



**Latest results from HERA**

**Phil Bull**  
Manchester/  
UWC

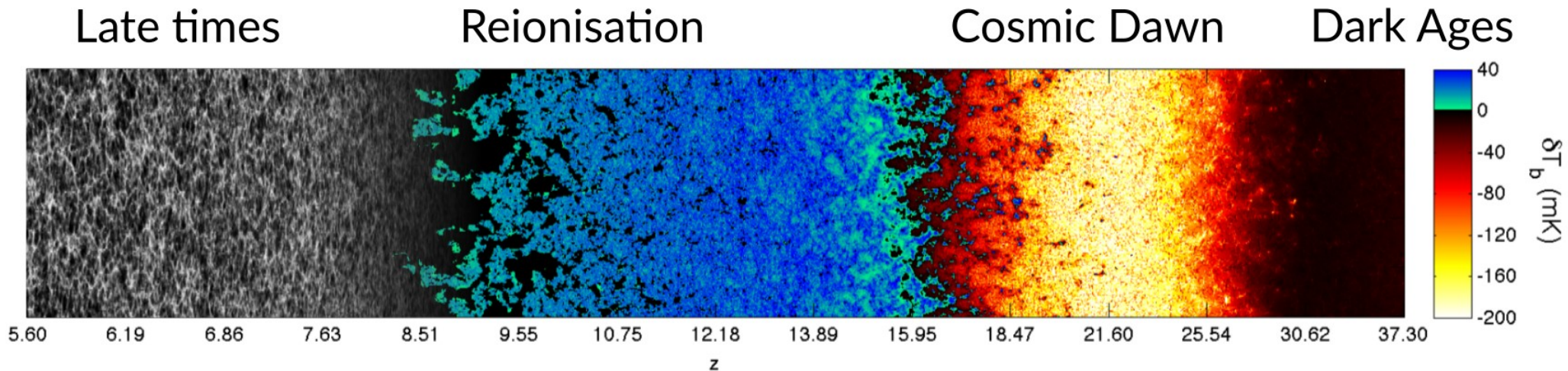
# Outline

- 21cm cosmology at high  $z$
- Overview of the HERA survey
- HERA upper limits



# Cosmology with the 21cm EoR power spectrum

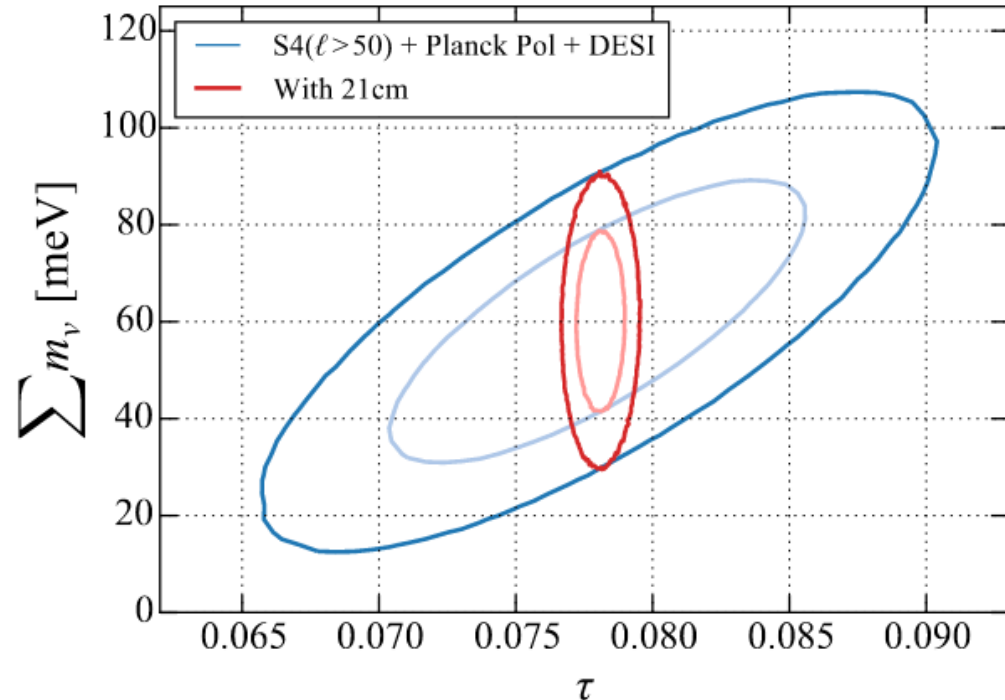
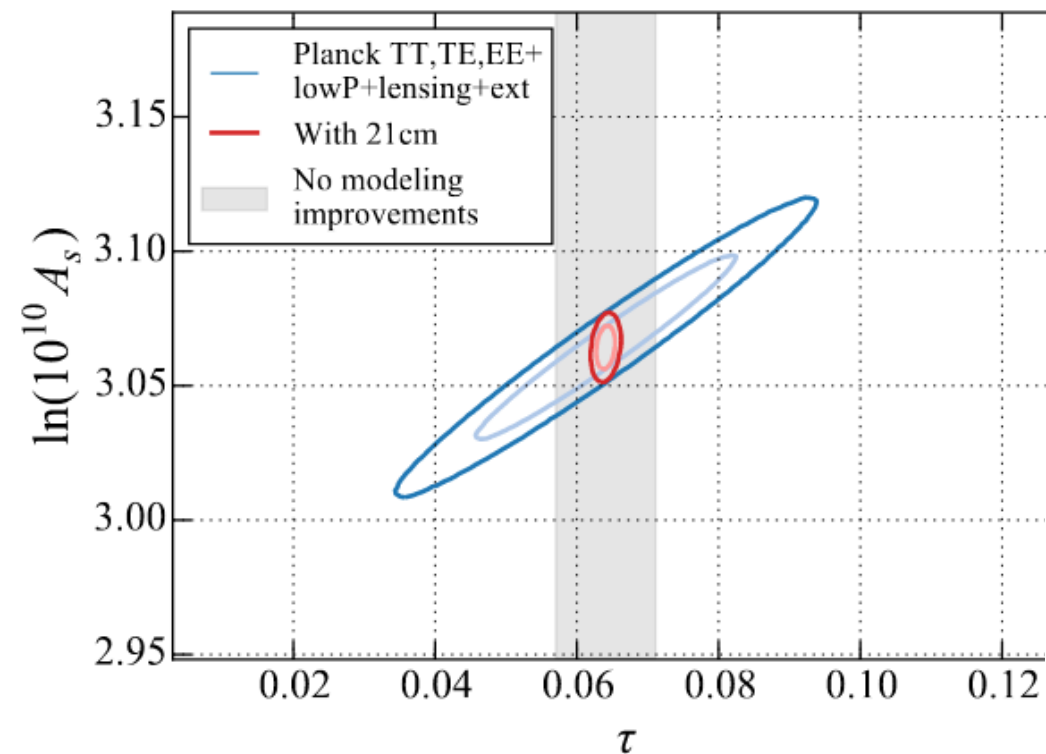
- EoR contributes to optical depth to the CMB,  $\tau$ , which controls TT power spectrum amplitude, low- $\ell$  BB spectrum etc.
- Properties of early galaxies and galaxy formation/evolution models
- Sensitivity to energy injection (e.g. DM annihilation)
- Gateway to the Dark Ages!



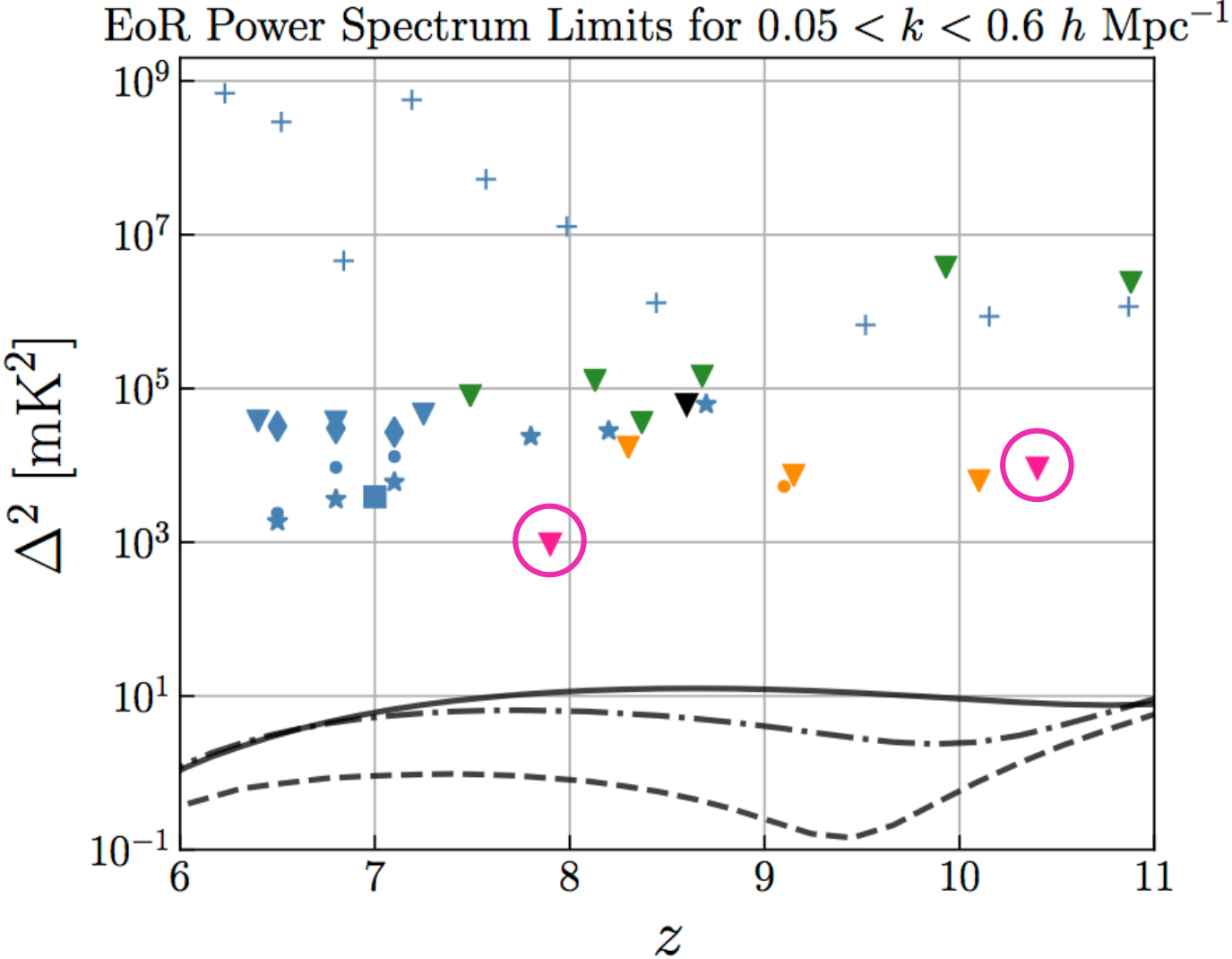
A. Mesinger (adapted)

# Impact of 21cm constraints on parameters

Optical depth is probably the biggest deal in the short-term



# Previous upper limits on the EoR power spectrum

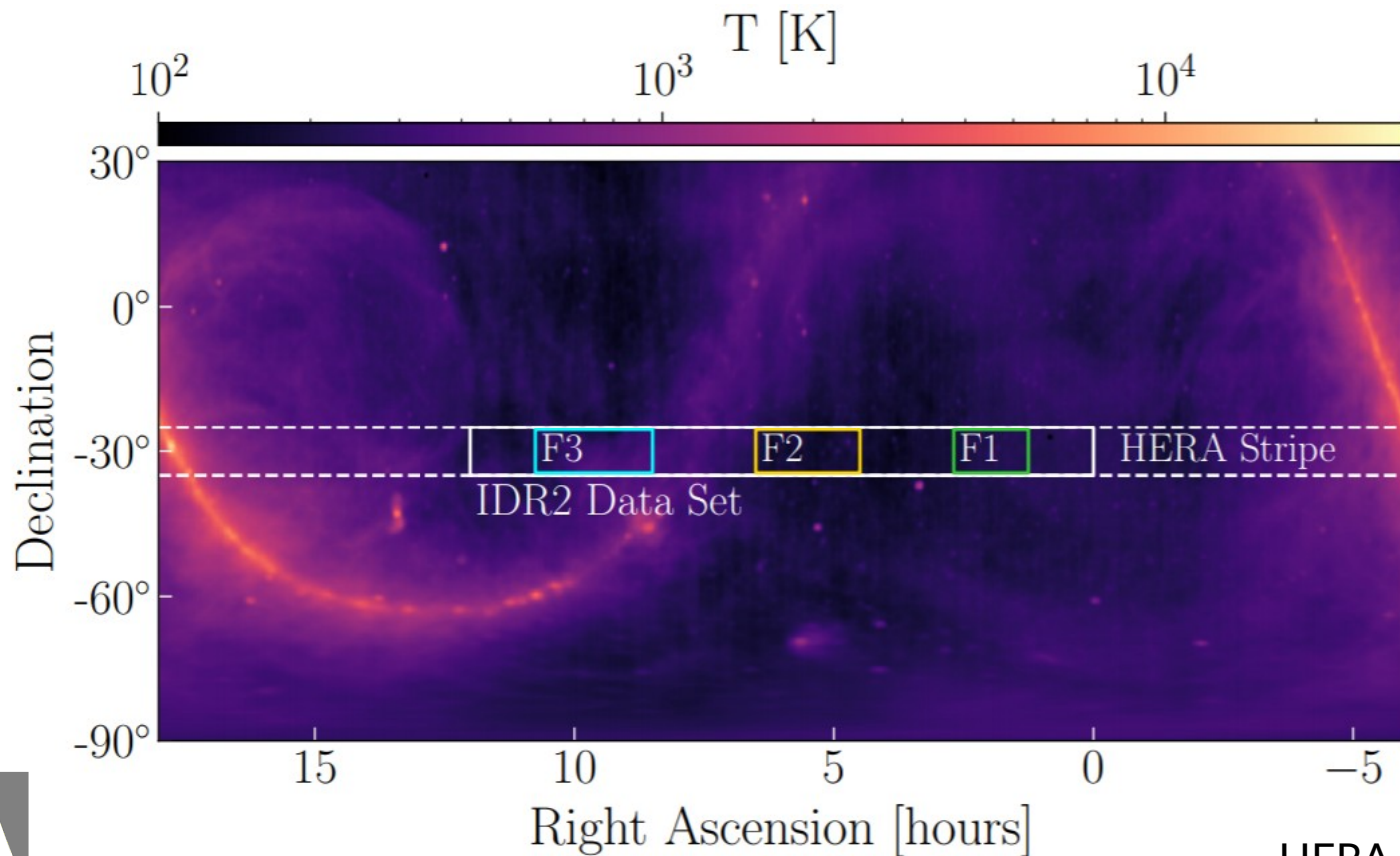


HERA Collaboration (2022)

# Hydrogen Epoch of Reionization Array

Built on SKA site in South Africa, run from UC Berkeley/NRAO

- Eventually: Approx. 350 dishes with dual-pol. receivers (50 – 220 MHz)
- 14m dishes, 14.6m minimum baseline length in hexagonal layout
- Drift scan telescope – it just points up!
- Redundant array – many copies of the same baseline





MeerKAT

HERA

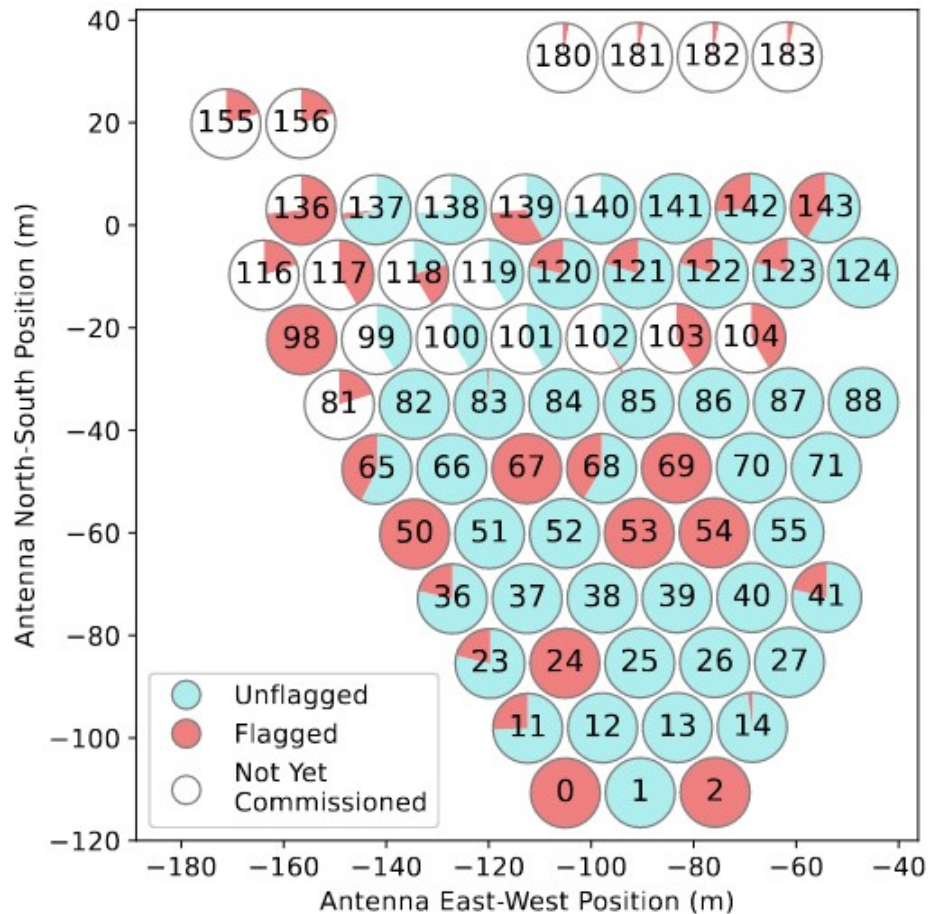




# HERA Phase I array

Data from between 35-41 good antennas (out of 47-71)

- 94 nights of good data (out of 182) between 29<sup>th</sup> Sep. 2017 - 31<sup>st</sup> Mar. 2018
- Using the old PAPER receivers (100 – 200 MHz, 1024 channels)
- Focused on 5 “clean” fields, 2 sub-bands

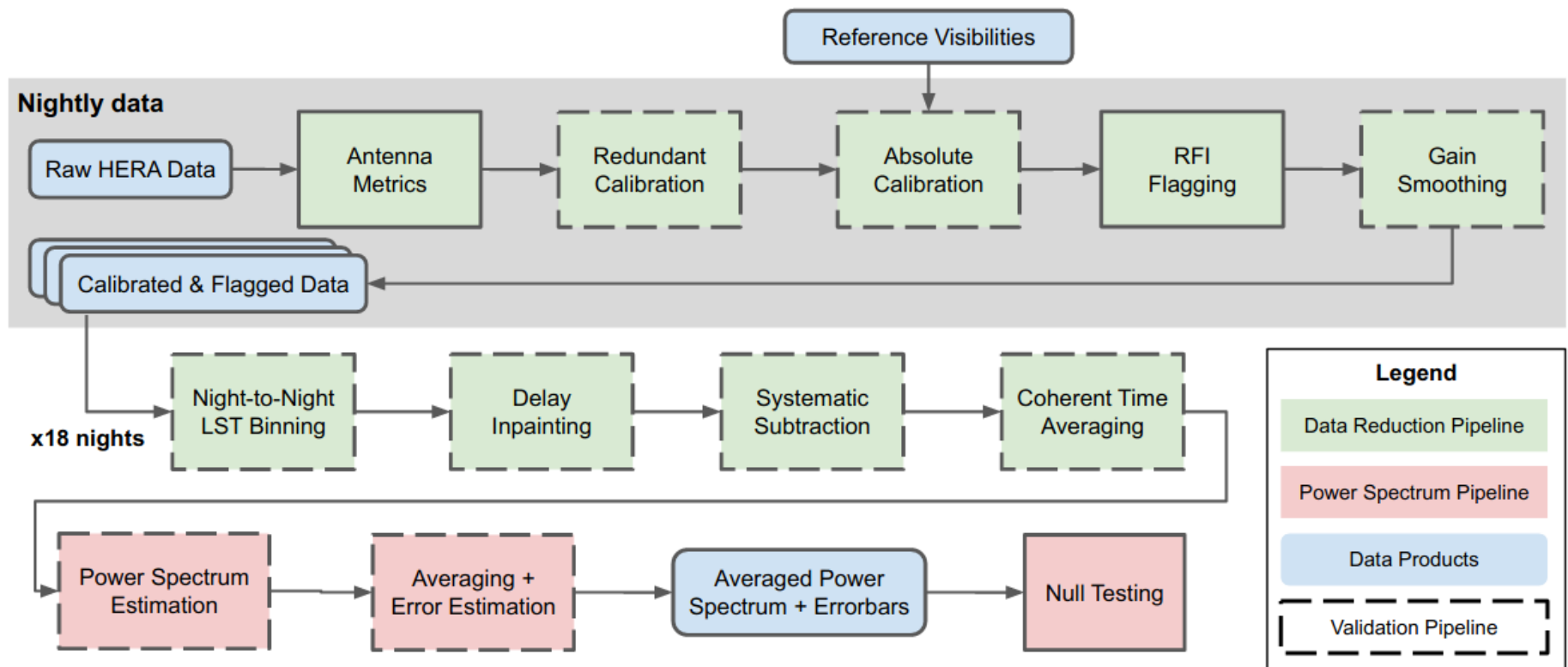


- Full redundant calibration
- Partial absolute calibration
- Smoothing of gains
- Reflection and cross-talk subtraction
- Foreground avoidance

# Analysis pipeline

We use *redundant calibration*, a restricted *absolute calibration*, and a *delay power spectrum* analysis.

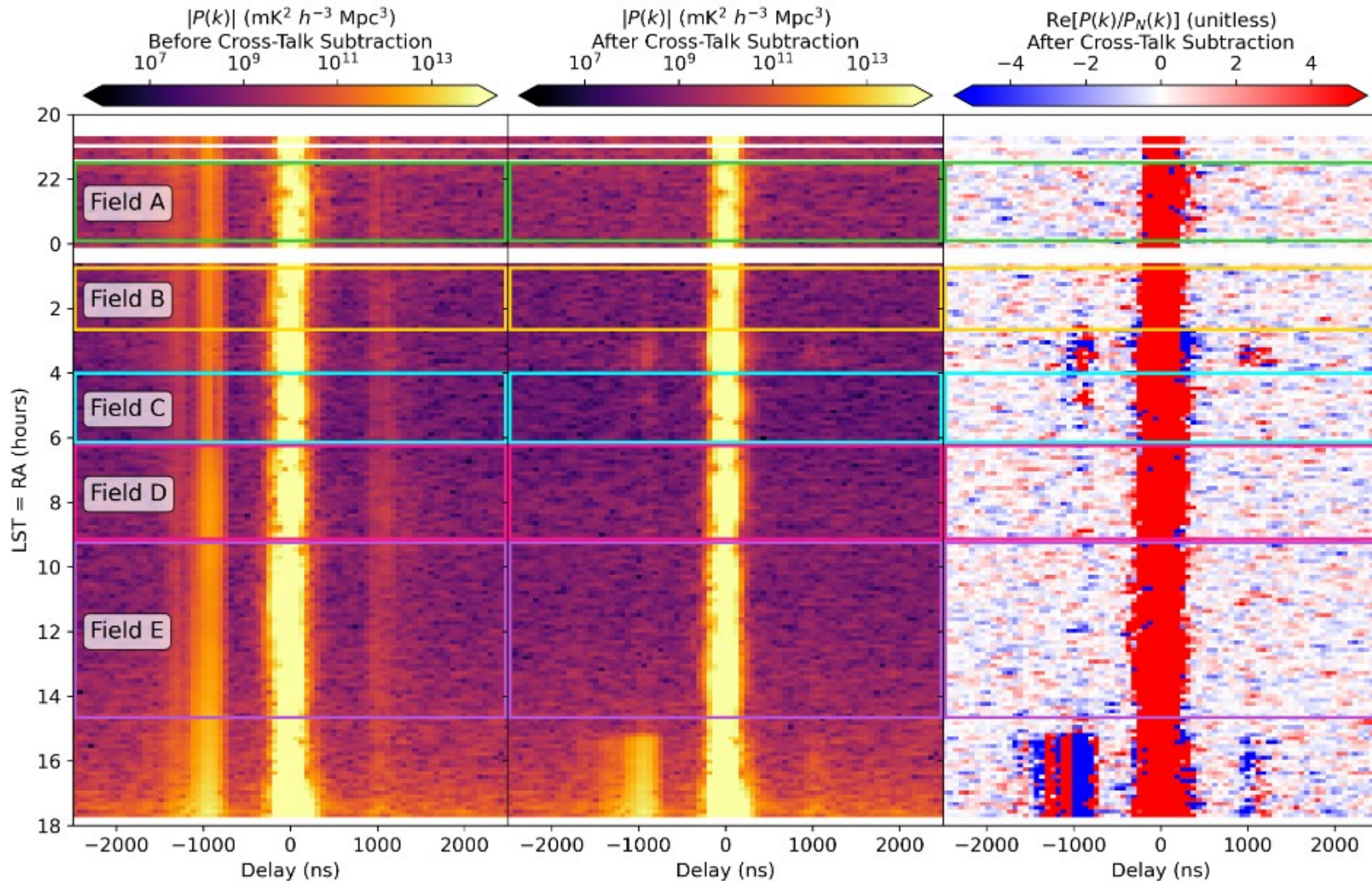
- Absolute calibration (using a sky model) only operates on a degenerate subspace of the gain solutions, which are also smoothed.
- We don't do any imaging! Power spectra are per-baseline.



# Reflection and cross-talk systematics

Significant contamination detected outside the foreground wedge

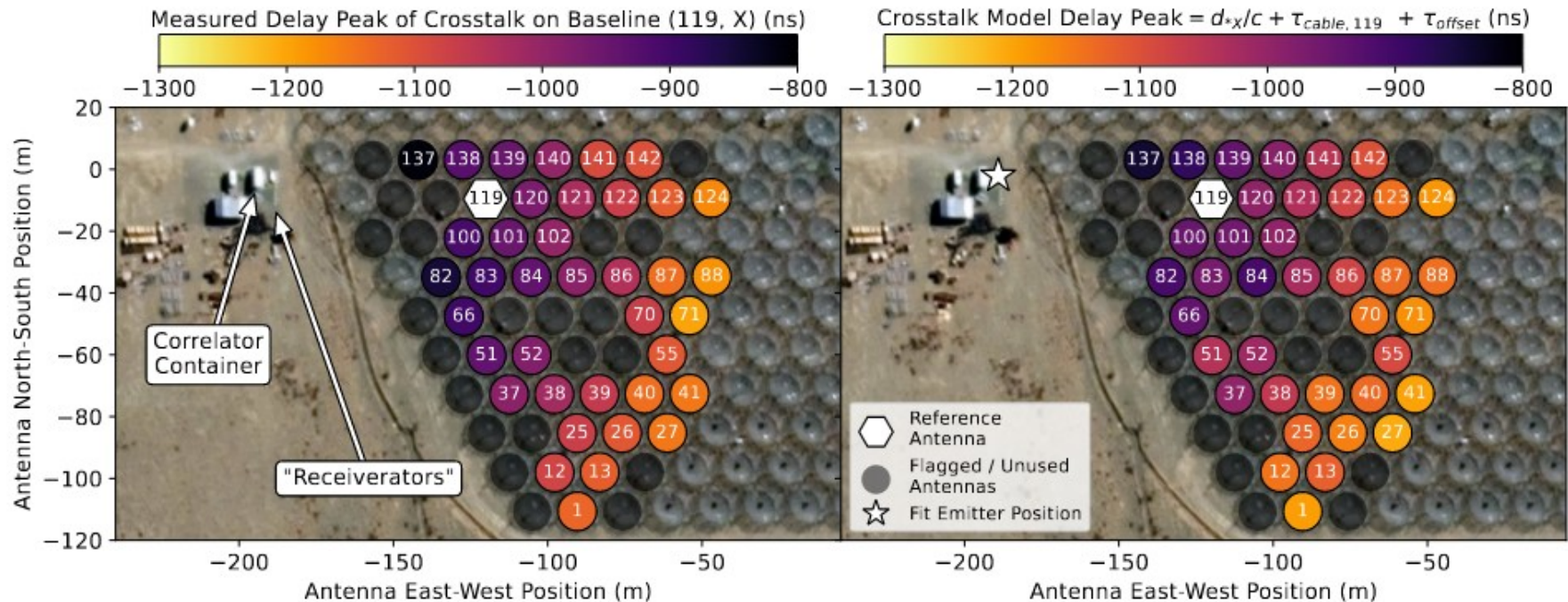
- Mostly due to cable reflections and over-the-air crosstalk (Kern et al. 2020)
- Cross-talk can be fitted out using 35-term Fourier-space model



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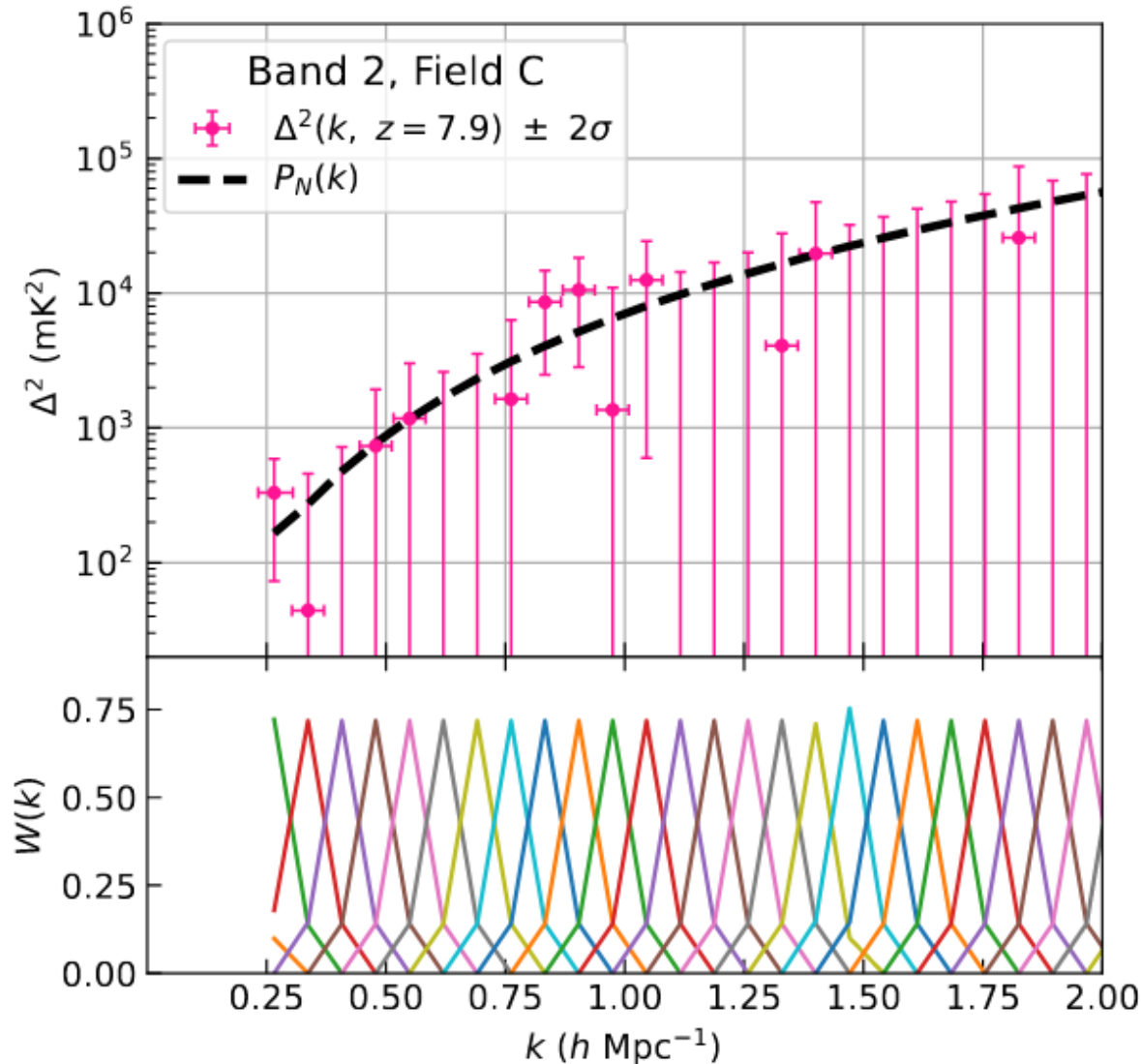
- Mostly due to cable reflections and over-the-air crosstalk (Kern et al. 2020)
- Cross-talk can be fitted out using 35-term Fourier-space model
  - Stable for particular array layout, but changes as the array is extended
  - This allowed us to identify the source!



Antenna → 150m cable → amplifiers in receiverator → broadcast back to antennas

# Power spectrum upper limits

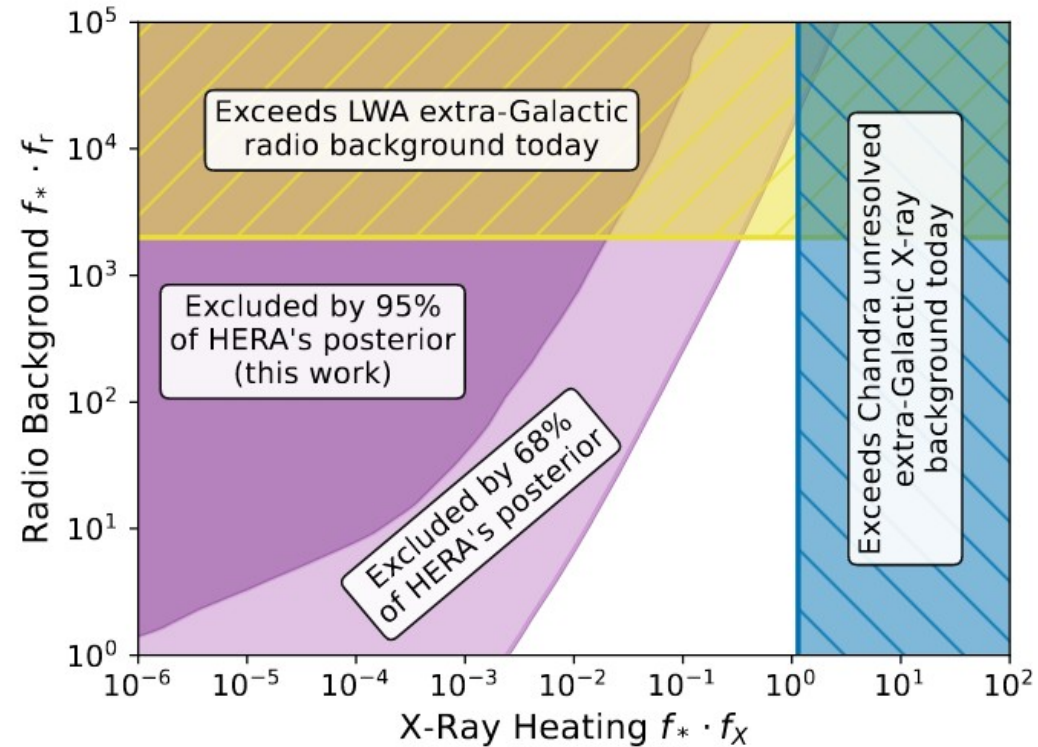
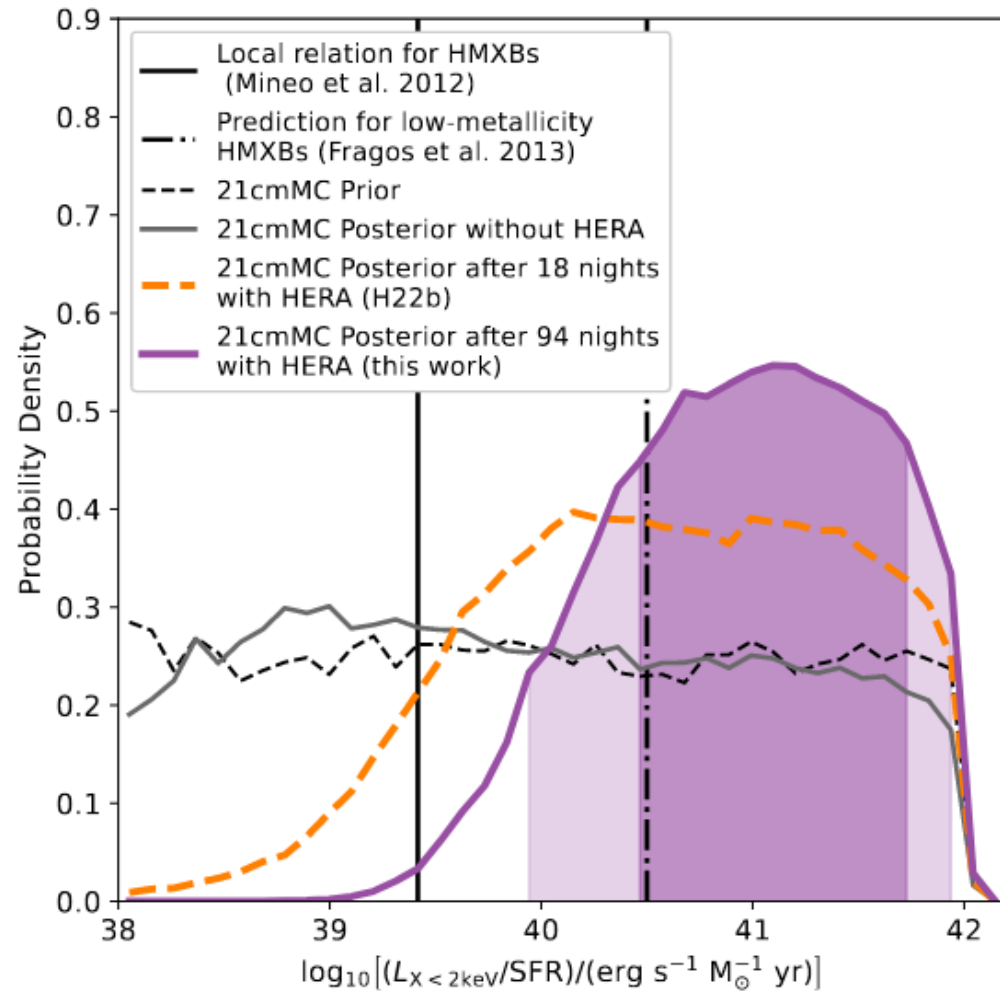
HERA Collab. [2210.04912]



- Robust power spectrum upper limits
- Mostly consistent with being noise-dominated (confirmed by statistical tests; see Mike's talk)
- Conservative analysis with low signal loss (6-7%; confirmed by validation sims etc.)

# Physical implications

HERA Collab. [2210.04912]



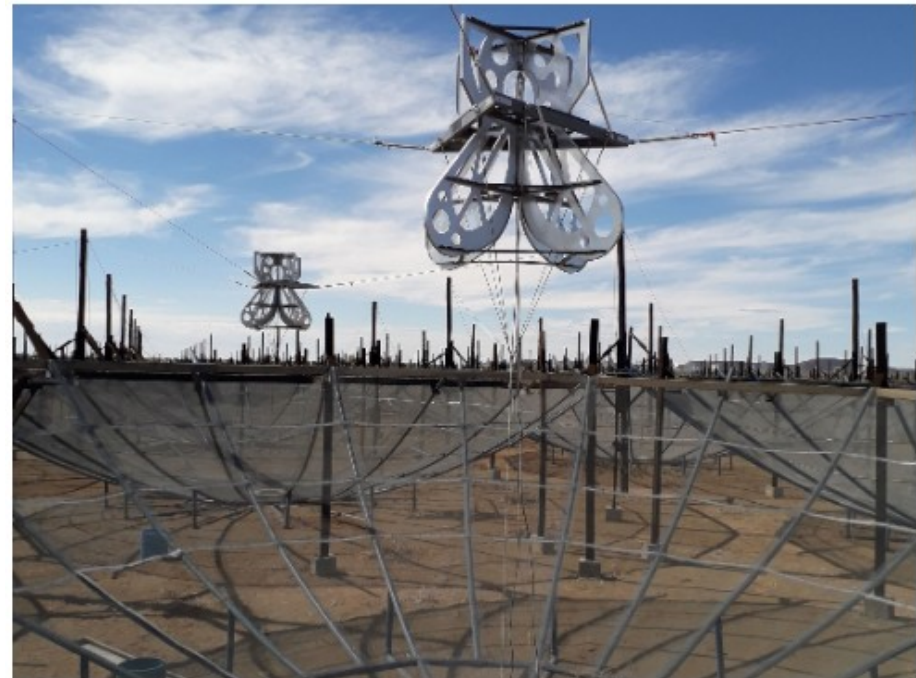
- Implies early galaxies had enhanced X-ray luminosity (for a given SFR)
- Disfavours unheated IGM at >99% level → heating definitely took place!
- Improved constraints on size of excess radio background

# What's next?

Fagnoni et al. [2009.07939]

## Phase II data

- Brand new Vivaldi feeds (50–220 MHz)
- Multiple new seasons of data with increasing number of dishes
- Fringe-rate filtering (see Hugh's talk)
- Lots of incremental improvements to analysis + new methods
- (Ask me about Hydra, our new Bayesian pipeline!)



# The HERA pipeline

Almost all of our analysis pipeline is open source

- Redundant and partial absolute calibration: **hera\_cal**
- Power spectrum estimation: **hera\_pspec**
- Visibility simulators (with systematics etc): **hera\_sim**
- Lots of other smaller utilities, RFI flagging etc.
- Very good unit test coverage, validation etc.

See <https://github.com/HERA-Team>

