Neutrino Physics and Machine Learning 2024 HCI J4-June 25th ETH Zurich, CH





# Deep Learning in NOvA

**Barbara Yaeggy** On behalf of the NOvA Collaboration



The NuMI Off-Axis v<sub>e</sub> Appearance Experiment NOvA



#### Muon neutrino beam (NuMI) at Fermilab

- Two configurations: neutrino mode (FHC) and antineutrino mode (RHC)

### NOvA is an accelerator-based neutrino experiment

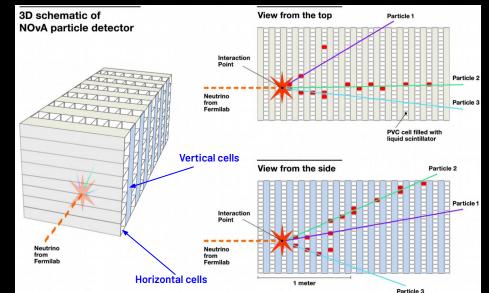
- Longest baseline in operation (810 km)
- Large matter effect, sensitive to mass ordering

# ~14 mrad off-axis, narrow-band beam around oscillation max

#### Barbara Yaeggy - University of Cincinnati

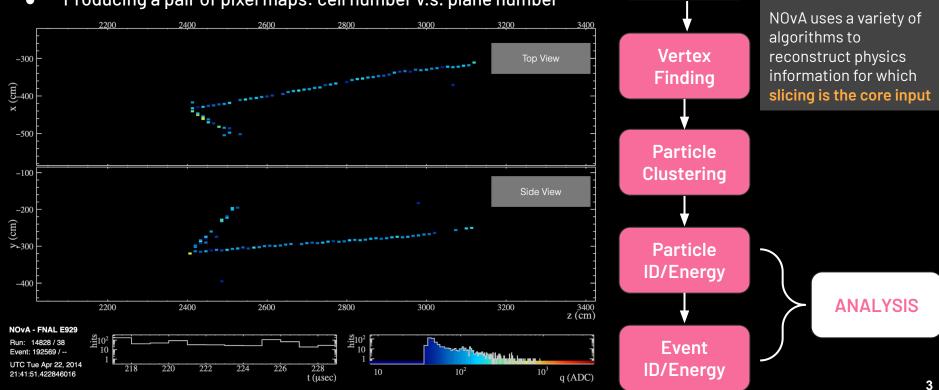
### **NOvA Detectors**

- ND (290-ton) and FD (14-kton) are functionally identical to minimize systematics
- Time resolution ~ ns, spatial resolution ~ cm, ND beam spill 10μm
- Composed of highly reflective extruded PVC cells filled with liquid scintillator. Scintillation light captured and routed to Avalanche Photodiode (APD) via wavelength shifting fiber (WLS)
- Cells arranged in planes, assembled in alternating horizontal and vertical directions → provide 3D views of the events



### **Event Reconstruction in NOvA**

- NOvA detectors are naturally segmented
- Producing a pair of pixel maps: cell number v.s. plane number



Slicer

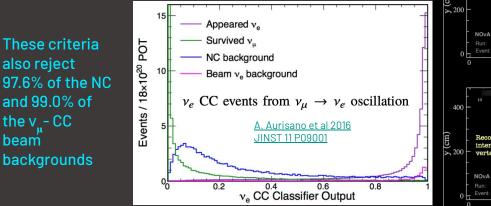
### NOvA use DL techniques for reconstruction and analysis NOvA is the first HEP experiment to apply CNNs to publish physics results: <u>Phys.Rev.Lett. 118 (2017)</u>

### Event Classifier - Oscillation related

Event classification (neutrino interactions or cosmic rays) and for particle classification starting from individual clusters

• **MobileNetv2** <u>arXiv:1801.04381</u> - Two-tower Siamese structure

Increase sensitivity to neutrino oscillation parameters over traditional methods (Phys. Rev. Lett. 116, 151806) equivalent to collecting 30% more exposure



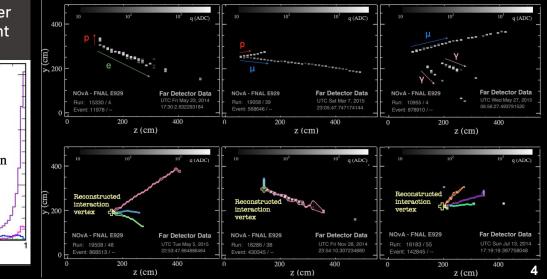
Barbara Yaeggy - University of Cincinnati

### Particle Classifier - Cross section measurements

Identify prongs as hadronic or electromagnetic (EM)

It takes the categories assigned by EventCVN and it labels by particle type, corresponding to the largest CVN score for the prong.

MobileNetv2 - Four-tower Siamese structure



### EventCVN Phys. Rev. Lett. 116, 151806

### MobileNetv2 arXiv:1801.04381 Two-tower Siamese structure

anti-neutrino mode

0.001

0.092

0.900

0.007

Nue

0.001

0.086

0.934

0.007

Nue

True Label

True Label

0.003

0.919

0.055

0.023

NC

0.003

0.858

0.057

0.024

NC

**NOvA Simulation** 

0.986

0.007

0.000

0.006

Cosmic

0.992

0.007

0.000

0.006

Cosmic

1.0

**NOvA Simulation** 

Color is efficiency

0.004

0.052

0.009

0.935

Numu

0.004

0.049

0.009

0.963

Numu

Color is purity

Cosmic

NC Label

Nue

Numu

Cosmic

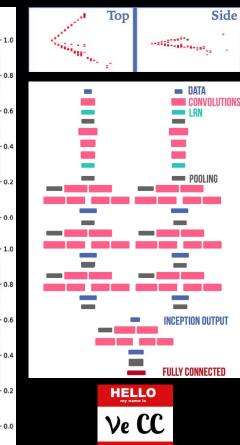
NC

Nue

Numu

Predicted Lab

dicted L

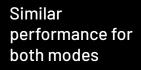


neutrino mode

**NOvA Simulation** 

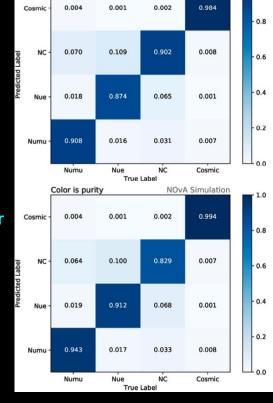
1.0

Color is efficiency



Anti-nu shows slightly increase in efficiency

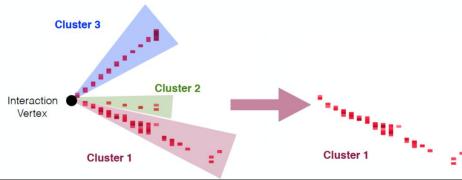
### Purity over 90% for all interactions flavors



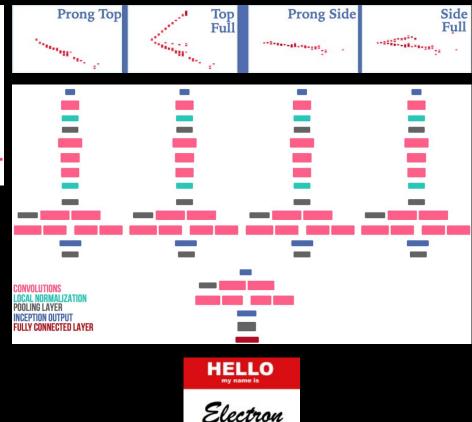


### ProngCVN Phys. Rev. Lett. 116, 151806

#### **MobileNetv2** <u>arXiv:1801.04381</u> Four-tower Siamese structure



- Single particles are separated using geometric reconstruction methods
- Hits with a common direction are grouped together to make clusters.
- Individual cluster identification
- Classifies particles using both views of the particle and both views of the entire event
- Currently, trained on separate samples of neutrino and anti-neutrino mode



### **TransformerCVN** *Event & Particle Classification*

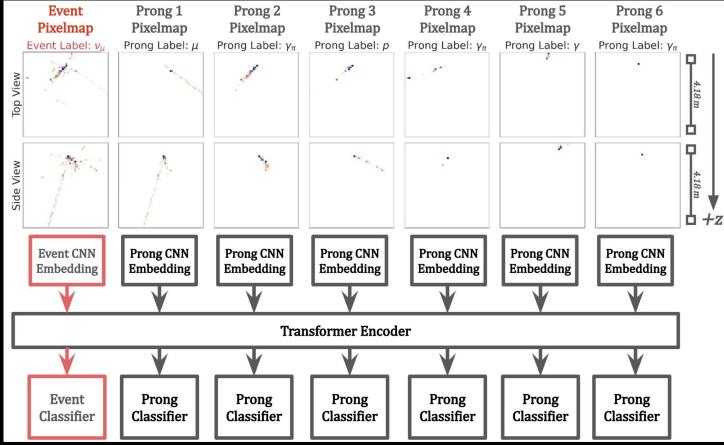
Spatial learning + attention mechanism: simultaneously classifies each event and reconstructs every individual particle's identity

Classifies events with 90% accuracy & improves the reconstruction of individual particles by 6% over baseline methods

## Great improvement in particle identification!

Thanks to the the additional context provided by all prongs and the transformer's attention mechanism

### Coming talk with details about Transformer CVN at NOvA, by Jianming Bian



#### arXiv:2303.06201

### Regression CNNs for Energy Estimation Phys. Rev. D 99, 012011

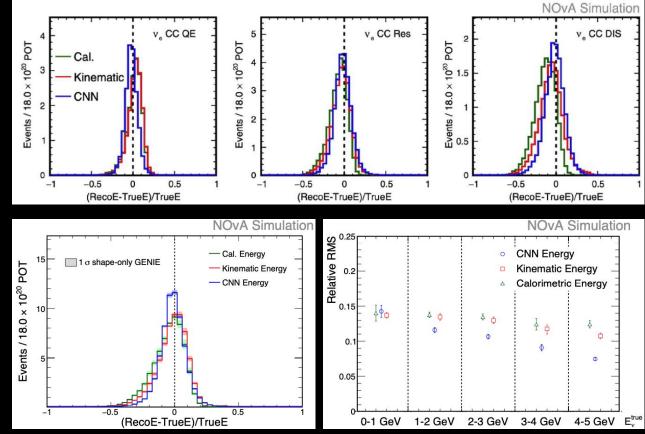
## The CNN architecture used is an adapted ResNet

- Weighting scheme, so the loss function sees a flat energy distribution, to control energy dependence
- Use mean absolute percentage error instead of square of errors to decrease the effects of outliers

Shows a better resolution compared with kinematics-based energy reconstruction

Provides improvements of 16% and 12% in RMS for v<sub>e</sub> CC and electron, respectively

Good stability over interaction types!



## LSTM for Energy Estimation

- Long Short-Term Memory (LSTM) is a type of recurrent neural network
- Takes a number of traditional reconstruction quantities as inputs (prong & slice level) → energy prediction
- Trained with artificially engineered sample to increase network resilience
- Resolution comparable with regression CNN, (used for quick cross-check)

### ML Vertexer, VertexCVN

- More accurate vertex finding, means more accurate on:
- Clustering hits to form individual particle tracks/showers
- Identifying particle types
- Energy estimation
- Shows good performance across interaction types!
- Developed to address several known failure modes of NOvA's existing algorithm "Elastic Arms"
- Forward failure tendency for main prong to be split into two
- Backward failure tendency for multiple, small prongs to be combined into one

#### Neutrino Beam **NOvA Simulation** 10000 Standard 8000 Mean -2.1% Mean -1.2% **RMS 10.8% RMS 9.4%** Events 6000 4000 2000 -0.25 0.00 0.25 -1.00-0.75-0.500.50 0.75 (Reco - True) / True Neutrino Energy **True Vertex** VertexCVN **Elastic Arms** z (cm)

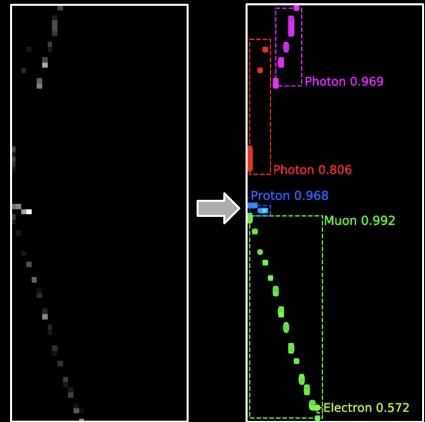
NOvA Simulation

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### We want to do a better reconstruction!

Semantic segmentation : pixel level ID Instant segmentation: semantic segmentation + object detection

- Full event reconstruction on a hit-by-hit basis using instance segmentation:
- Bounds: Create a bounding box around each particle with a Region-based CNN (RCNN)
- ID Score: Use a softmax function to classify the particle contained within each box
- Clusters: Group together hits, identify hits, then individual hits are combined to form clusters
  - Very powerful in PID and clustering efficiency
  - No dependence on other reconstruction (vertex, etc)
  - However, it's quite slow to run on CPUs, and more work needs to be done to run at scale



## **Summary**

• NOvA pioneered the use of CNNs for event classification in HEP and implemented improved networks for recent analyses

- In NOvA, machine learning has been developed to:
- Identify events and final state particles
- Singularly simulated particle for ND analyses (no contextual information)
- Cosmic filtering, based on ResNet18 with a siamese structure
- Reconstruct neutrino energy, final state particle energy, vertex
- Perform full event reconstruction

• **Other ongoing ML efforts in NOvA:** improve ProngCVN with both neutrino and antineutrino sample, graph neural networks, unsupervised training.

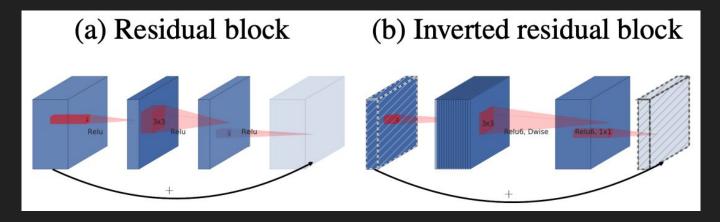
• NOvA has been performing expansive data comparison, impact analysis, uncertainty studies and cross-checks to improve robustness and interpretability of ML techniques



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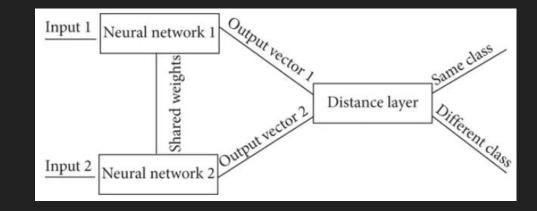
### MobileNetV2



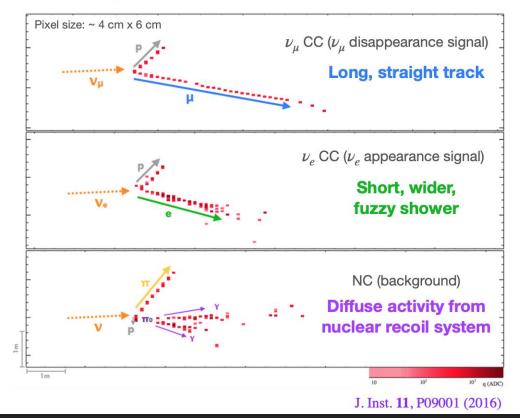
- CNN
- Thickness of each block to indicate its relative number of channels
- The MobileNet v2 architecture is based on an inverted residual structure where the input and output of the residual block are thin bottleneck layers opposite to traditional residual models which use expanded representations in the input. MobileNet v2 uses lightweight depthwise convolutions to filter features in the intermediate expansion layer.

## Siamese Neural Networks

- Innovative technique known as similarity problems which finds if two inputs are similar or not which is known as a siamese neural network
- We require only one training example for each class
- This kind of neural network architecture is scalable and does not require much retraining
- Siamese networks are not learning to classify an image to any of the output classes.
- But, it is learning with help of a similarity function, which takes two images as input and gives the probability of how similar these images are.



# **Event Topology**



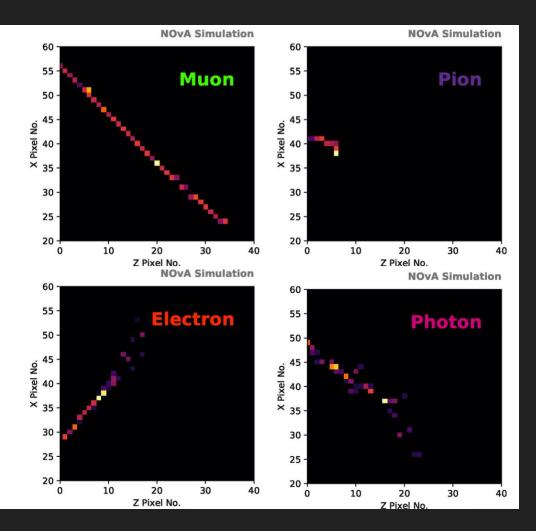
# **Cosmic filtering with a NN**

- Network based on ResNet18 backbone with a siamese structure
  - Takes in two event images (top-view and sideview) as input
- Softmax output with five labels:  $v_{\mu},\,v_{e},\,v_{\tau},\,NC,$  and cosmic score
- Training sample contained 1M+  $v_{\mu},\,v_{e},\,and$  NC events in both beam modes and 5M+ cosmic events
  - Not trained separately for neutrino/ antineutrino mode
- Performs better than traditional cosmic rejection in all samples

Data Sample	Traditional Cosmic Rejection	Cosmic Rejection Neural Network
$\nu_e$	93.21	99.71
$\overline{\nu_e}$	92.81	99.82
$ u_{\mu}$	93.22	99.20
$\overline{\nu_{\mu}}$	92.82	99.20
ν ΝΟ	93.24	97.08
ν̄ NC	92.79	96.82
Cosmic v	7.80	5.00

# Single particle ID

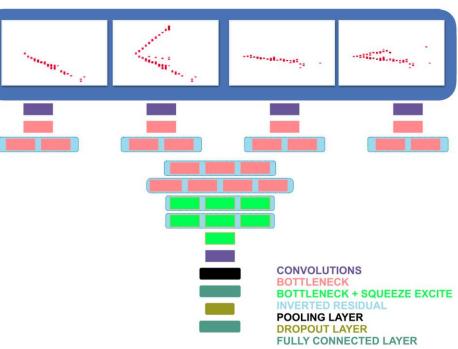
- NOvA also has trained a network using singularly simulated particles for ND analyses → no contextual information
- Also developing a network designed for neutron identification using these samples



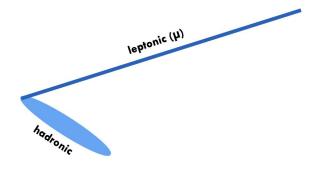
# Improved ProngCVN

#### Akshay Chatla, DAE 2022

- Modifies ProngCVN (modified MobileNetv2) architecture by adding Squeeze-Excite block for channel attention
- Trained on a combined sample of neutrino and antineutrino mode
- Shows good performance for particle classification



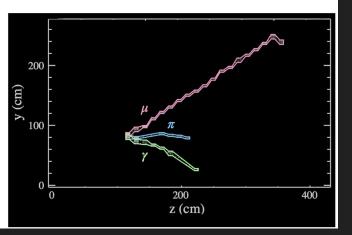
## **Reconstruction Uses**

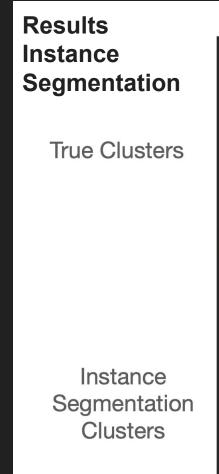


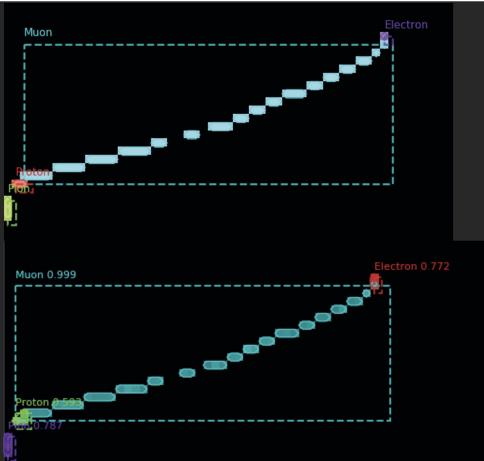
Particle clusters are the base unit for neutrino energy estimation.

 $E_v = E_{lep} + E_{had}$ 

Clusters are used to identify specific final states for cross section measurements.







### Performance of LSTM EE

To assess the performance of both standard EE and LSTM EE, we define the energy resolution as

$$R = rac{E_{reco} - E_{true}}{E_{true}}.$$

A smaller Root Mean Square(RMS) of resolution indicates a better performance of the estimator.

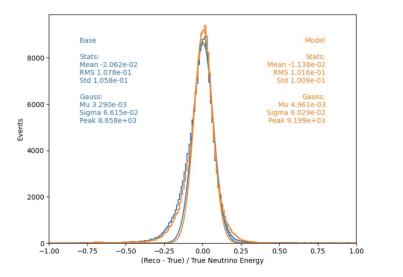


Figure 1: Plot of muon neutrino energy resolutions of standard EE and LSTM EE.

The default neutrino sample we used is an FD FHC sample generated by Genie and GEANT4

LSTM EE has better performance than standard EE in NOvA.

#### It has

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smaller resolution of energy and higher stability with respect to systematics in NOvA.

It help NOvA to determine the neutrino oscillation parameters.