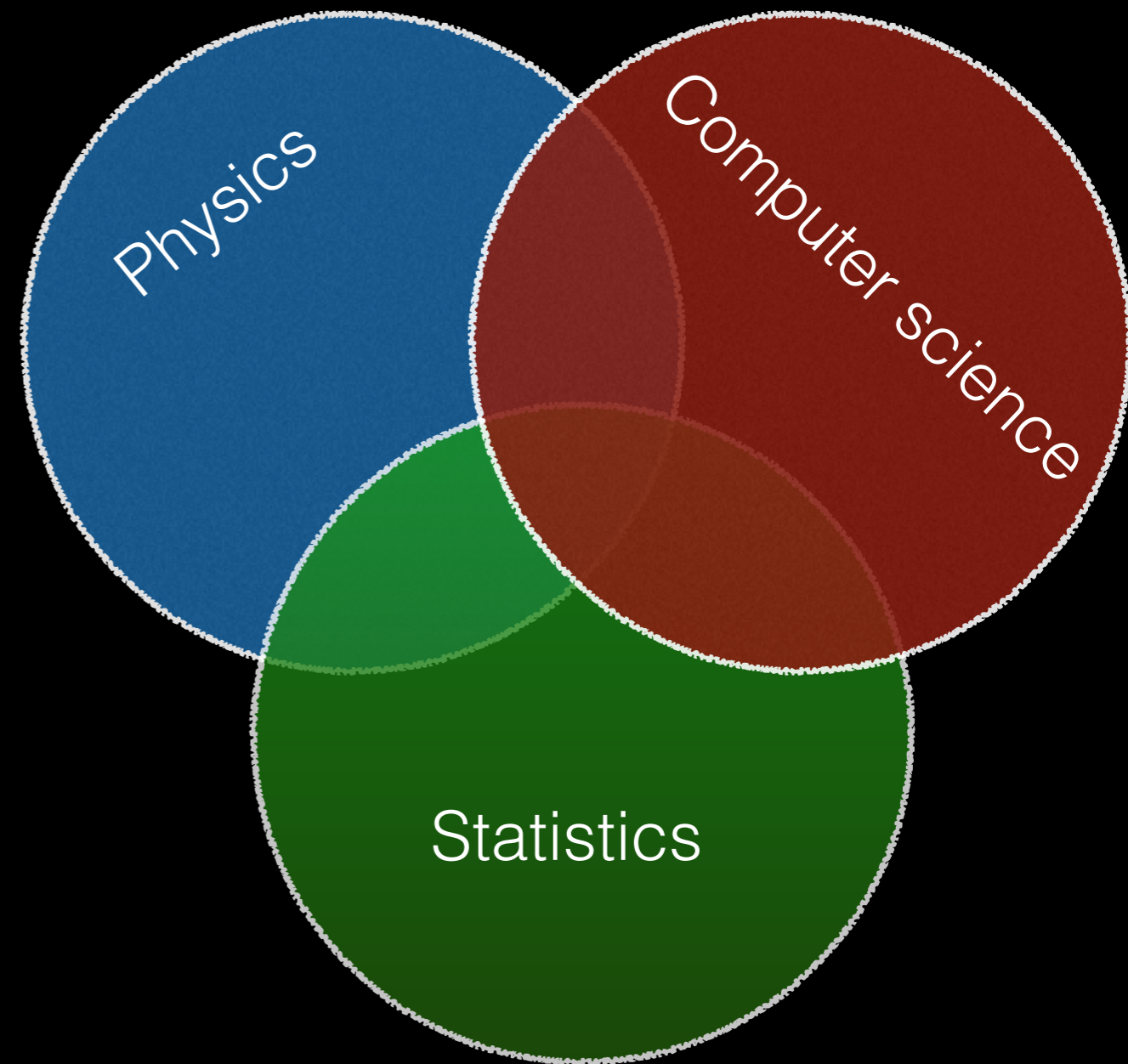
Current challenges

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Current challenges

- Fast sampling in high dimensions
- Diagnostics & Convergence criteria
- Handling of systematic effects



CosmoHammer References

- Akeret J., Seehars S., Amara A., Refregier A., and Csillaghy A. (2013). Astronomy and Computing, Volume 2, Pages 27-39
 - <http://dx.doi.org/10.1016/j.ascom.2013.06.003>
- Available at <http://www.astro.ethz.ch/refregier/research/Software/cosmohammer>