

Testing the EP with the QE

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based on work in progress

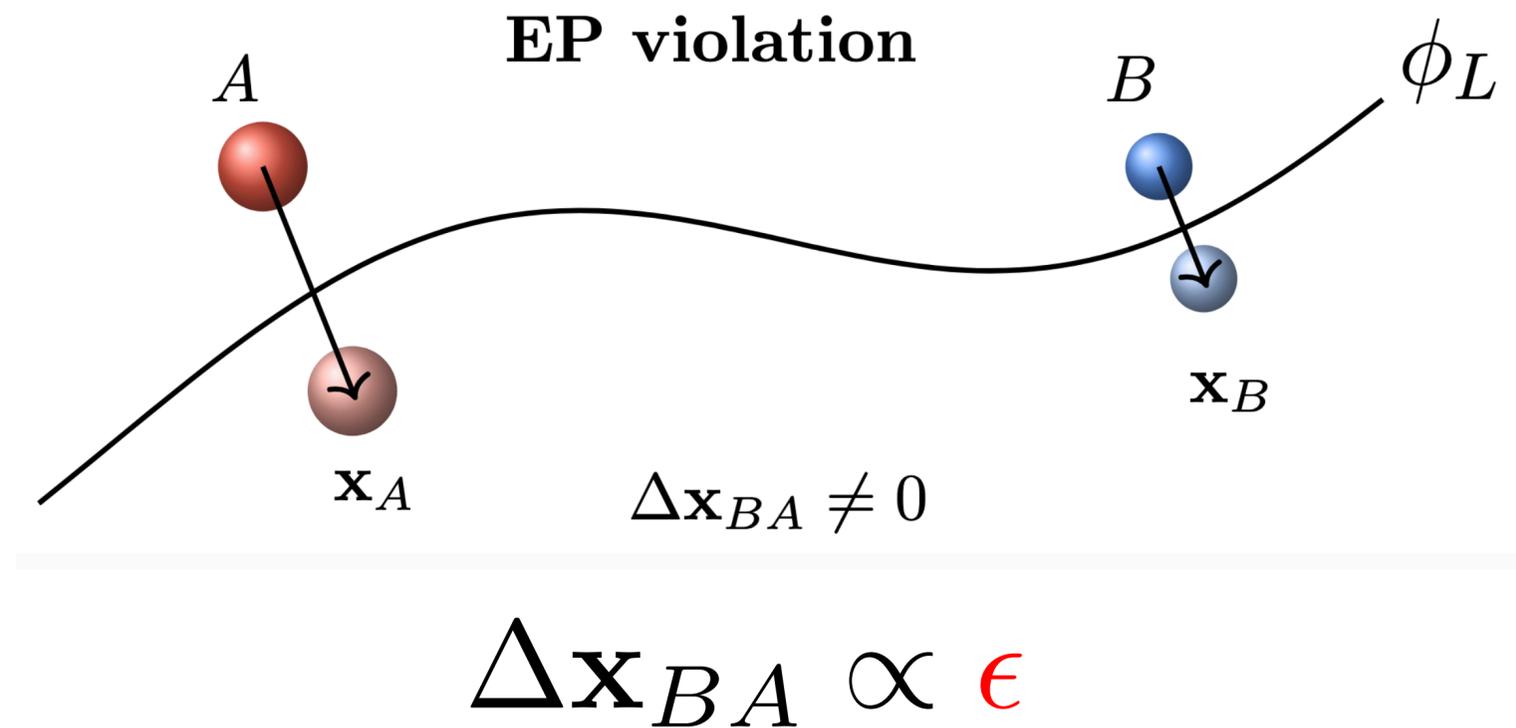


Swiss Cosmology Days 2025

Can QE test universal free fall?



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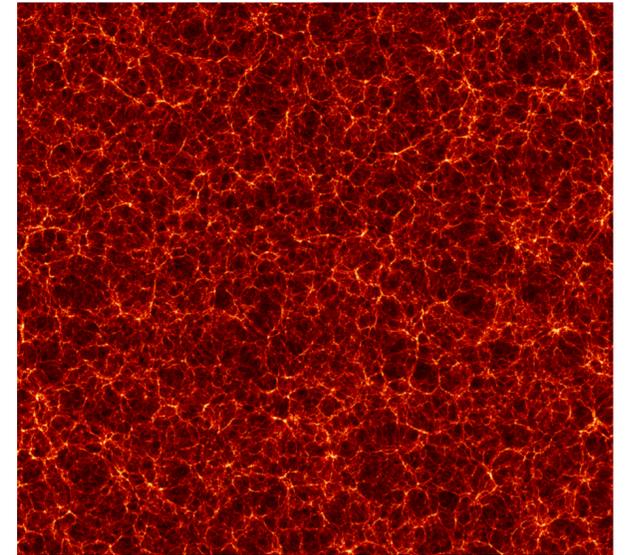


$$\langle \delta_A \delta_B \rangle_L \supset \epsilon \delta_L$$

Falling Dark Matter

Dark-matter non-linear due to gravity

$$\delta \sim \delta_1 + \boxed{\delta_2} + \dots$$



Credit: AbacusSummit

$$\delta_1^2$$

Non-linear growth (G)

$l=0$

$$\vec{\Psi} \cdot \vec{\nabla} \delta_1$$

Shift (S)

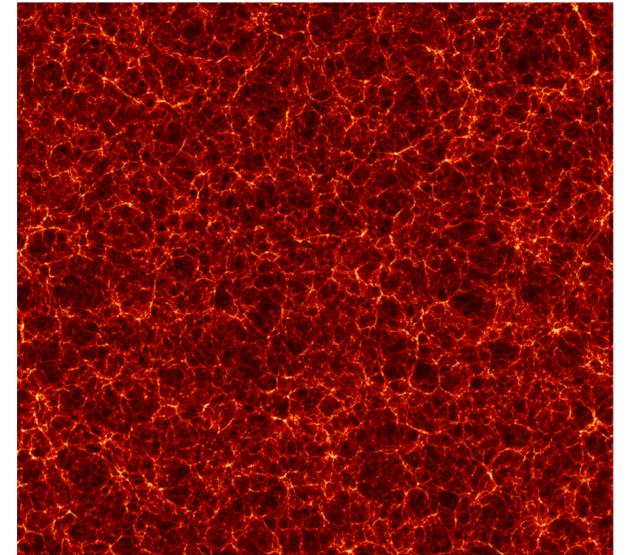
$l=1$

$$\left(\left[\frac{\partial_i \partial_j}{\partial^2} - \frac{1}{3} \right] \delta_1 \right)^2$$

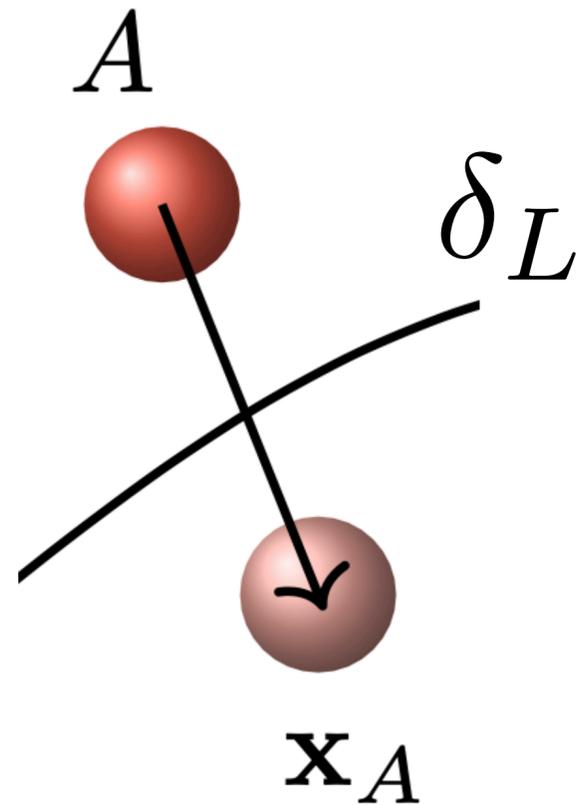
Tidal (T)

$l=2$

Falling Dark Matter



Credit: AbacusSummit



$$\delta \sim \delta_1 + \delta_2 + \dots$$



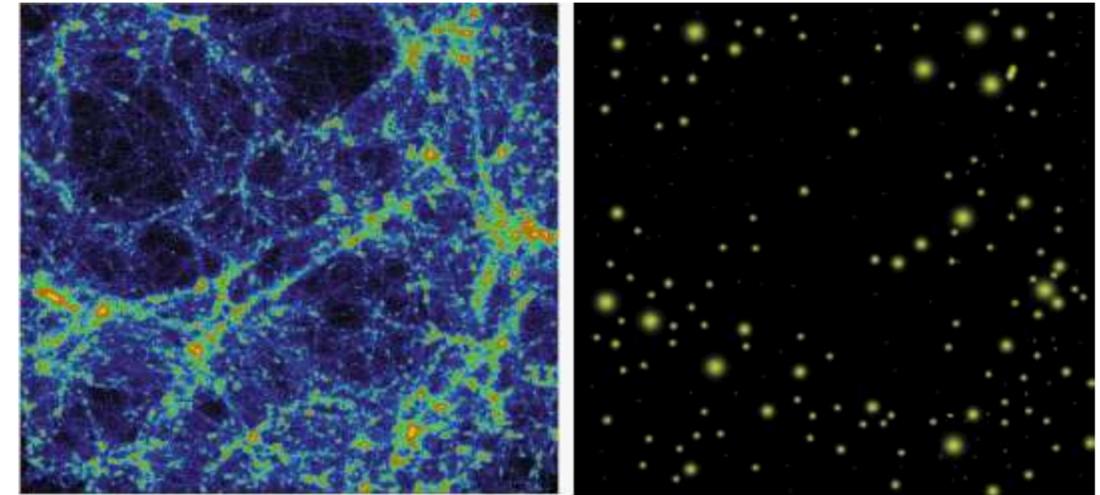
$$\vec{\Psi} \cdot \vec{\nabla} \delta_1$$

$$\delta_1(\vec{x} + \vec{\Psi}_L)$$

$$T_{\text{cmb}}(\hat{n} + \vec{d}_L)$$

Falling Galaxies

Galaxies are biased tracers of non-linear dark matter



Credit: Cooray & Sheth

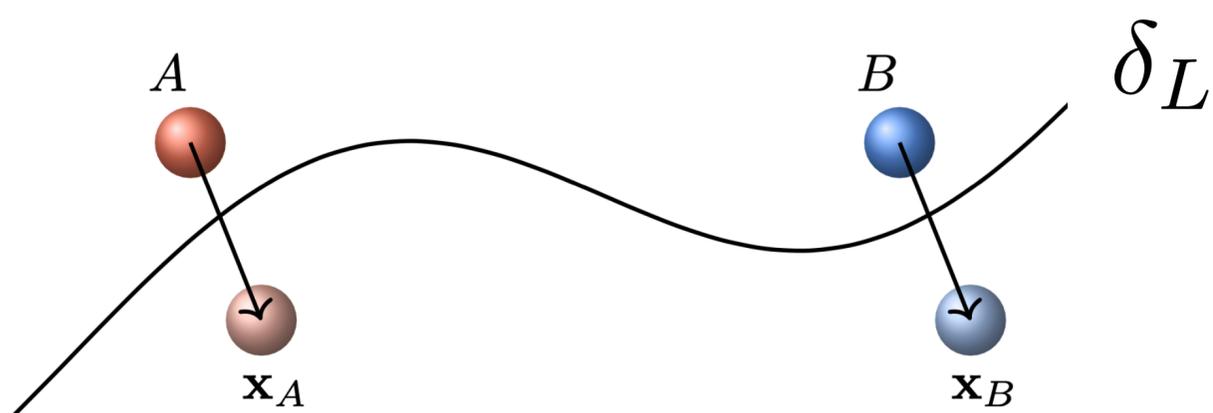
$$\delta_g \sim b_1 \delta_1 + (b_1 + 21/17b_2) \delta_G + b_1 \delta_S + (b_1 + 7/2b_{s^2}) \delta_T$$



No new additional bias!

$$b_1 \delta_1 (\vec{x} + \vec{\Psi}_L)$$

Shifting under a long-mode



$$\delta_L \quad B_{ABL} \sim \langle \delta_L(\vec{K}) \delta_A(\vec{k}_A) \delta_B(\vec{k}_B) \rangle$$

$$f_{AB} \delta_L = \langle \delta_A(\vec{k}_1) \delta_B(\vec{k}_2) \rangle_L \quad f = \frac{B(\vec{K}, \vec{k})}{P_L}$$

e.g. Lewis+2011, Lewis 2011, Barreira & Schmidt 2017, Carron&Lewis 2024

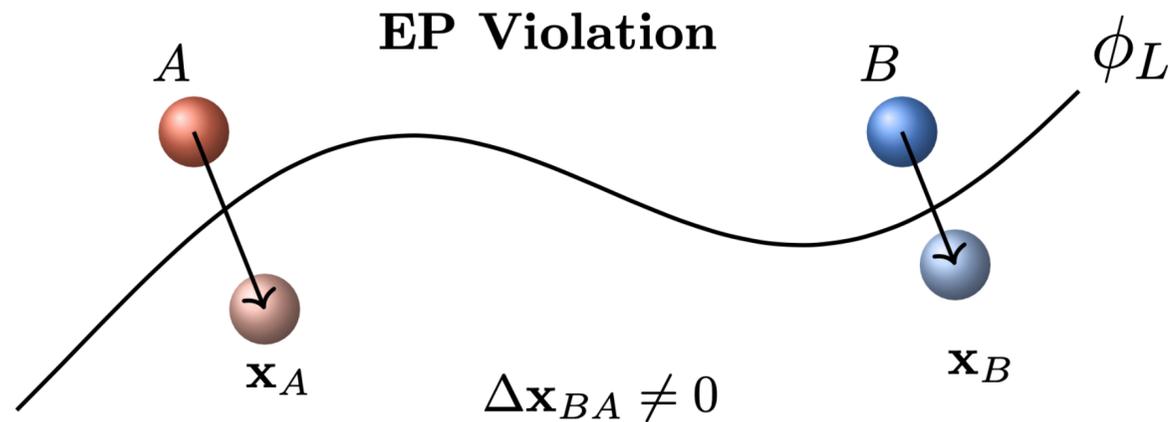
Symmetric

Anti-symmetric

$$\begin{aligned} f_{AB}(\mathbf{k}_1, \mathbf{k}_2) &= C_{(AB)}^G f_G^{(+)}(\mathbf{k}_1, \mathbf{k}_2) + C_{[AB]}^G f_G^{(-)}(\mathbf{k}_1, \mathbf{k}_2) \\ &+ C_{(AB)}^S f_S^{(+)}(\mathbf{k}_1, \mathbf{k}_2) + \boxed{0} \\ &+ C_{(AB)}^T f_T^{(+)}(\mathbf{k}_1, \mathbf{k}_2) + C_{[AB]}^T f_T^{(-)}(\mathbf{k}_1, \mathbf{k}_2), \end{aligned}$$

Preliminary

Now break the EP



$$f_{AB}(\vec{K}, \vec{k}) \supset \boxed{C_{[AB]}^S} f_S^{(-)}(\vec{K}, \vec{k}) \underset{K \rightarrow 0}{\sim} \epsilon \frac{\vec{k} \cdot \vec{K}}{|\vec{K}|^2}$$

$$\epsilon \frac{b_{1B} b_{\epsilon, SA} - b_{1A} b_{\epsilon, SB}}{2} + \frac{b_{1B} b_{1A} - b_{1A} b_{1B}}{2}$$

Kehagias, Riotto, Petroni, Peloso, Creminelli, Simonović, Valageas, Lewandowski, Vernizzi, ...

Motivation to study (mixed) bispectrum to detect violations of EP

• **But covariance is hard**

• **Combination with trispectrum? Shot noise? Window functions? Systematics?**

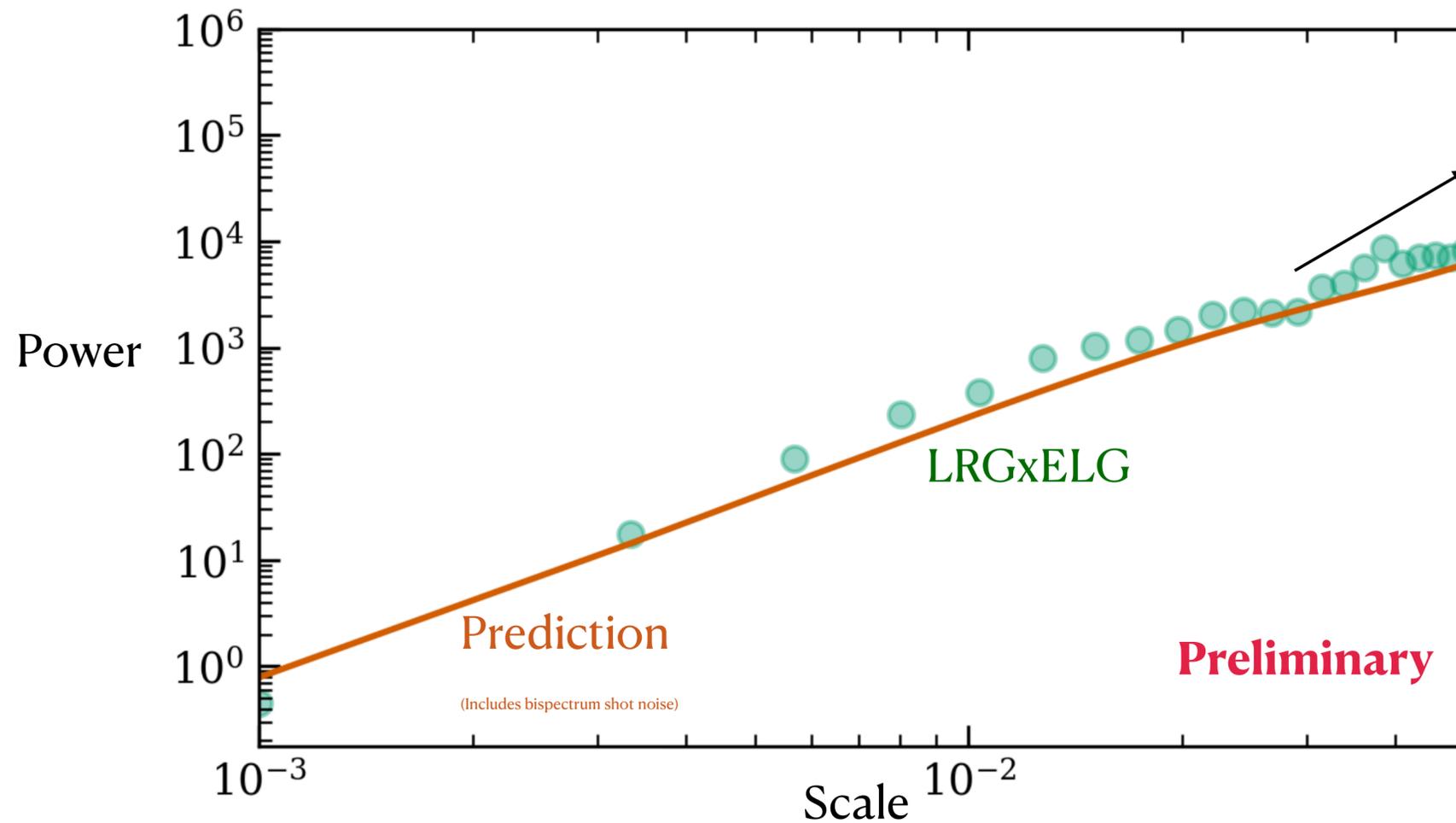
LSS QE

$$\langle \delta_A(\vec{k}_1) \delta_B(\vec{k}_2) \rangle_{\text{fixed } \delta_L} \sim f(\vec{k}_1, \vec{k}_2 = \vec{K} - \vec{k}_1) \delta_L$$

$$\widehat{\delta}_L \sim \frac{\delta_A \delta_B}{f}$$

e.g. Lewis 2011, Schmitfull+ 2015, de Putter 2018, Dai+ 2020, Li+2020, Darwish+ 2020, Carron&Lewis 2024

Abacus DESI sims



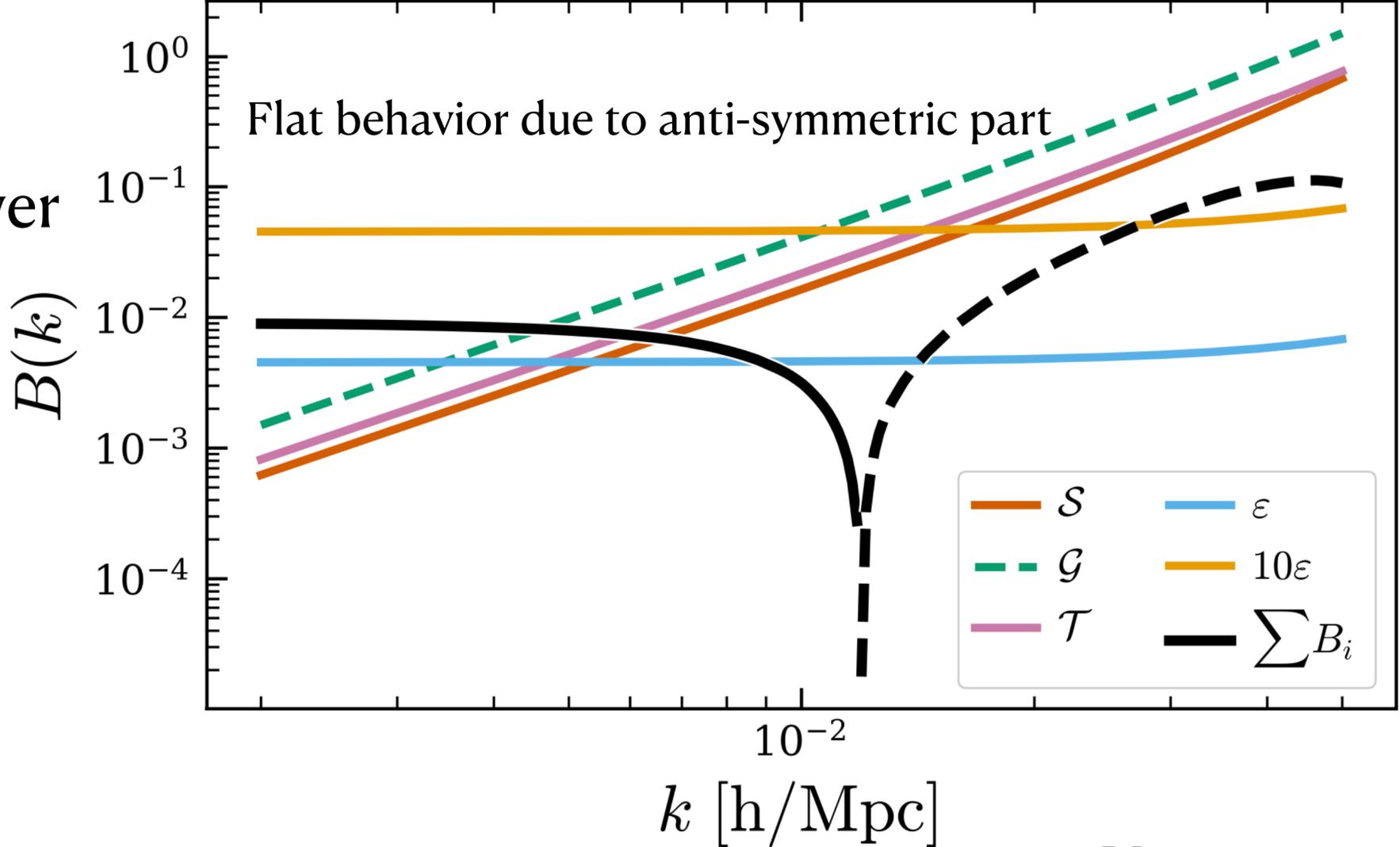
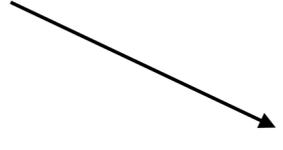
$$\langle \widehat{\delta}_L^{AB} \delta_{\text{matter}} \rangle \sim B^{ABm}(\vec{K}, \vec{k}_1, \vec{k}_2)$$

Main cost a bunch of FFTs. Widely applied for CMB secondaries.

GPU code with applications on EP, non-linearity and PNG

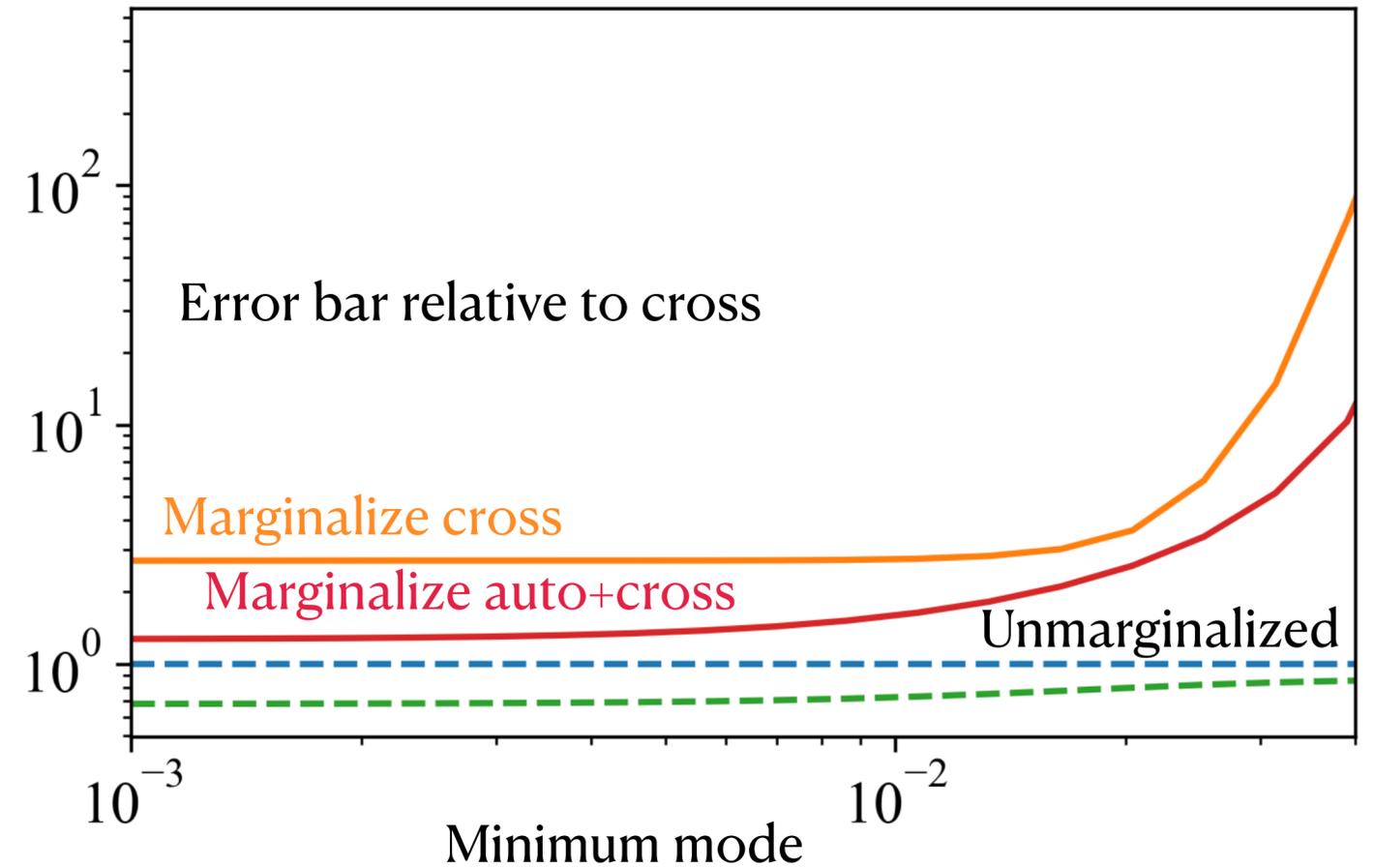
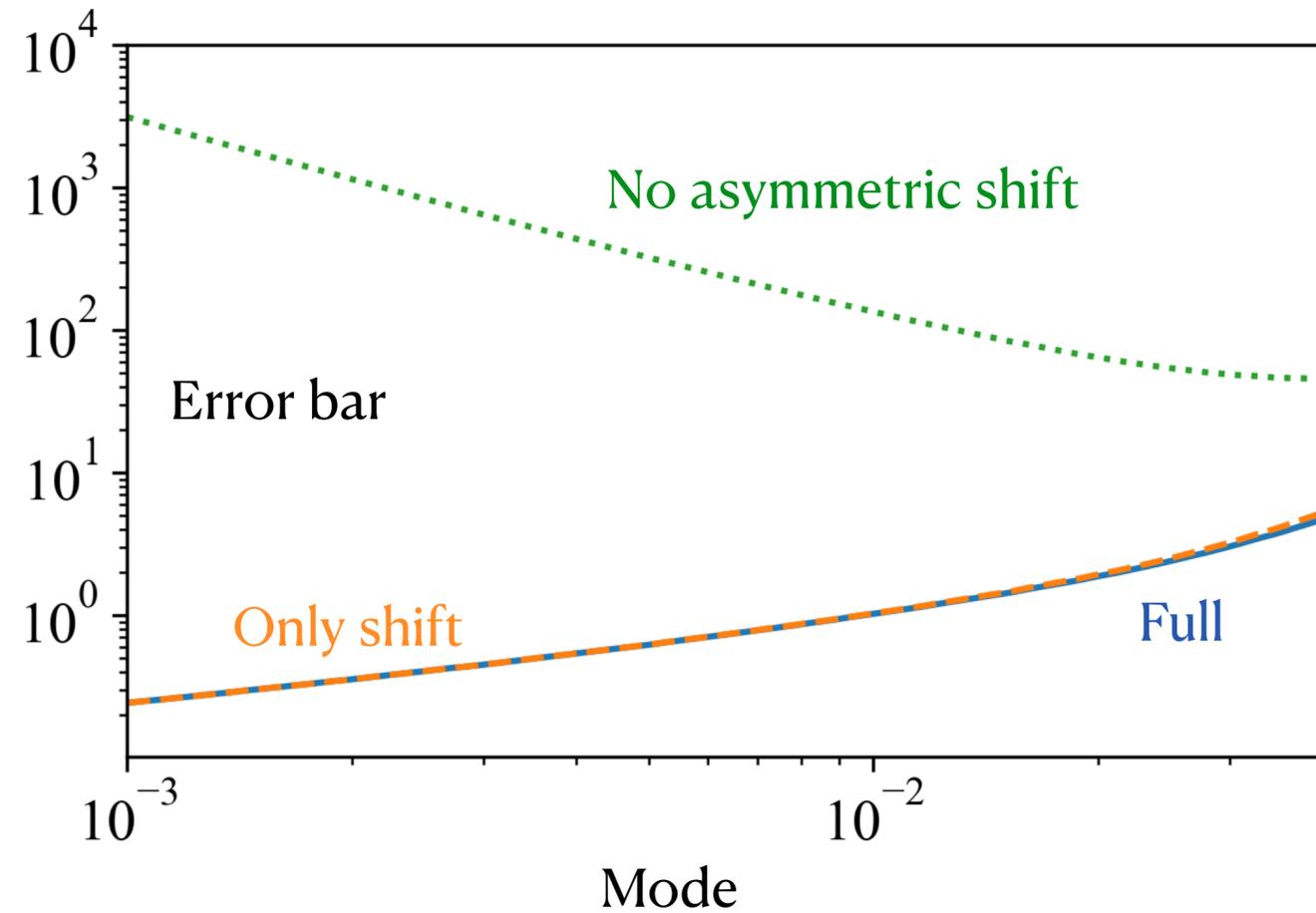
Effective reconstruction bias

Cross-Power/Linear Power



Very exaggerated example

Preliminary forecasts



Preliminary

Imagine a Stage V survey

Ideally.... $\lim_{x \rightarrow 0} \sigma^2(K)_{\text{joint}} \sim \sigma_{\text{single}}^2 k_{\text{max}}^{-3} \sim \sigma_{\text{single}}^2 / N_{\text{modes}}$

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