Event reweighting and generative models

In neutrino experiments

IPA workshop on Machine Learning for particle physics and astrophysics March 21, 2023



LABORATÓRIO DE INSTRUMENTAÇÃO E FÍSICA EXPERIMENTAL DE PARTÍCULAS partículas e tecnologia

Cristóvão Vilela

Reweighting and generating data

- Machine learning often used to **summarise** multidimensional data.
 - Given high-dimensional representation of the event *X*, produce a classification score *y*.

Features,
$$X \longrightarrow ML \longrightarrow$$
 Labels, y

• But we may be interested in **generating** some data *X*, starting from the summary *y*, or some modified data *X'*, starting from our nominal data *X*.

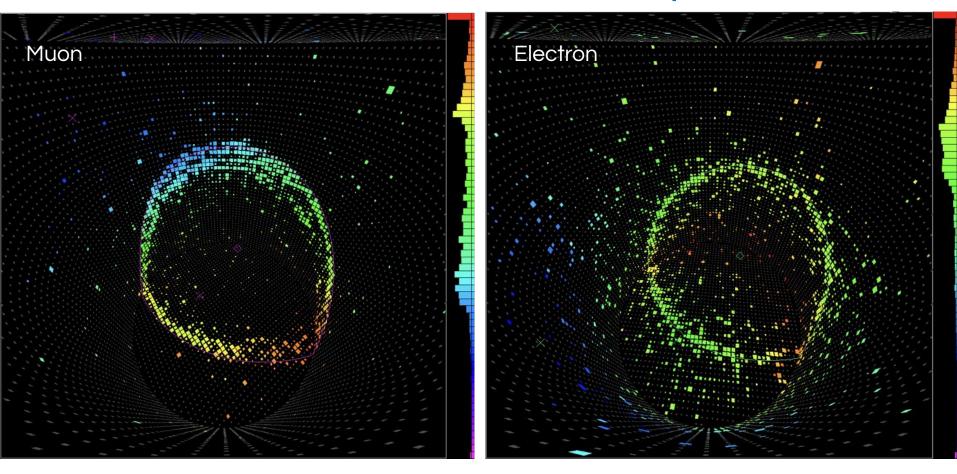
"Labels",
$$y \longrightarrow ML \longrightarrow$$
 "Features", $X \longrightarrow ML \longrightarrow$ Features, $X \longrightarrow ML \longrightarrow$ Features, X'

- I will briefly describe three applications of ML in neutrino physics where we are interested in obtaining a detailed description of events, *X*, rather than the summary, *y*.
 - A generative model for Super-Kamiokande events.
 - Reweighting between neutrino interaction models in DUNE.
 - This is actually a classification task.
 - Translating DUNE events from the near to the far detector.

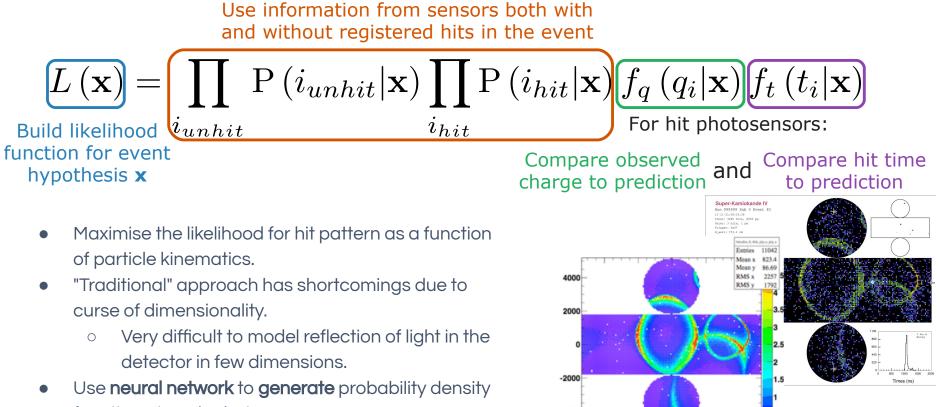
Super-Kamiokande event reconstruction

Super-Kamiokande 50 kt H₂O 11k photomultiplier tubes

Particle identification at Super-K



Super-K event reconstruction

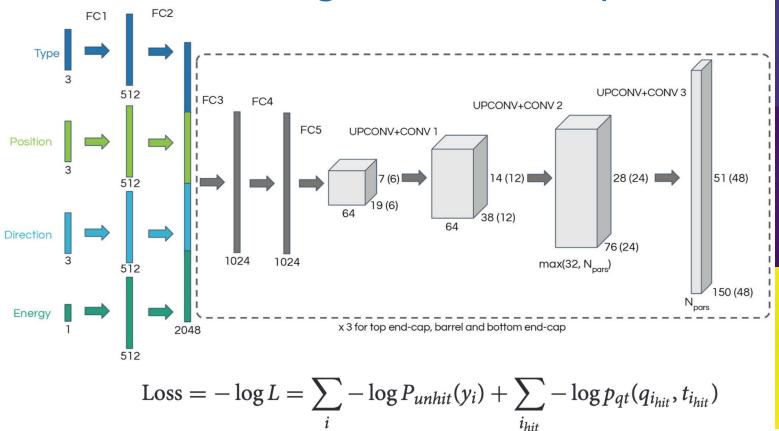


function at each photosensor.

2000

4000

Generating PDFs for Super-K

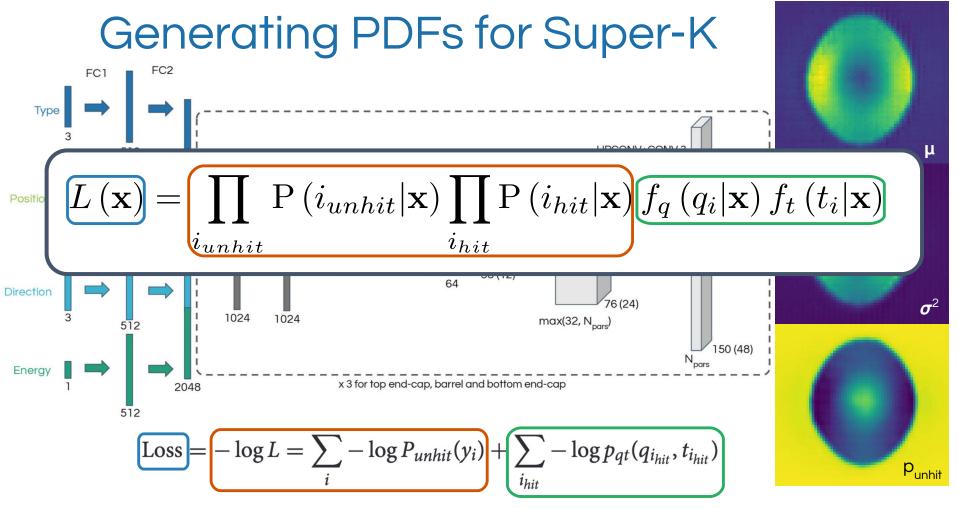


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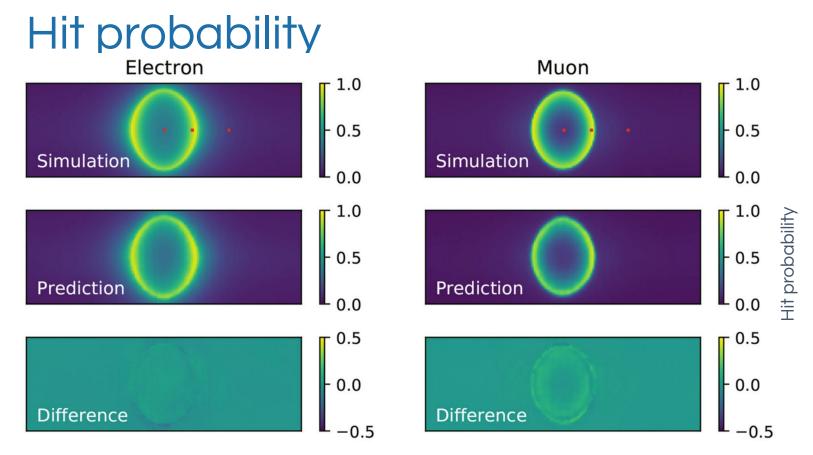
μ

 σ^2

P_{unhit}



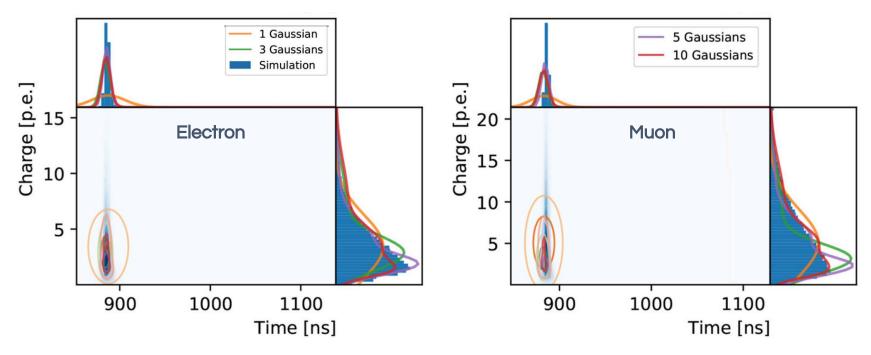
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• Compare neural network output to average of 50k events generated with the same parameters.

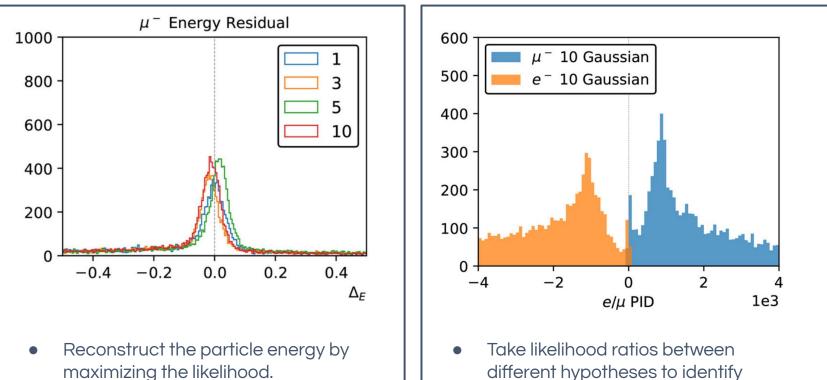
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Charge and time PDF



PMT at the edge of the Cherenkov ring.

Energy reconstruction

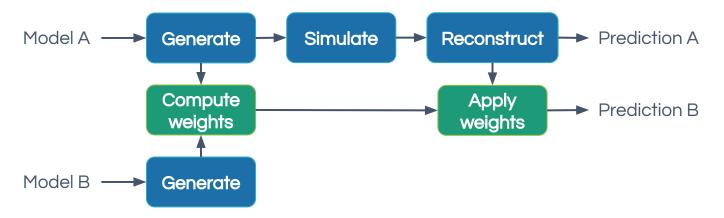


different hypotheses to identify particles.

DUNE event reweighting

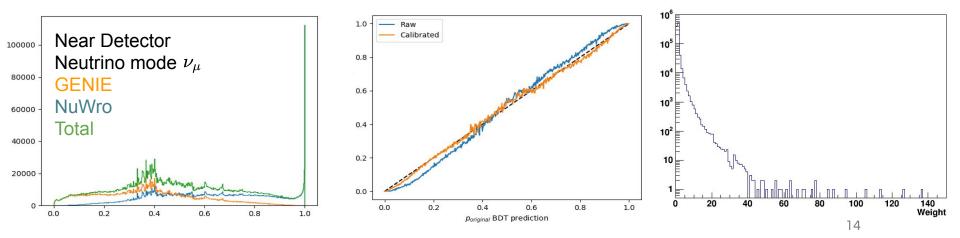
Event reweighting: what is it good for?

- We use Monte Carlo methods to predict what our data should look like under different hypotheses.
- Roughly factorizes into:
 - a. Generate neutrino-nucleus interactions (fast)
 - b. Simulate the detector response (slow)
 - c. Reconstruct the events (slow)
- There are significant uncertainties in the neutrino-nucleus interaction models, so we want to test the impact of different models on our sensitivity.
 - However, re-running the full simulation chain is often prohibitively expensive.

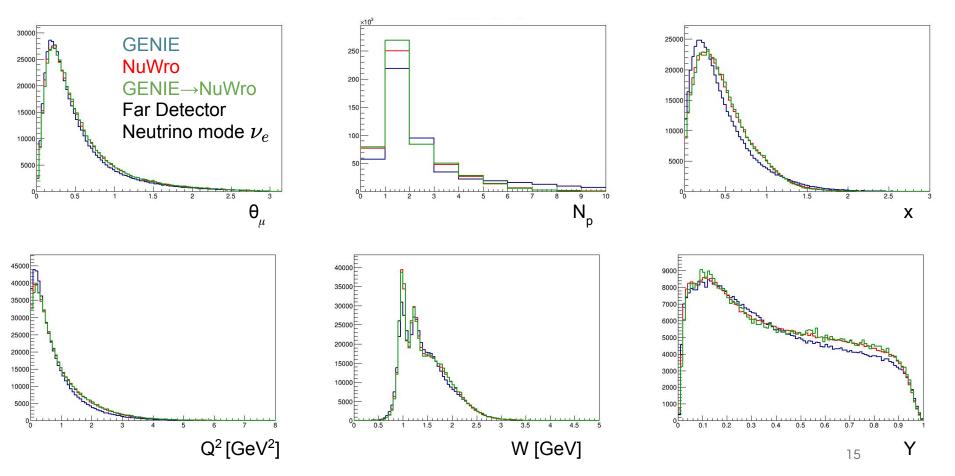


Reweighting between models

- Traditional reweighting methods make use of histograms.
 - Allows for reweighting up to 2 or 3 dimensions but not more.
 - Can only be used to reweight low-dimensional parts of the model that can be factorized from the rest of the model.
- But we can **reweight in high number of dimensions** using ML.
- DUNE TDR: **boosted decision trees** used to **classify** generated events between two different models using 18 variables for describing the events.
 - Classifier output can be interpreted as a **probability** and expressed as a **weight**.

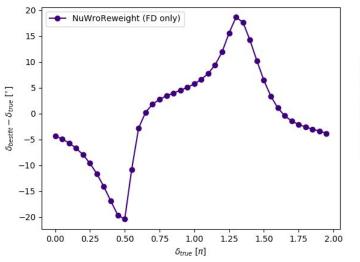


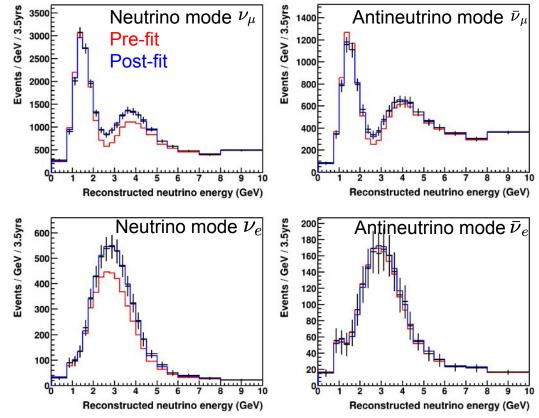
Reweighting examples



Check model impact on analysis

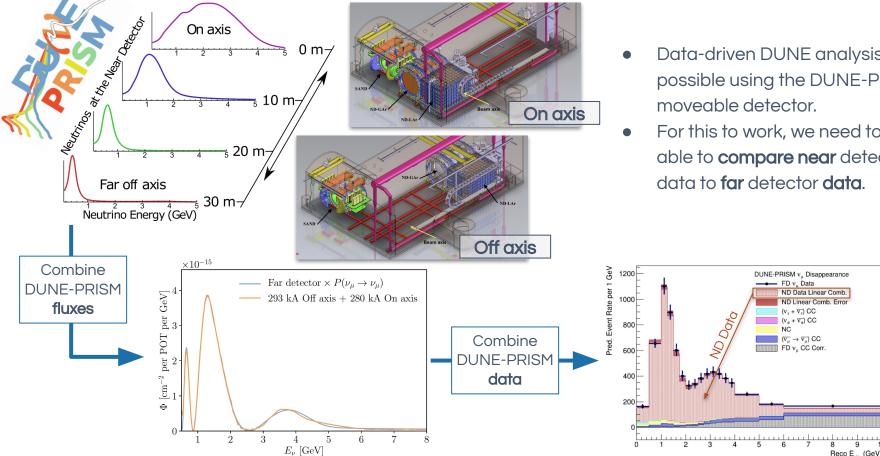
- Run neutrino oscillation analysis on reweighted simulation.
- Get significantly different results if the wrong model is assumed.
 - But near detector saves the day!
- Important result, only possible using ML.





DUNE near to far detector translation

Precision Reaction-Independent Spectrum Measurement



- Data-driven DUNE analysis possible using the DUNE-PRISM moveable detector.
- For this to work, we need to be able to compare near detector data to **far** detector **data**.

DUNE-PRISM v., Disappearance

 $(v_r + \overline{v_r}) CC$ $(v_{o} + \overline{v_{o}}) CC$

 $(\overline{v_u} \rightarrow \overline{v_u}) CC$ FD v., CC Corr.

ND Data Linear Comb.

ND Linear Comb. Error

- FD v. Data

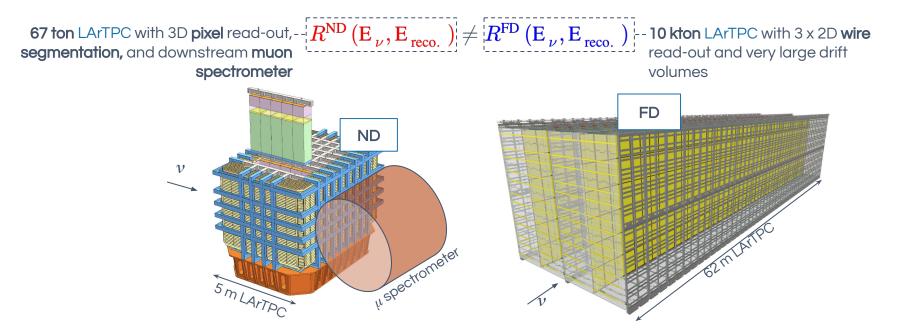
NC



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Reco Euro (GeV)

Detector response in data-driven analyses



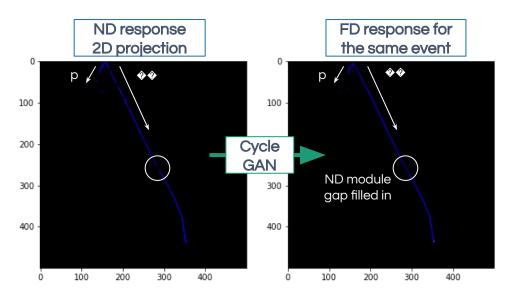
- Need a model-independent method to account for differences in the detector responses.
 - If an ND event had occurred instead in the FD, what would be its reconstructed energy at the FD?

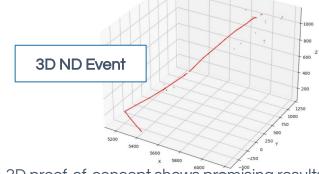
Learning the differences between ND and FD



Monet → photo CycleGAN arXiv:1703.10593 (2017)

- Traditional approach: response matrices out of high-level reconstructed variables.
 Leads to model dependence.
 Model
 Curse of
 dependence
- Reduce model dependence by using **image-to-image translation** techniques to generate FD-like events from ND events at **hit level**.





- 2D proof-of-concept shows promising results.
- Next step:
 - \circ ND 3D readout \rightarrow FD 3x2D readout
 - Needs novel neural network architectures.
 - Work in progress with R. Radev (CERN).

Summary

- Lots of interesting applications of machine learning in neutrino physics.
- Showed three examples of applications not focused on classification:
 - Generative model for Super-Kamiokande events.
 - Produces probability density functions for the detector sensors given an event hypothesis.
 - Event reweighting for DUNE.
 - Reweight between interaction models in many dimensions.
 - Enables important studies that would not be possible without ML.
 - Near-to-far detector event translation in DUNE.
 - Get far detector response given a near detector event.
 - Exploit full information content of the data to avoid model dependence.
 - Work in progres...

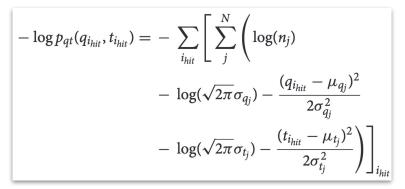


PDF parameterisation

- Requirements:
 - Statistically robust (i.e., integrates to 1) and smooth (for gradient descent)
 - General-ish do not assume what the PDFs look like a priori

• Strategy:

- Use combinations of Gaussians (a la Gaussian Mixture Model)
 - 2x1D or 2D, correlated or uncorrelated.



Uncorrelated

Note: q-t correlations can still arise from the different means of each component.

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$$f(\boldsymbol{\eta}|\boldsymbol{\theta}) = \sum_{j}^{N} \frac{n_{j}}{(2\pi)|\Sigma_{j}|^{1/2}} \exp(-\frac{1}{2}(\boldsymbol{\eta}-\boldsymbol{\theta}_{j})^{\mathrm{T}}\Sigma_{j}^{-1}(\boldsymbol{\eta}-\boldsymbol{\theta}_{j}))$$

Correlated Extra DOF for each component: the q-t correlation.

$$\Sigma^{-1} = \begin{pmatrix} \alpha_{11} & 0 \\ \alpha_{12} & \alpha_{22} \end{pmatrix} \begin{pmatrix} \alpha_{11} & \alpha_{12} \\ 0 & \alpha_{22} \end{pmatrix}$$

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Work with components of the Cholesky-decomposed triangular matrix.

Add conditions to the alphas to guarantee covariance matrix properties.